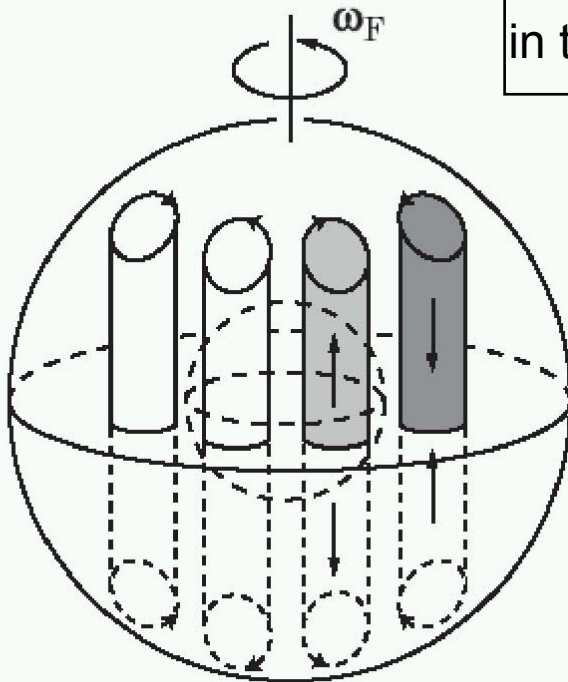


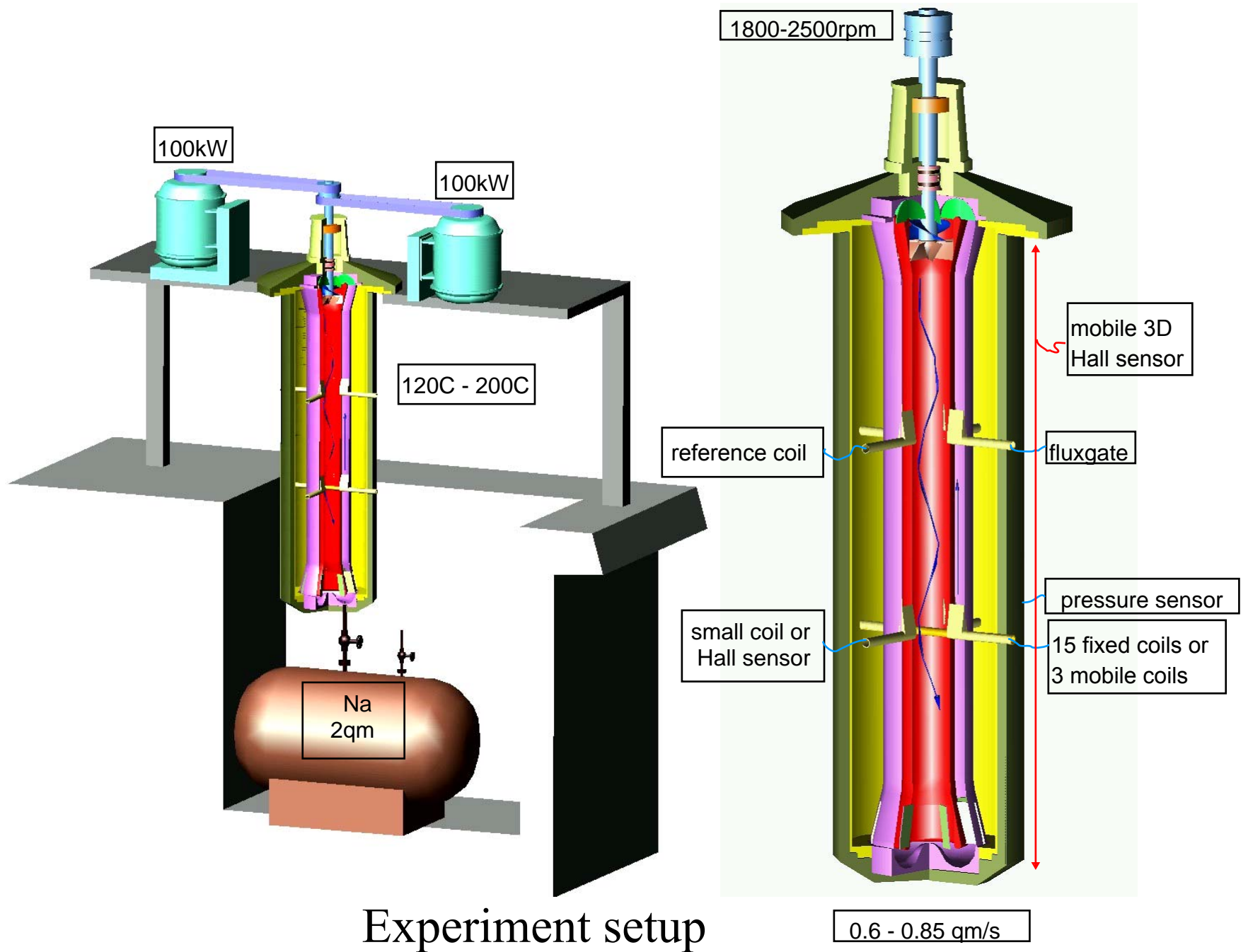
RIGA DYNAMO EXPERIMENT

A.Gailitis, O.Lielausis, E.Platacis,
G.Gerbeth & F.Stefani

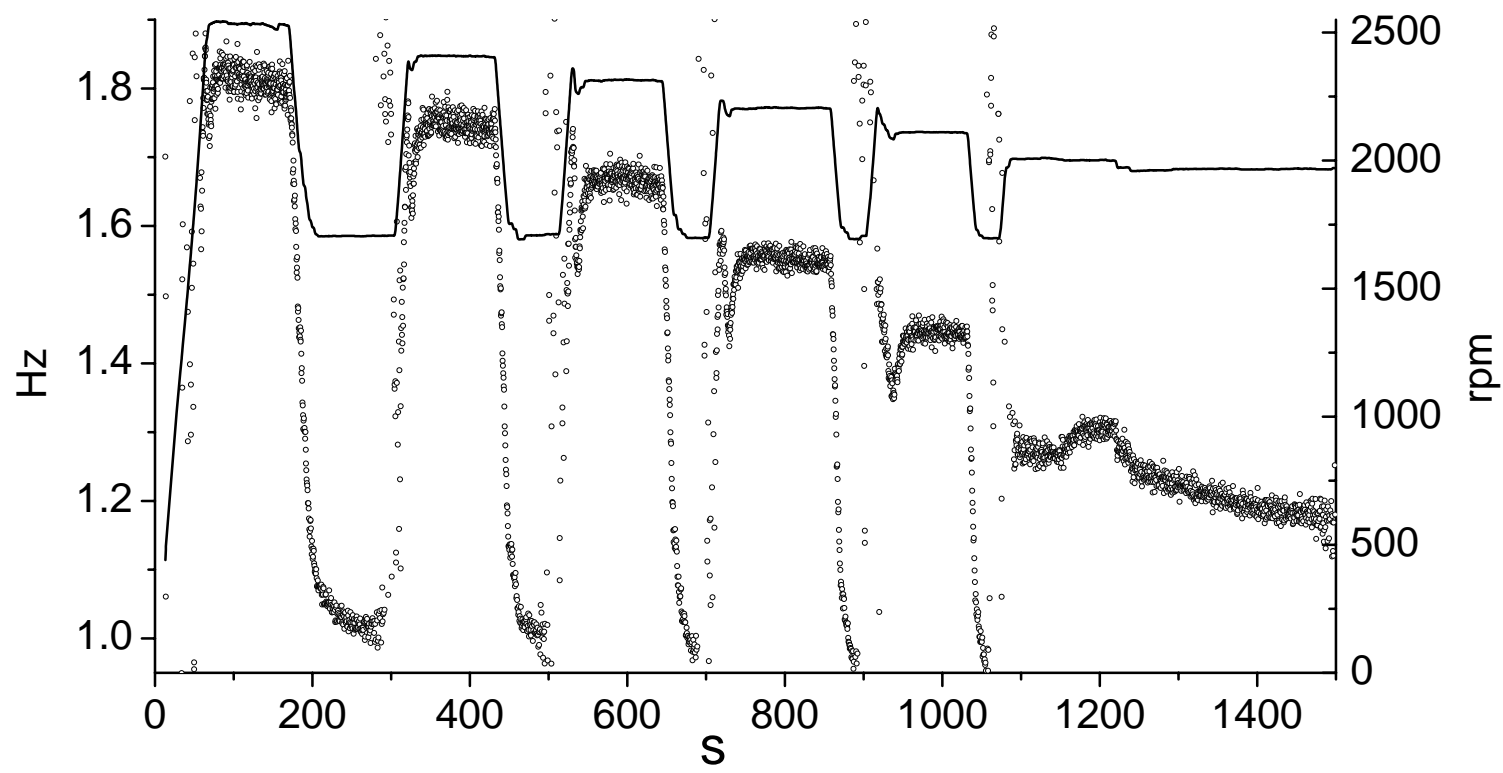
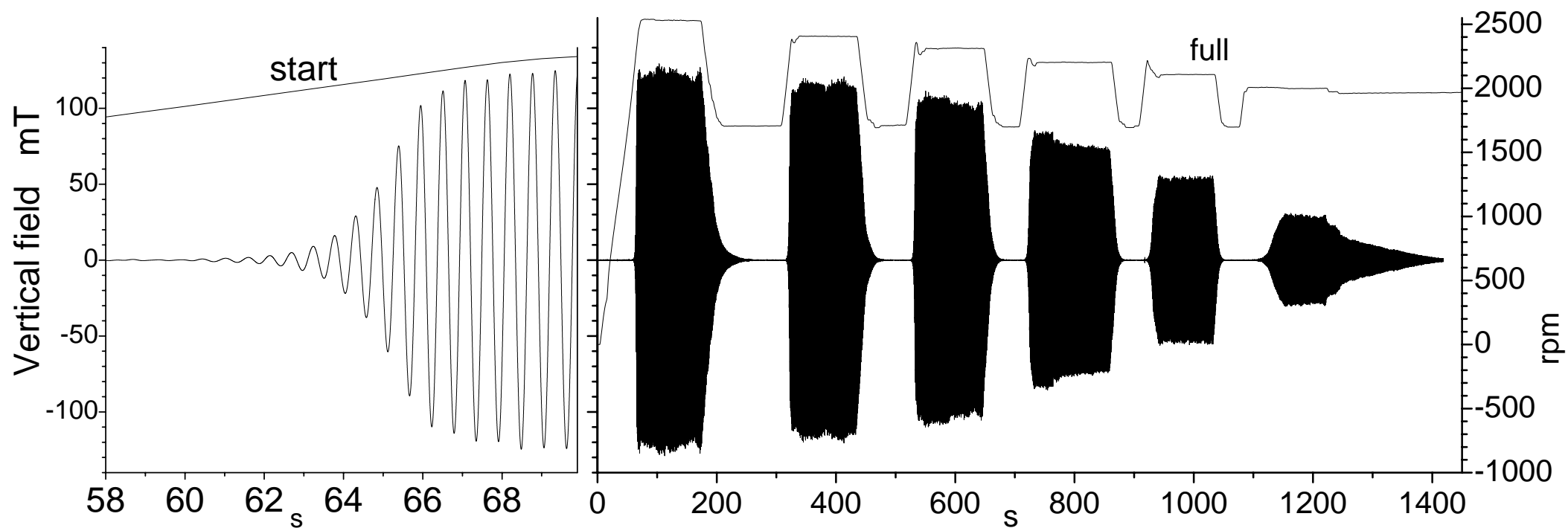
Convection columns in the Earth core

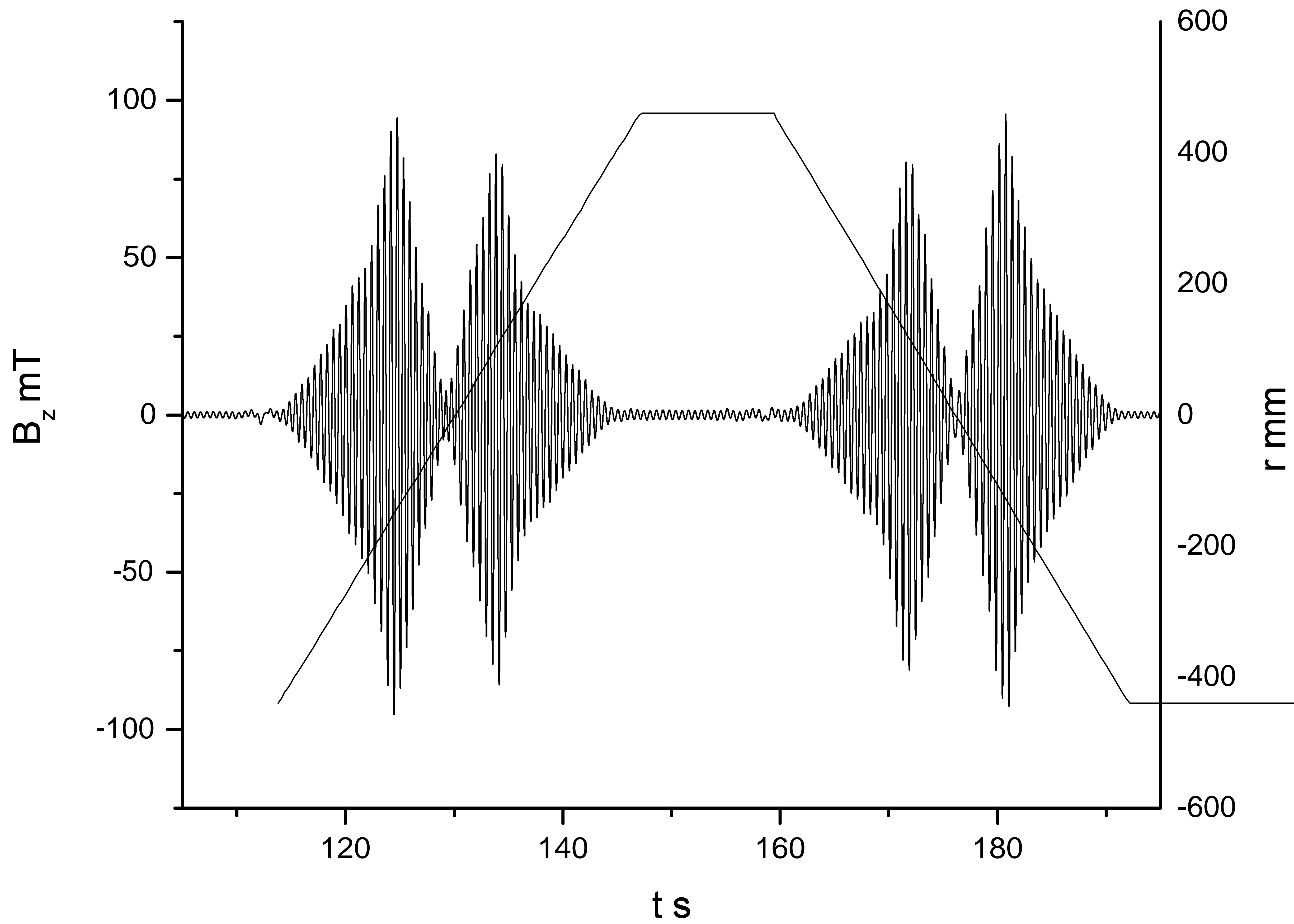




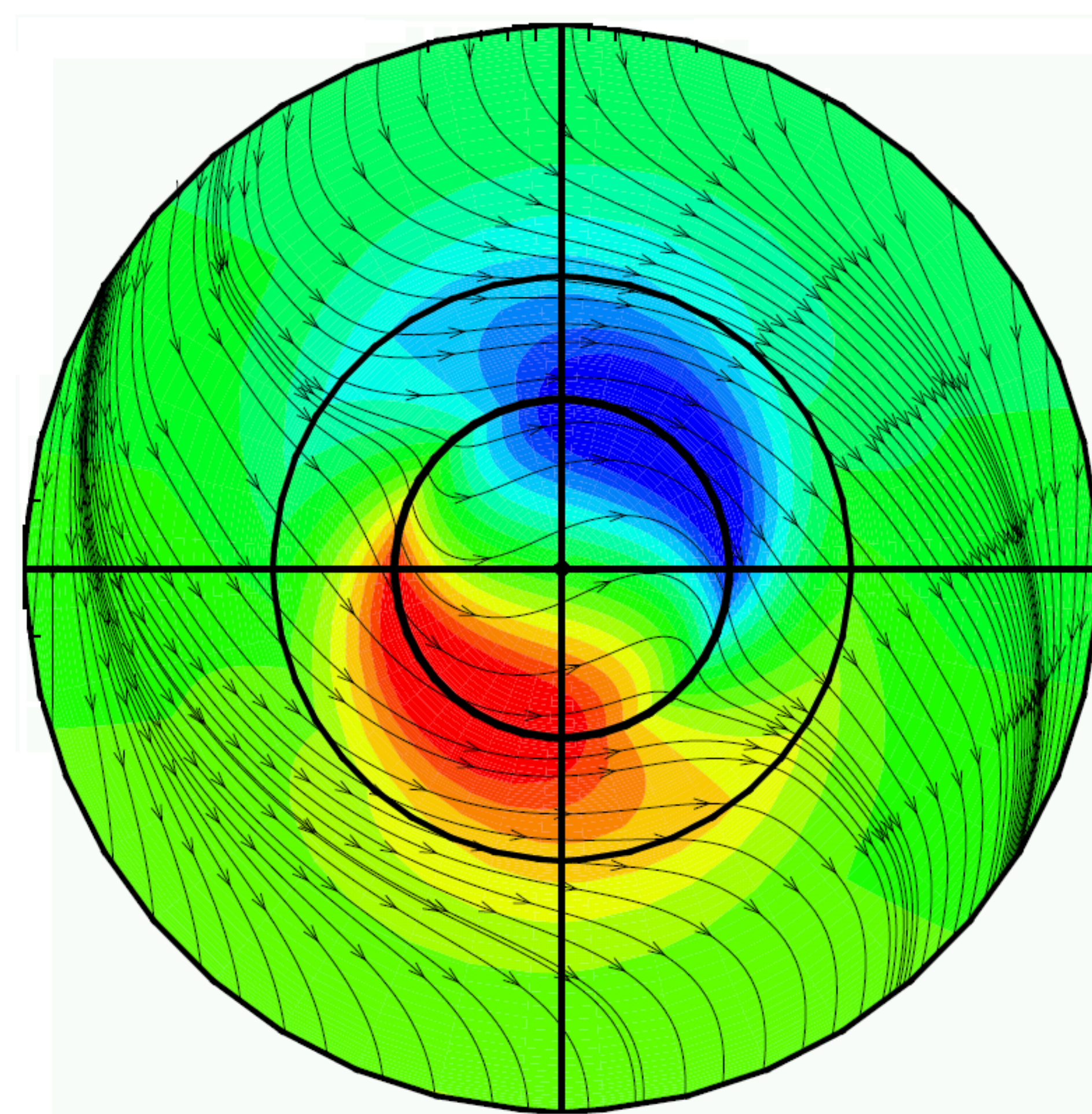
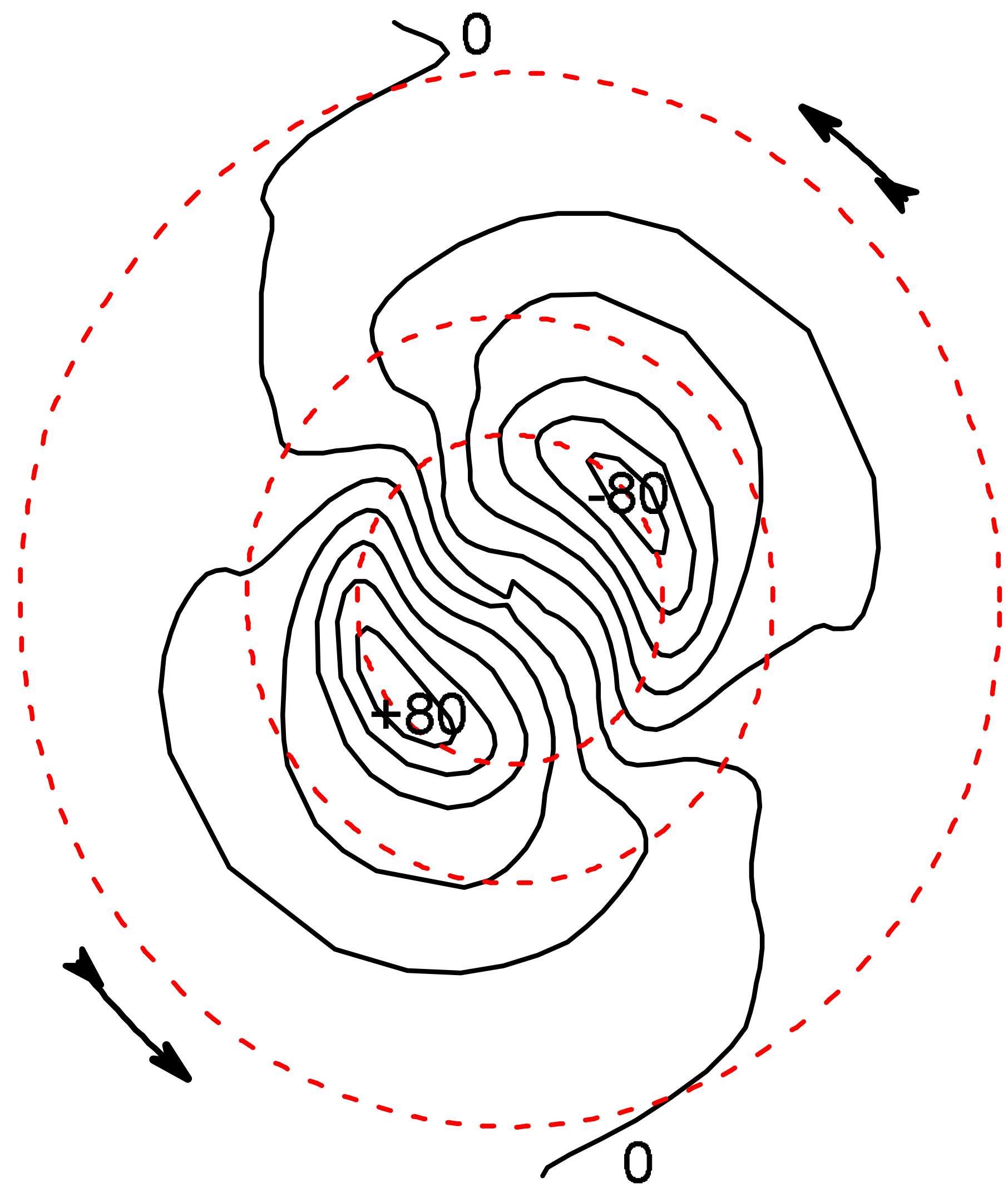
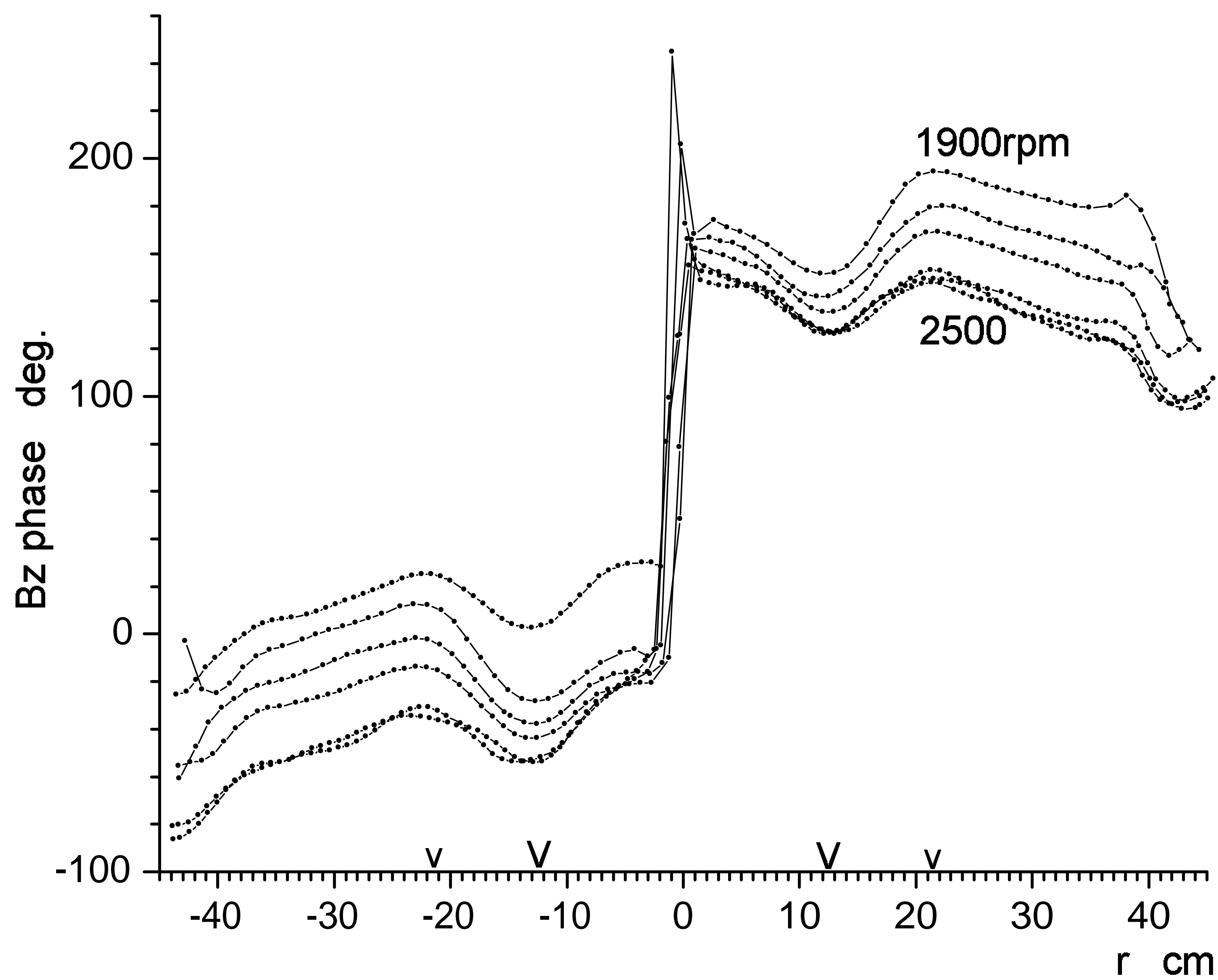
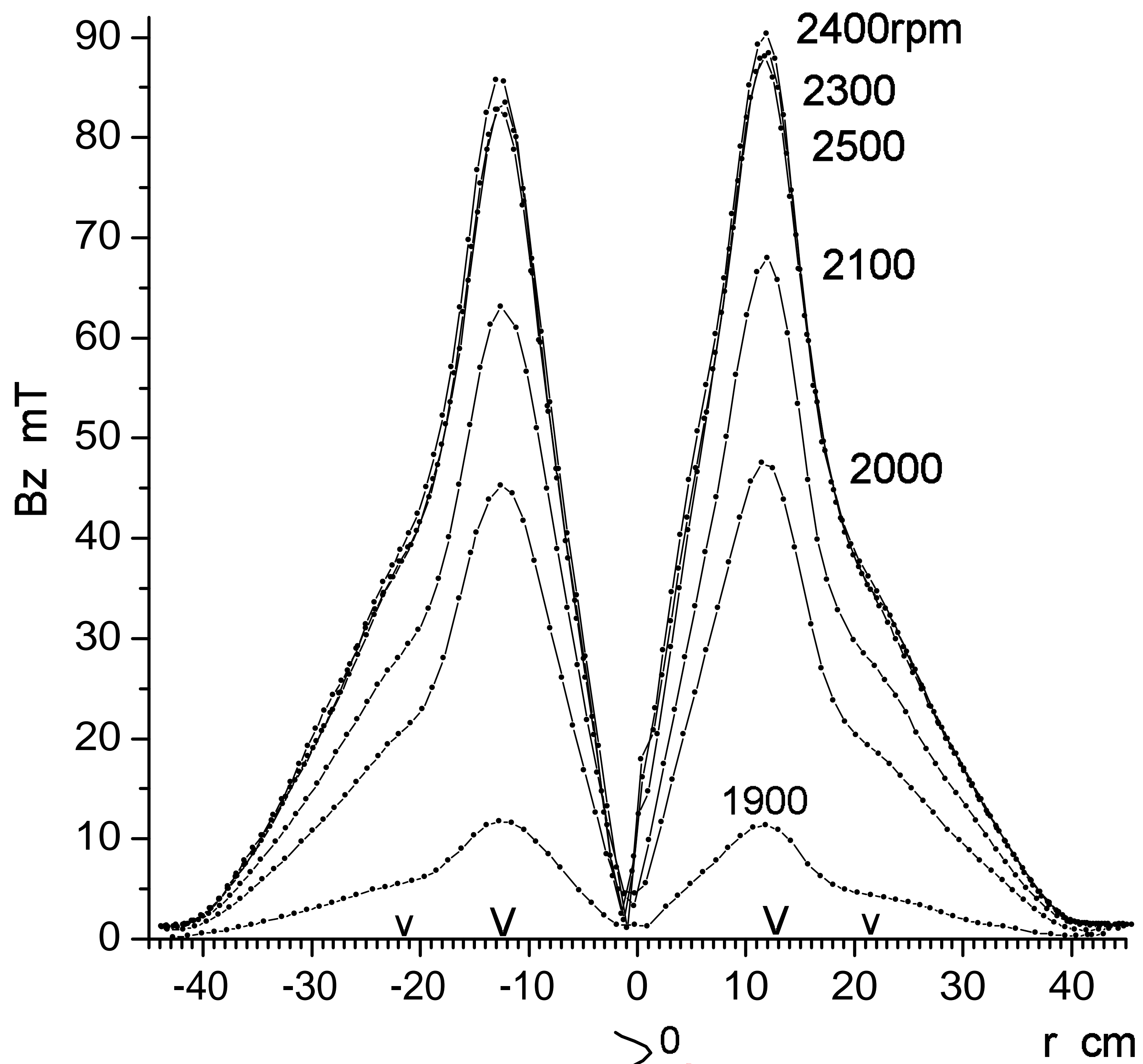


Experiment setup

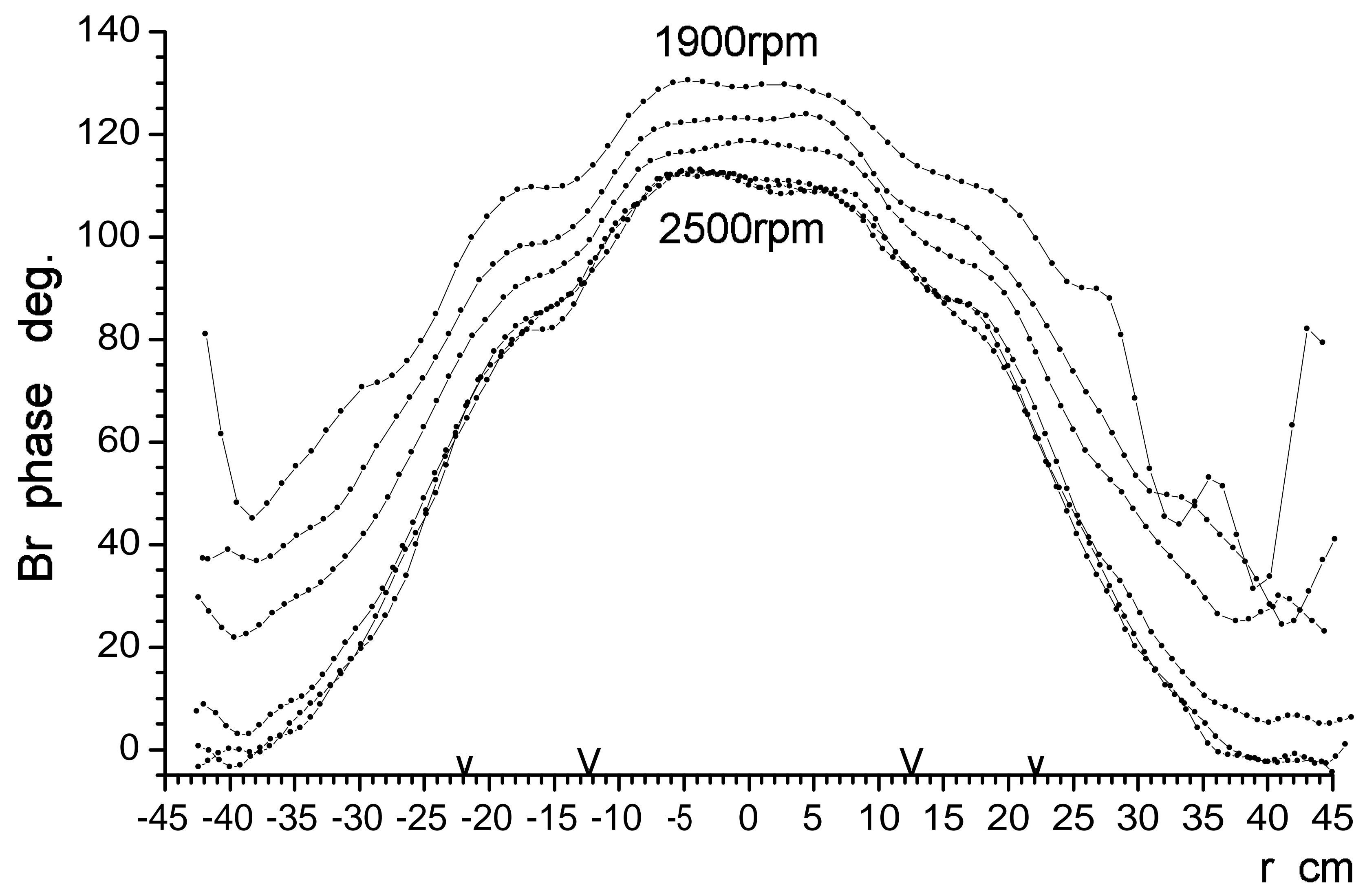
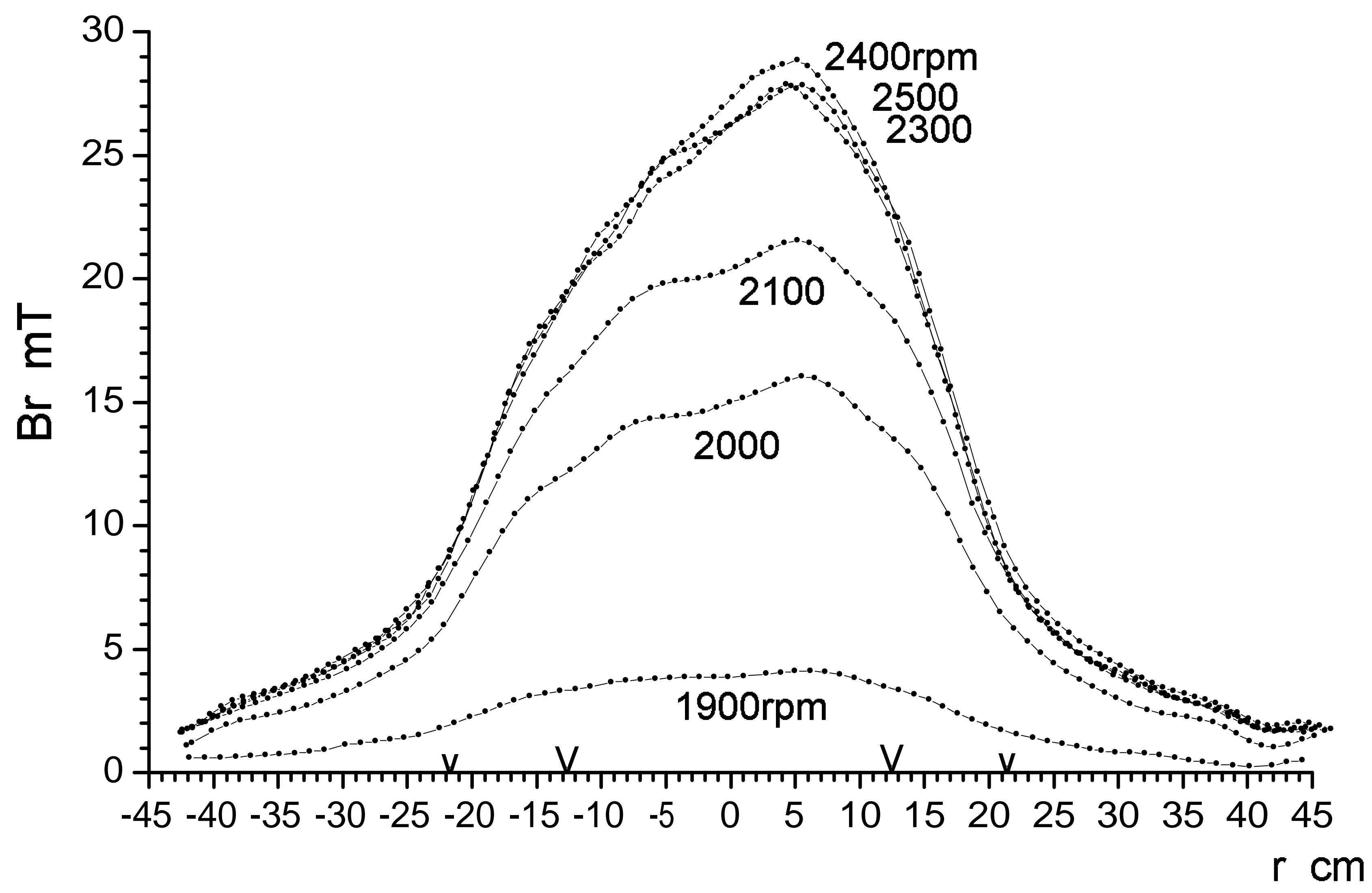
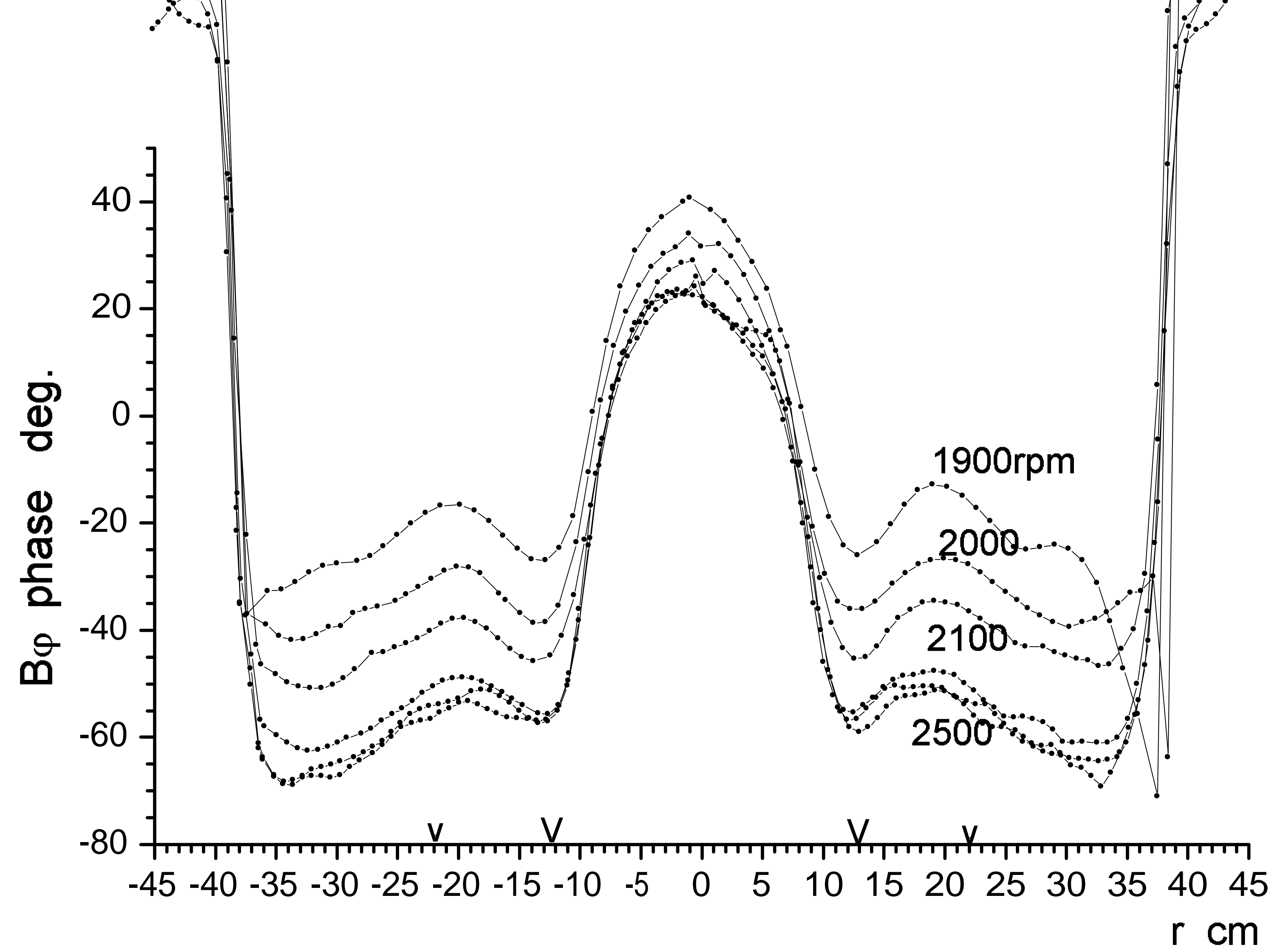
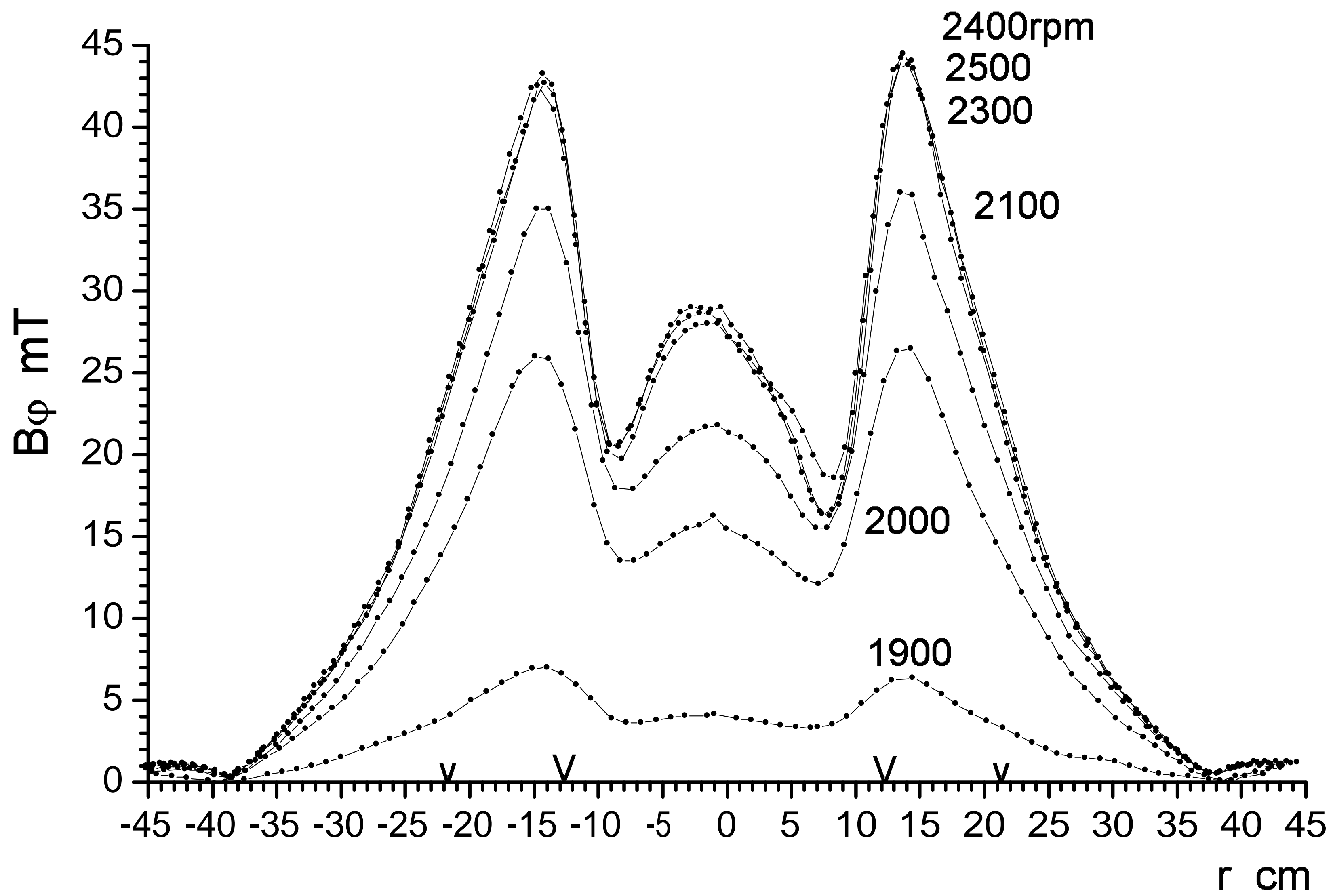


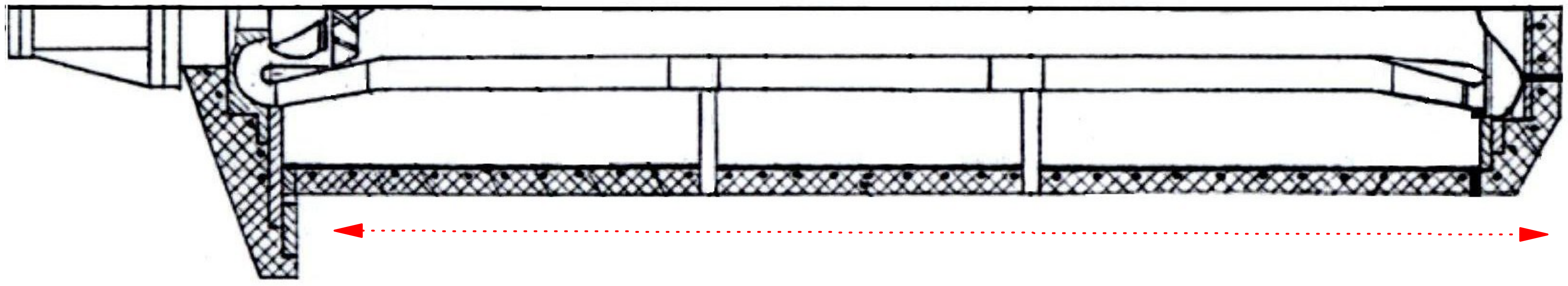
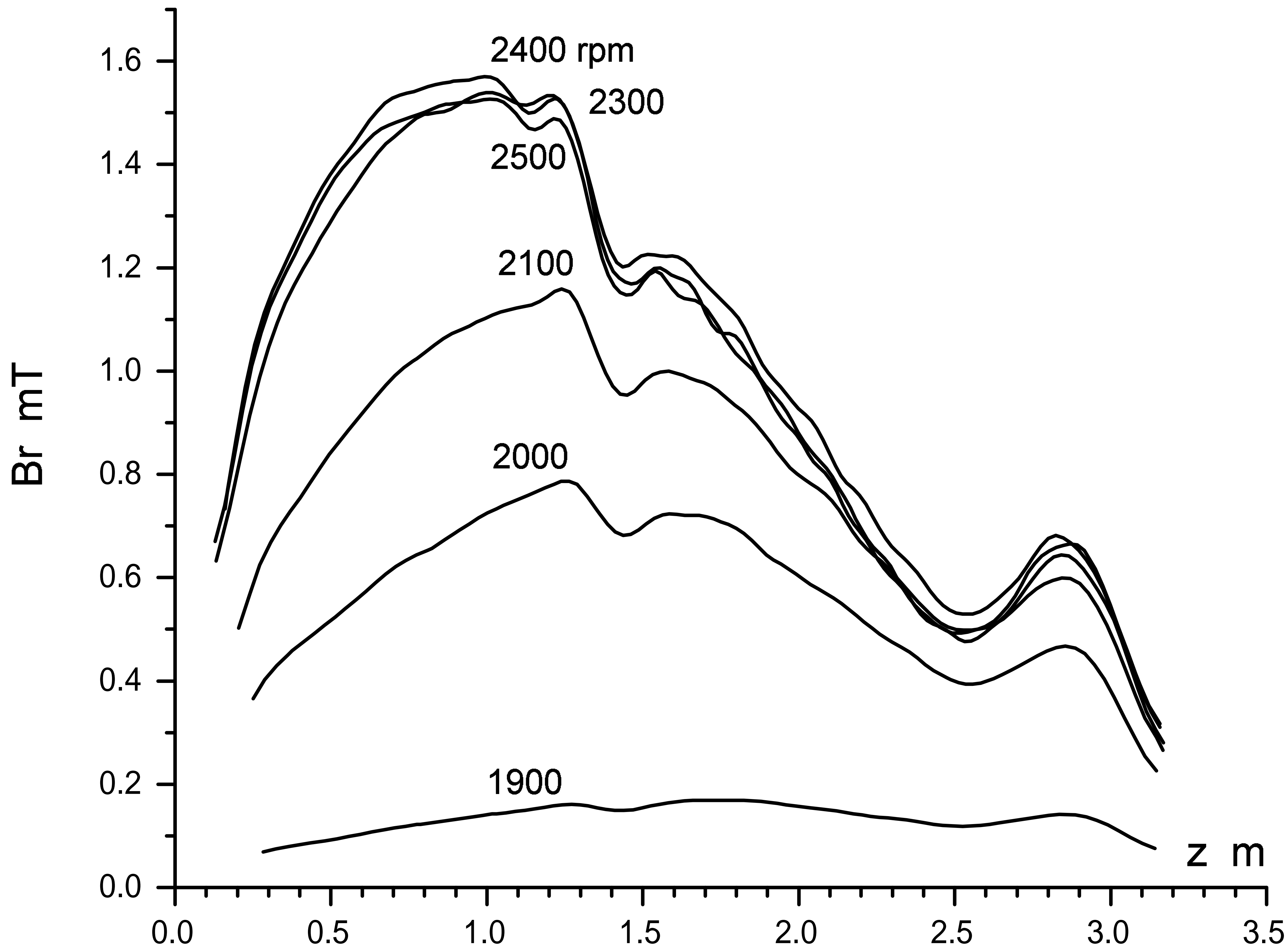


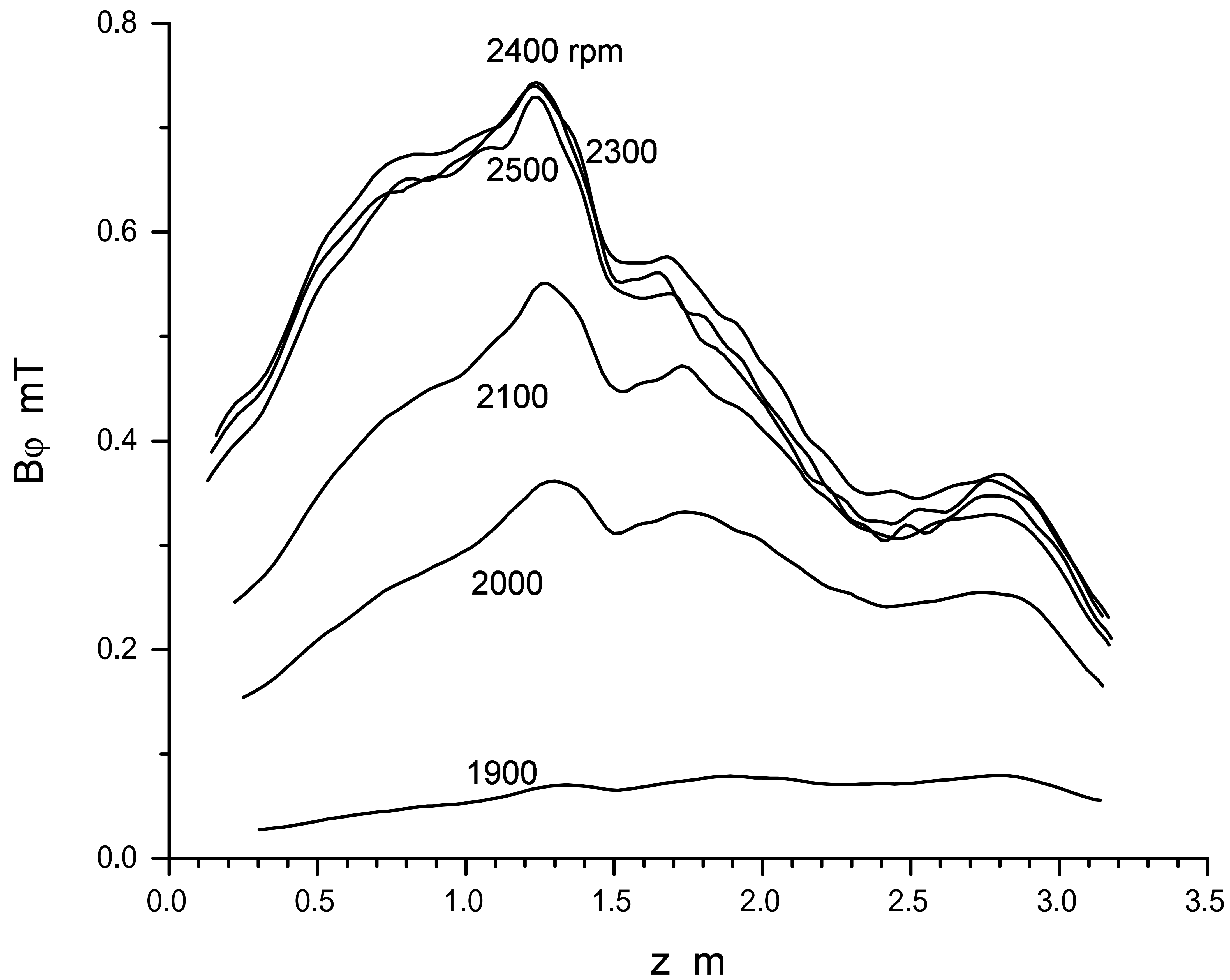
VERTICAL FIELD

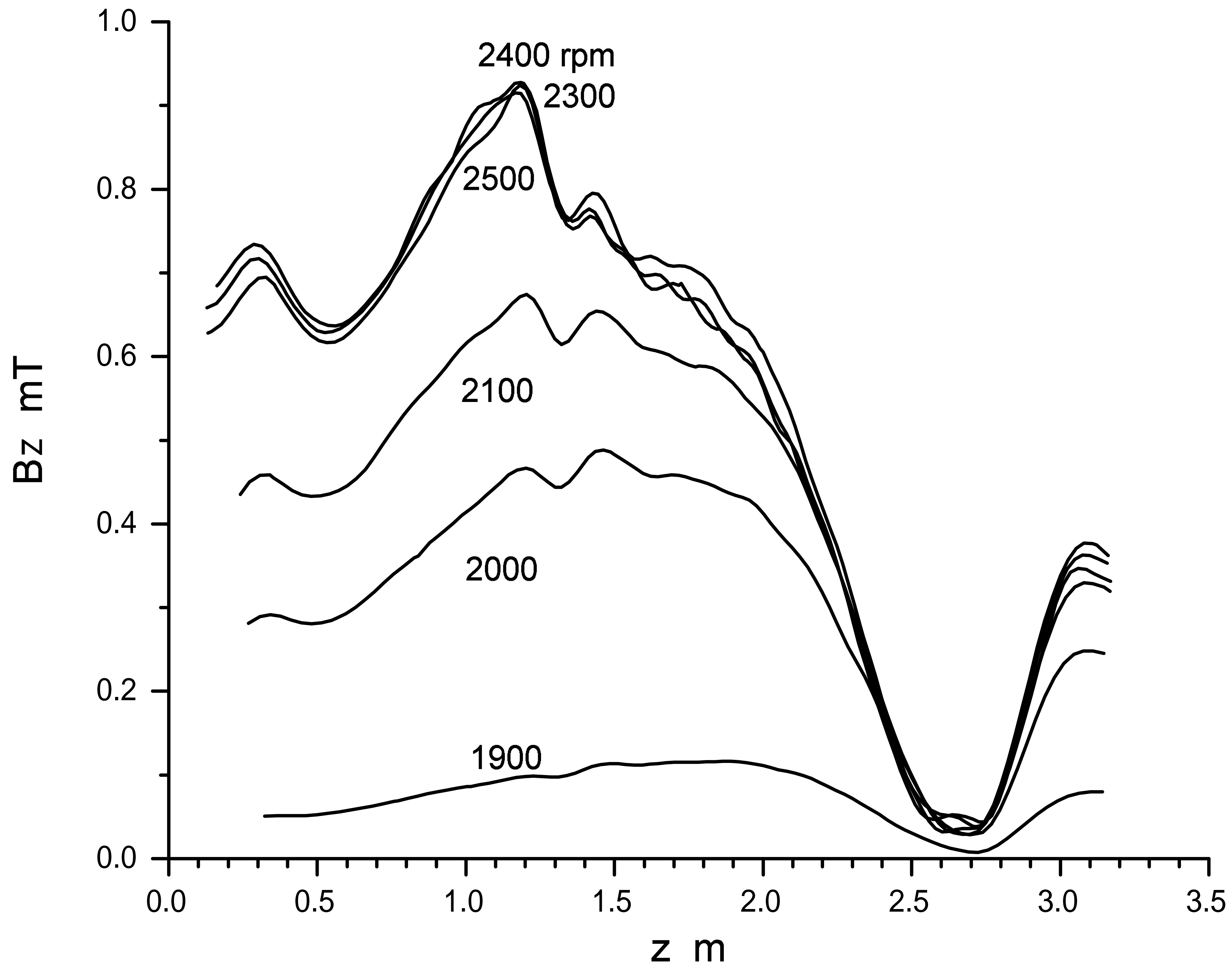


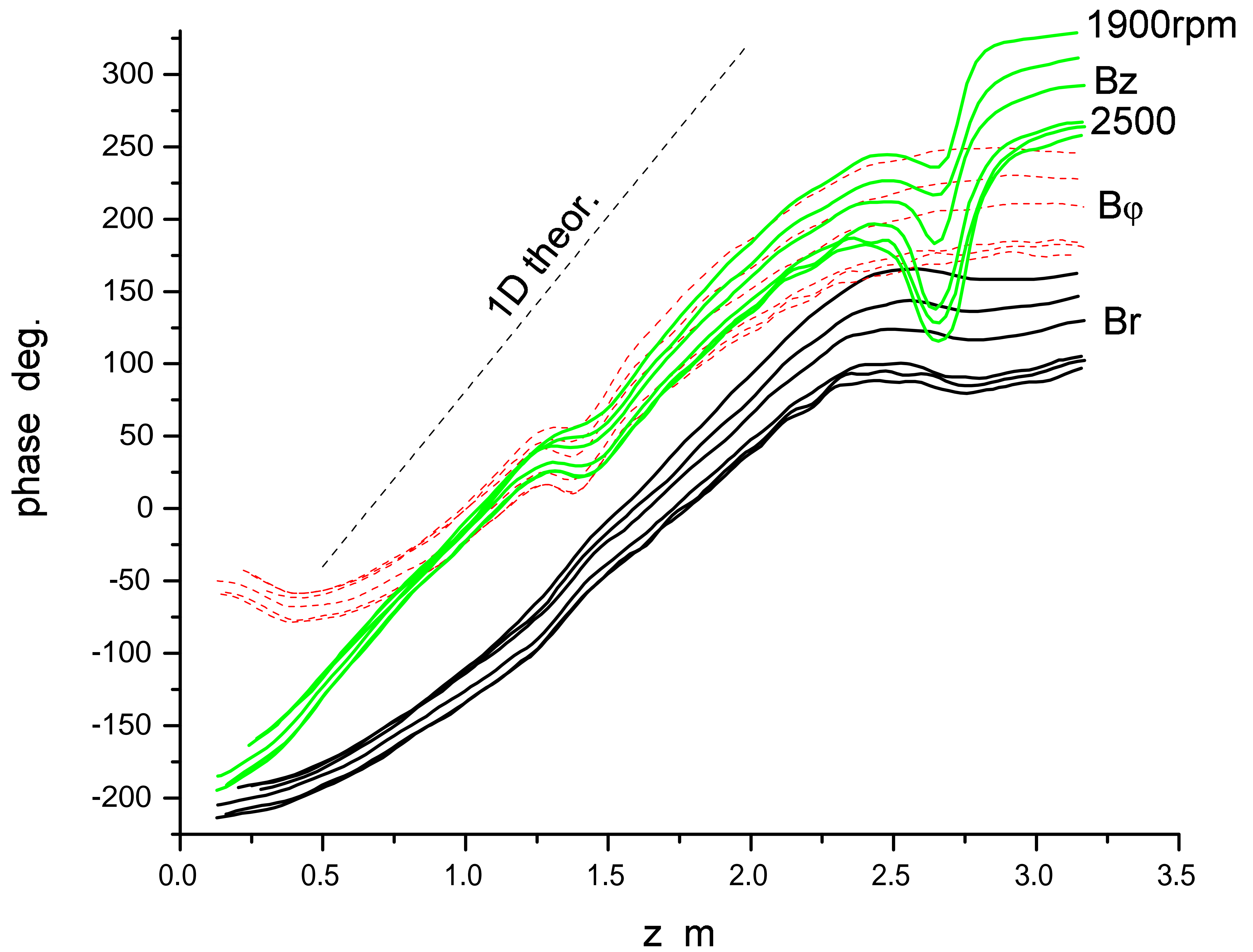
Horizontal field

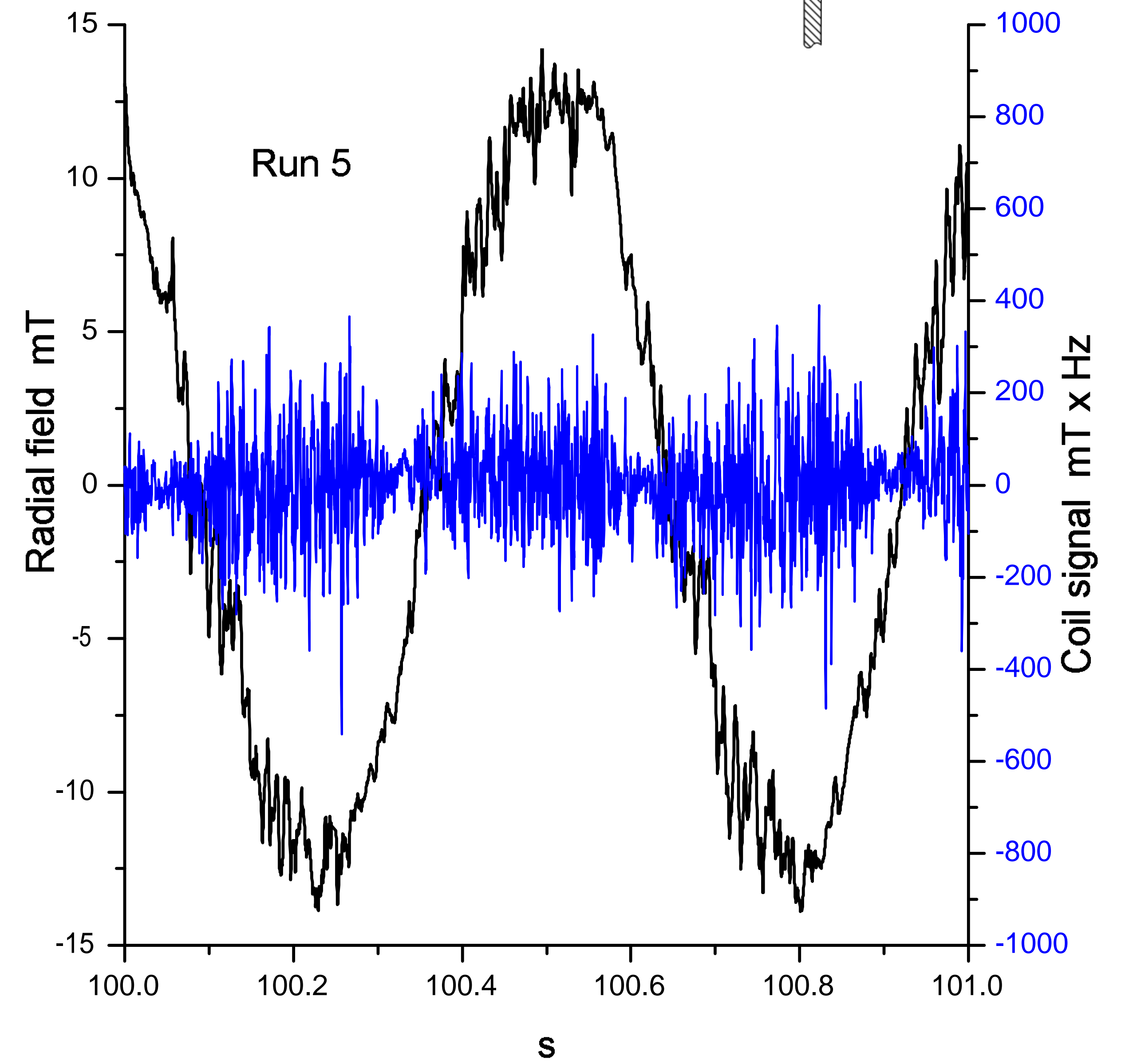
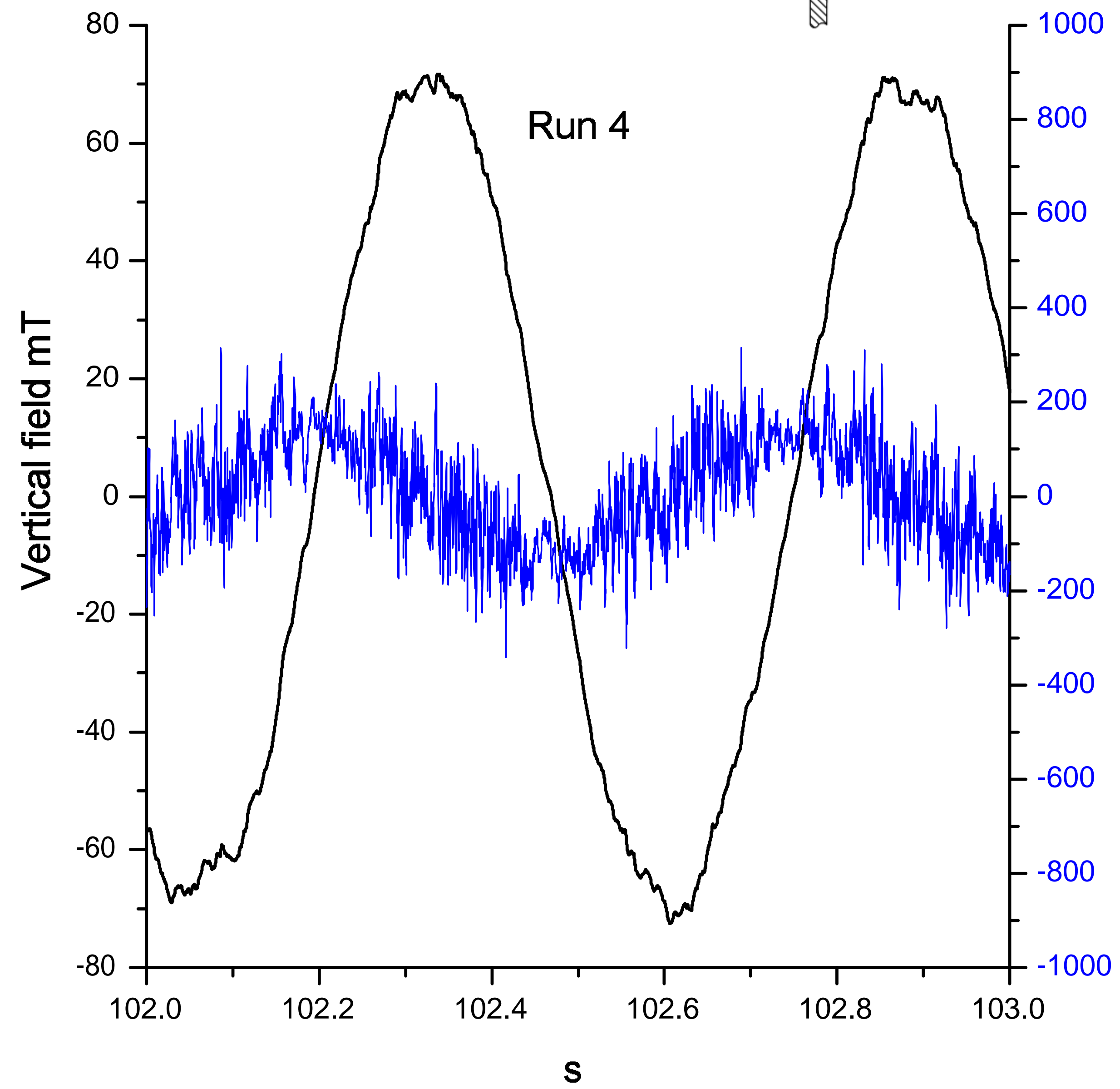
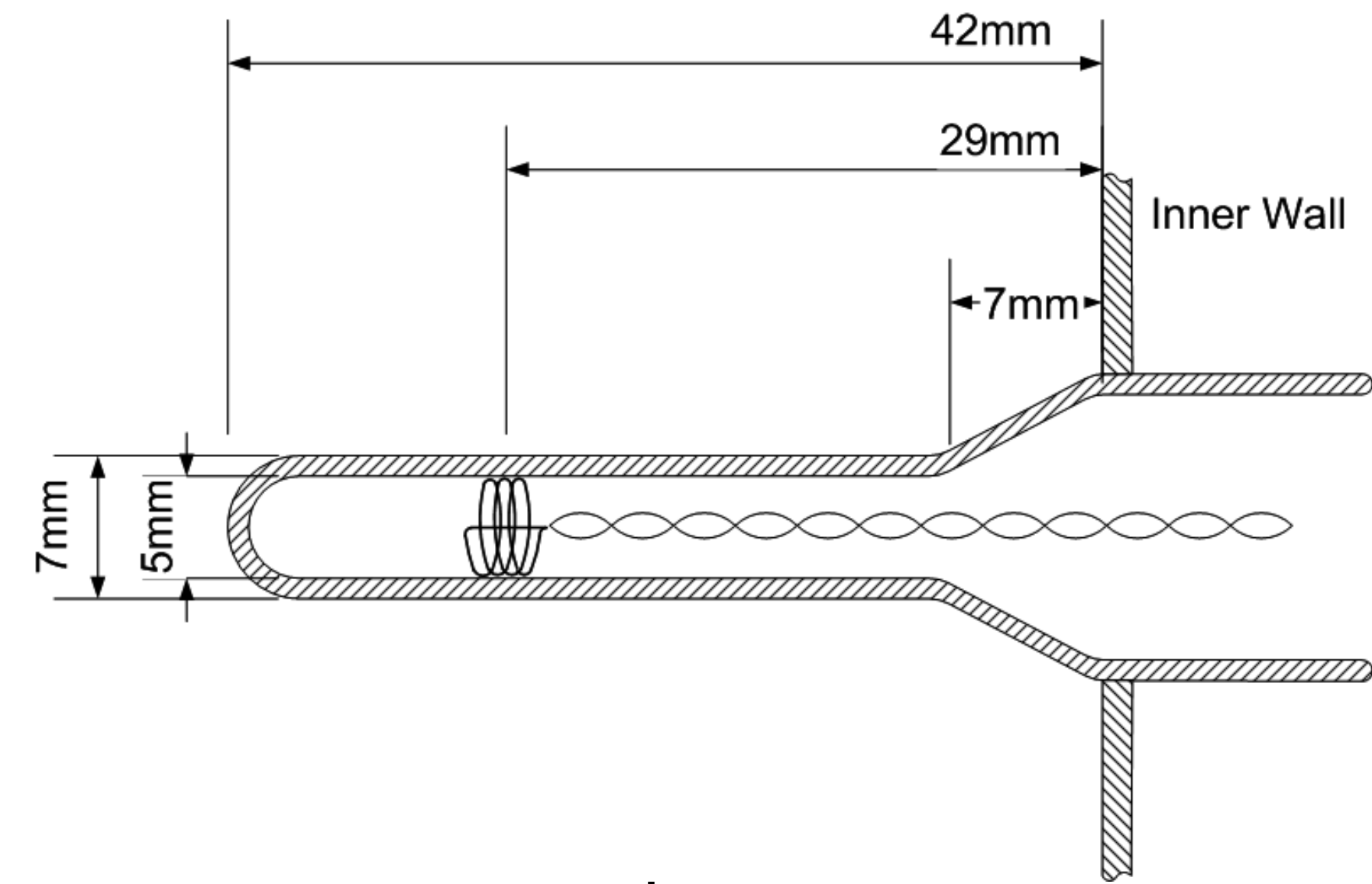
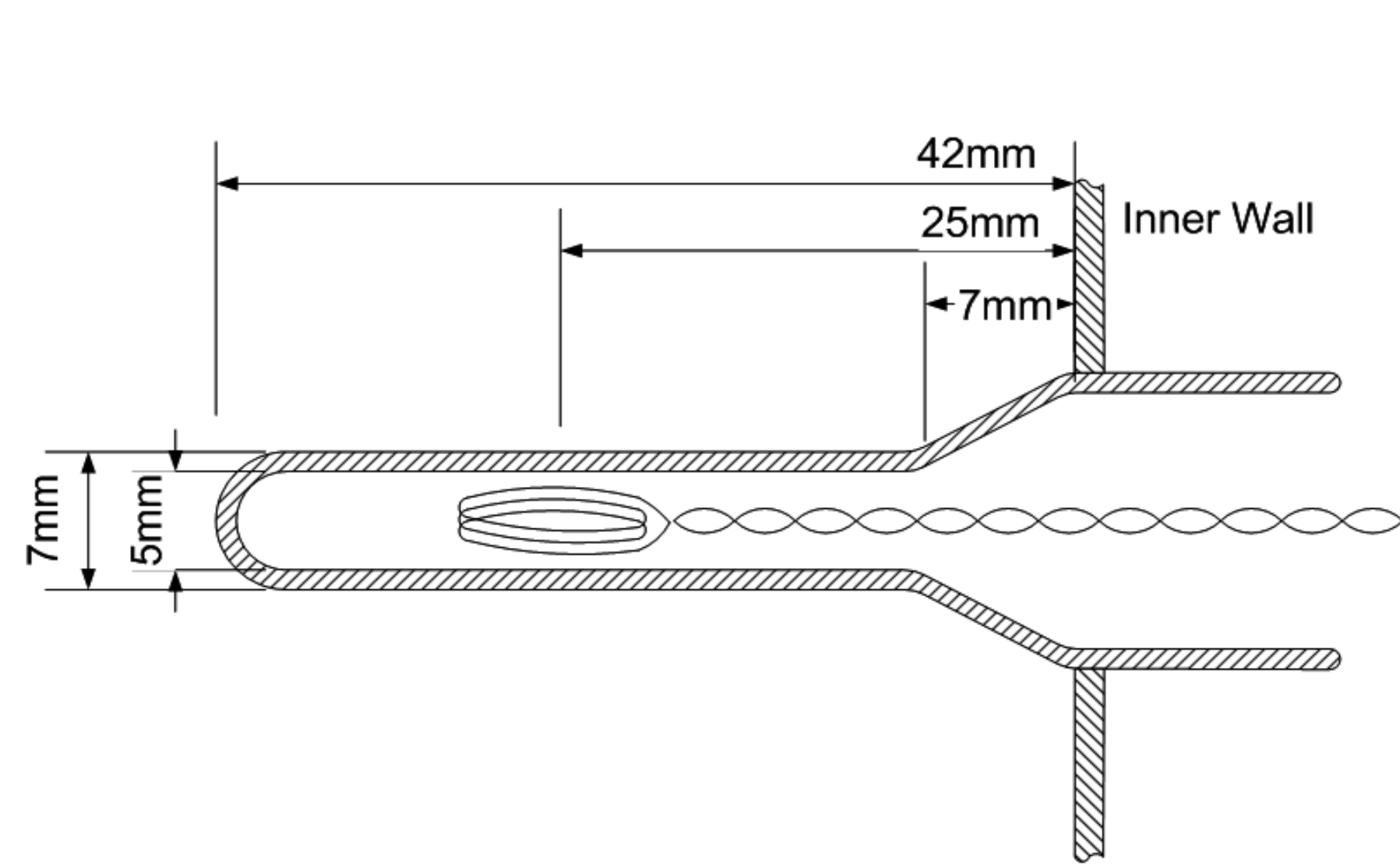


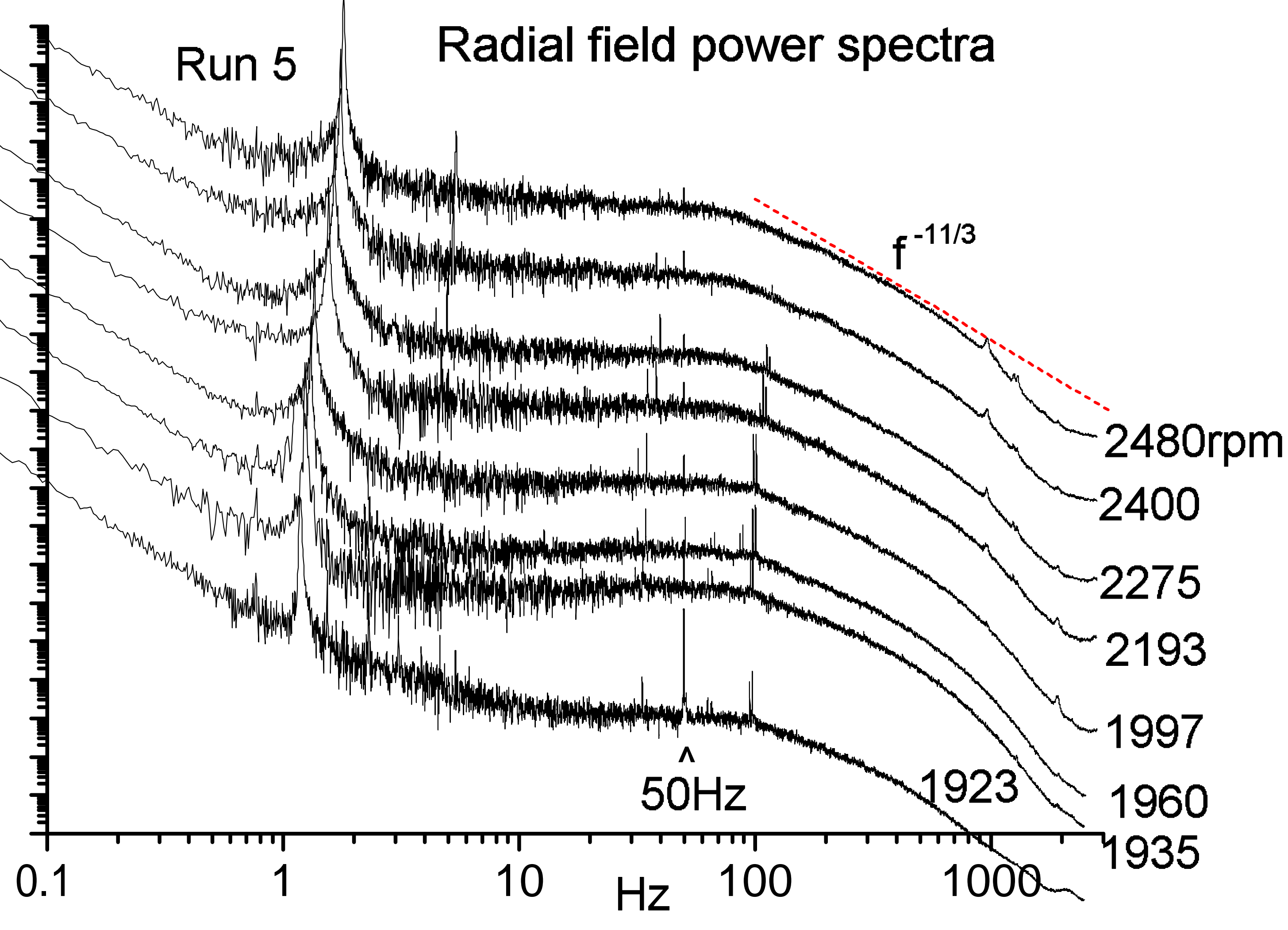
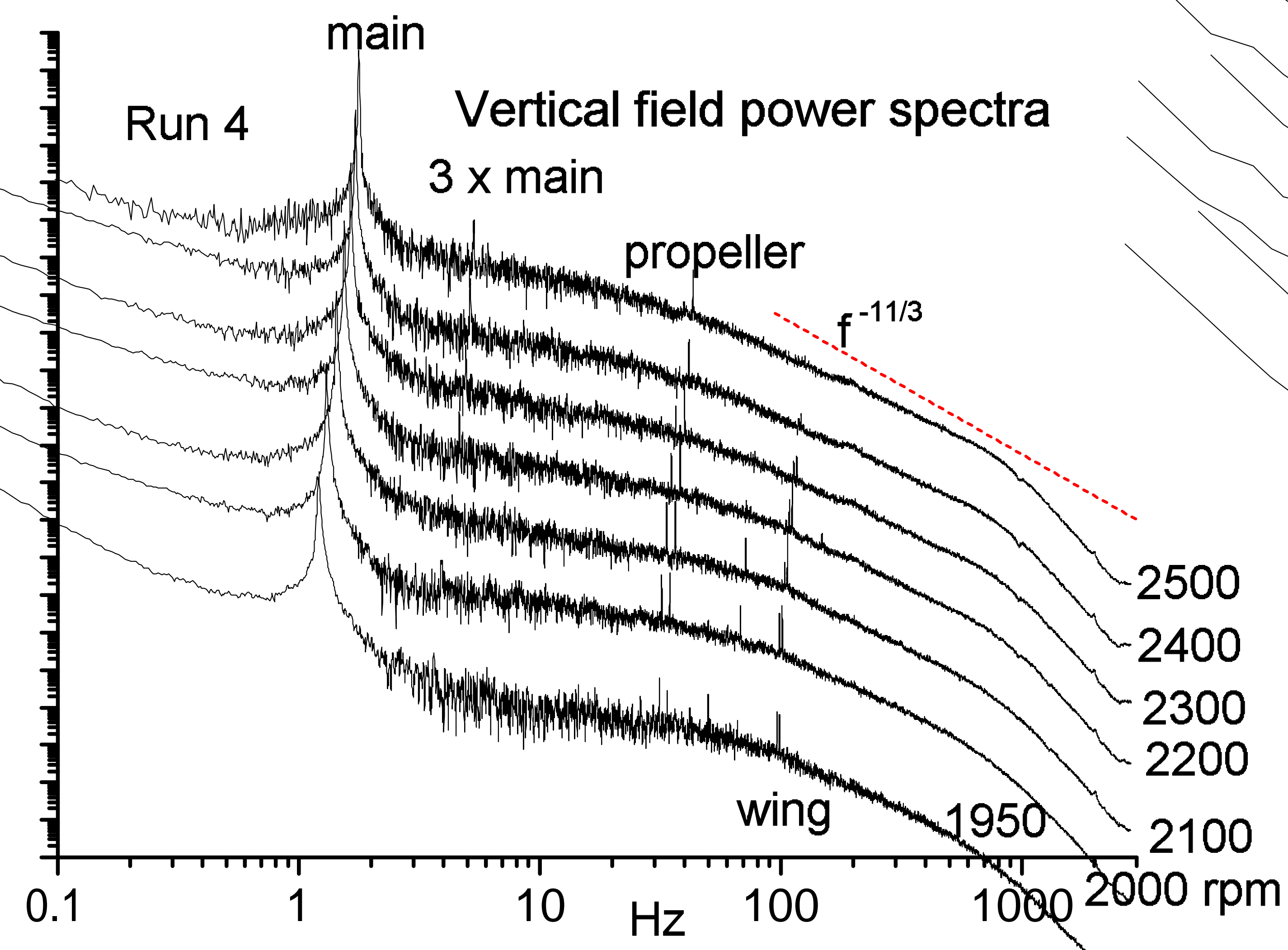


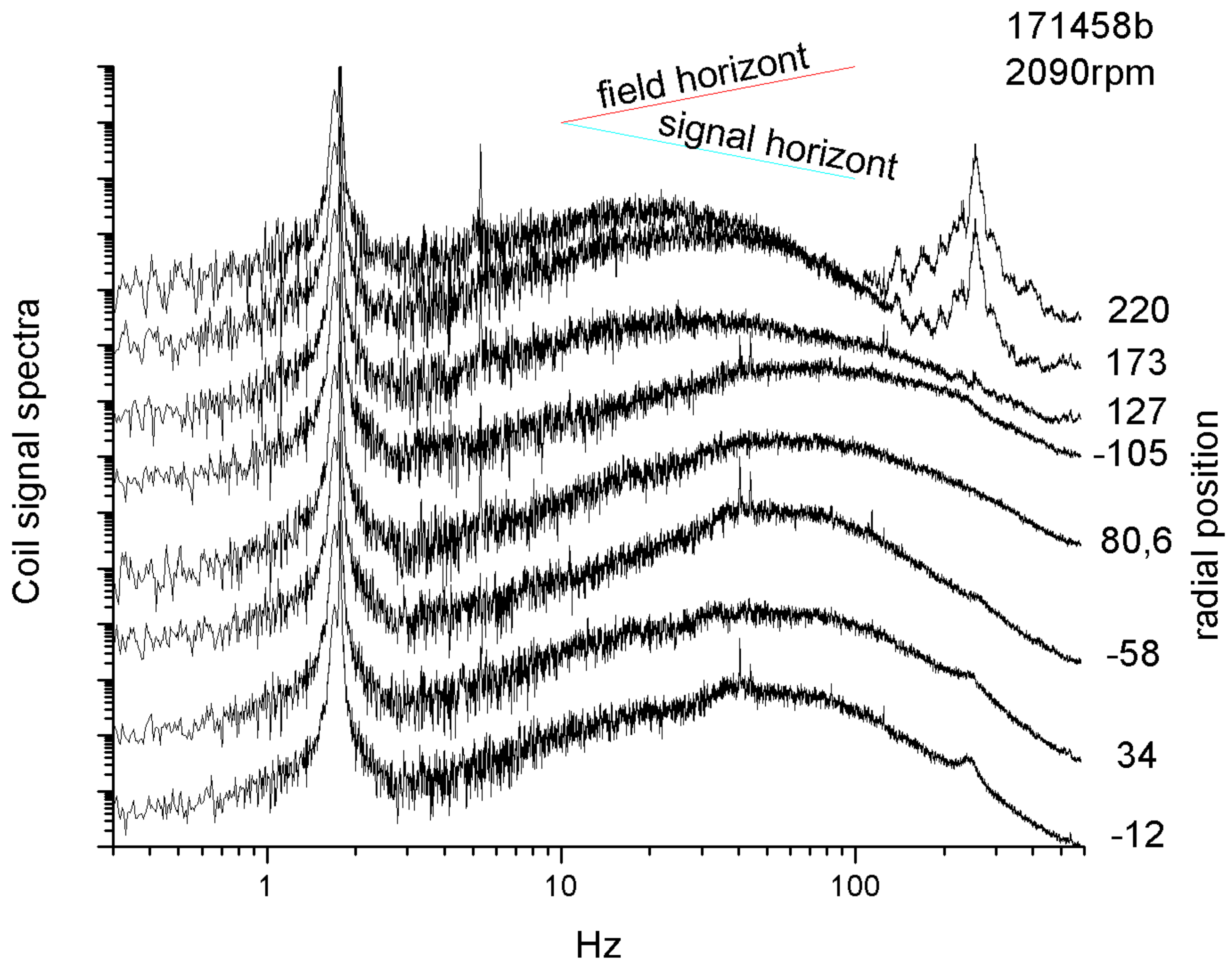


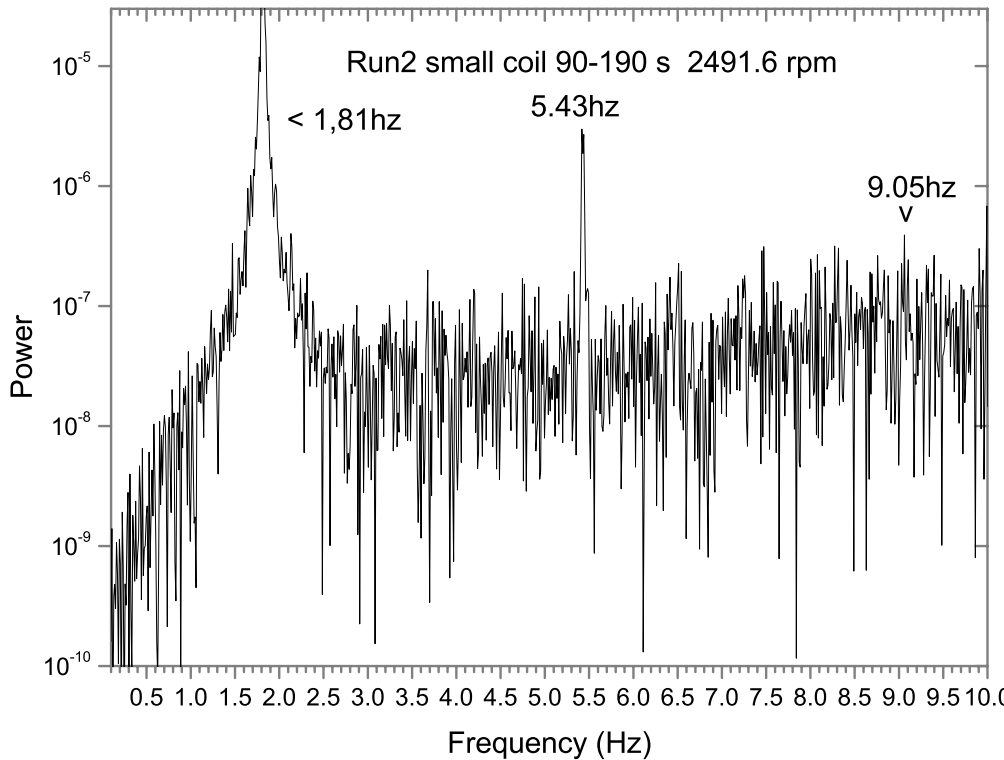




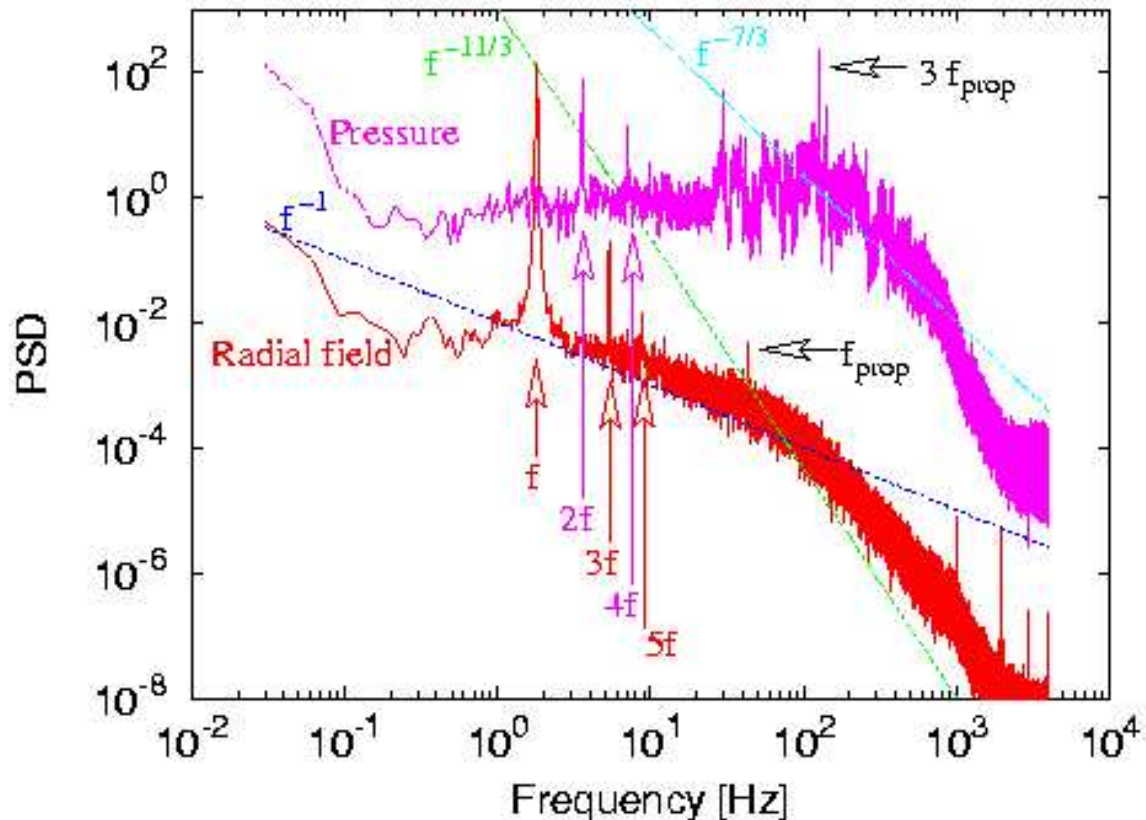


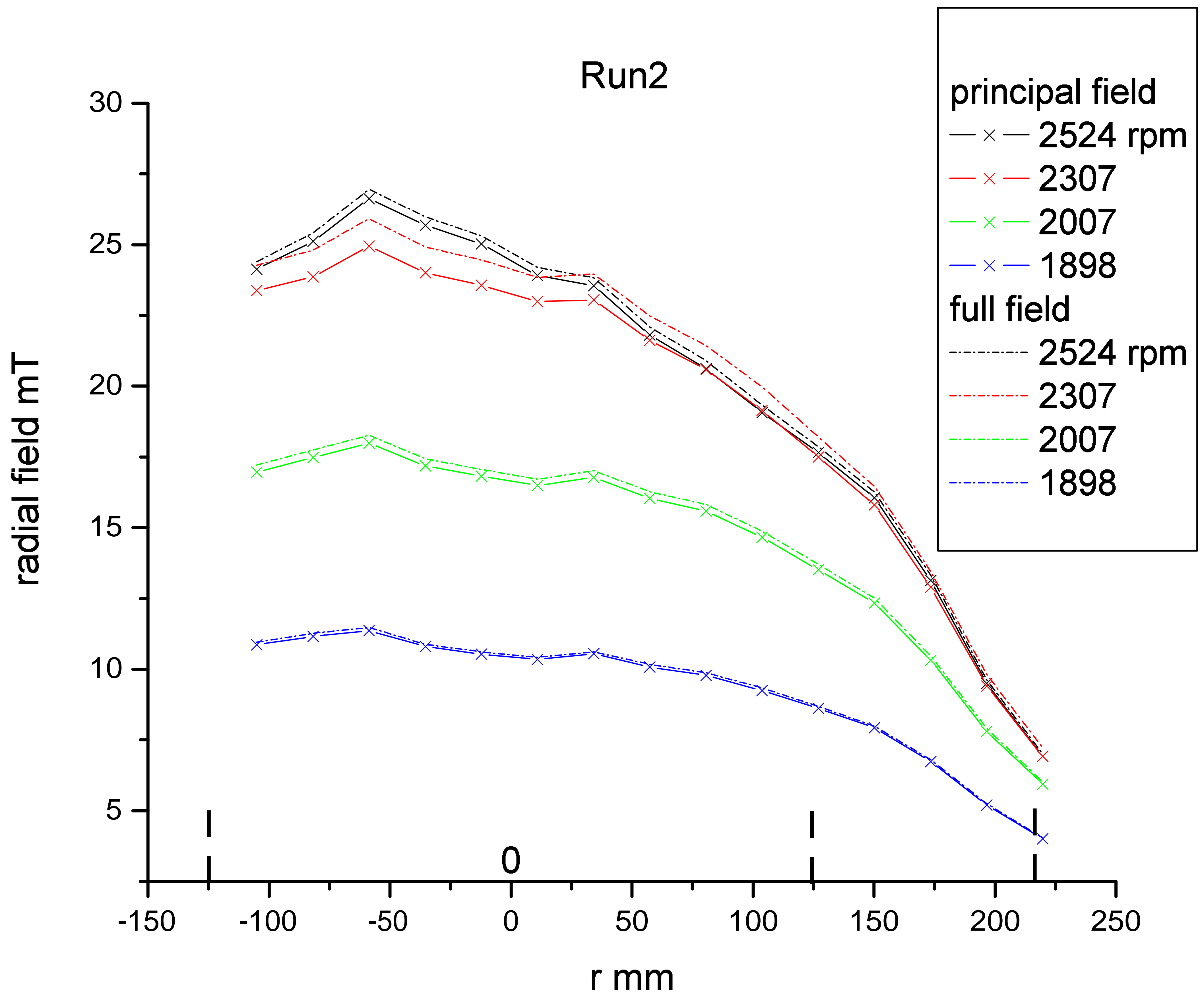




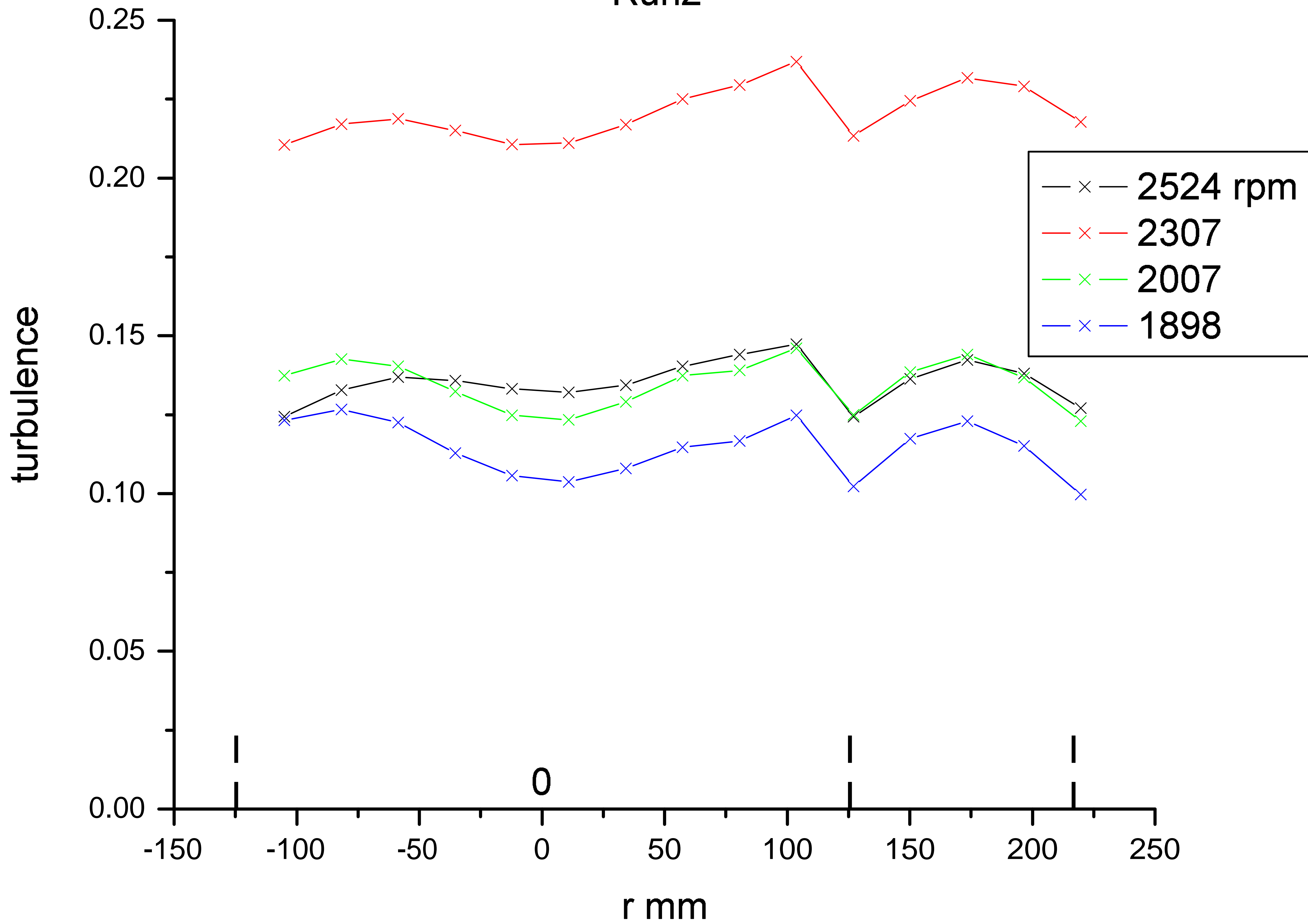


Run 1 - Spectra of radial field (tip) and pressure





Run2



β -- effect:

$$\begin{aligned}\sigma_T &= \sigma / (1 + (\mu_0 \sigma v^T \lambda)^2 / 9) \approx \\ &\approx \sigma / (1 + (B^T / B_0)^2 / 9) \approx \\ &\approx \sigma / (1 + (B_r^T / B_r \times B_r / B_z)^2 / 3) \approx \\ &\approx \sigma / (1 + (0.15/3)^2 / 3) \approx 0.999\sigma\end{aligned}$$

Energy balance at 2300 rpm:

Unperturbed kinetic energy:

in direct flow 5000 J/m

in rotation 1300 J/m

in back flow 2500 J/m

total 8800 J/m

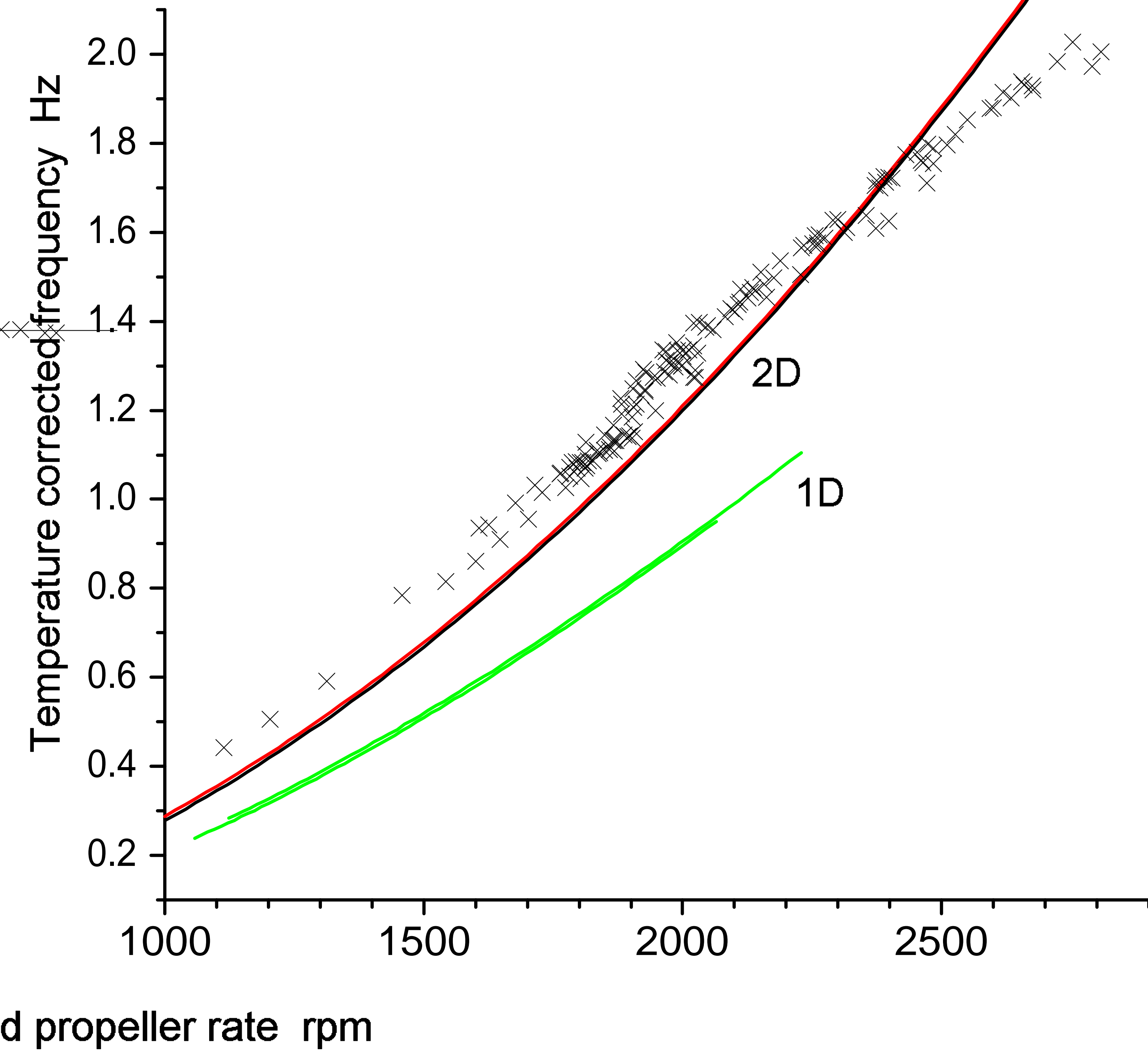
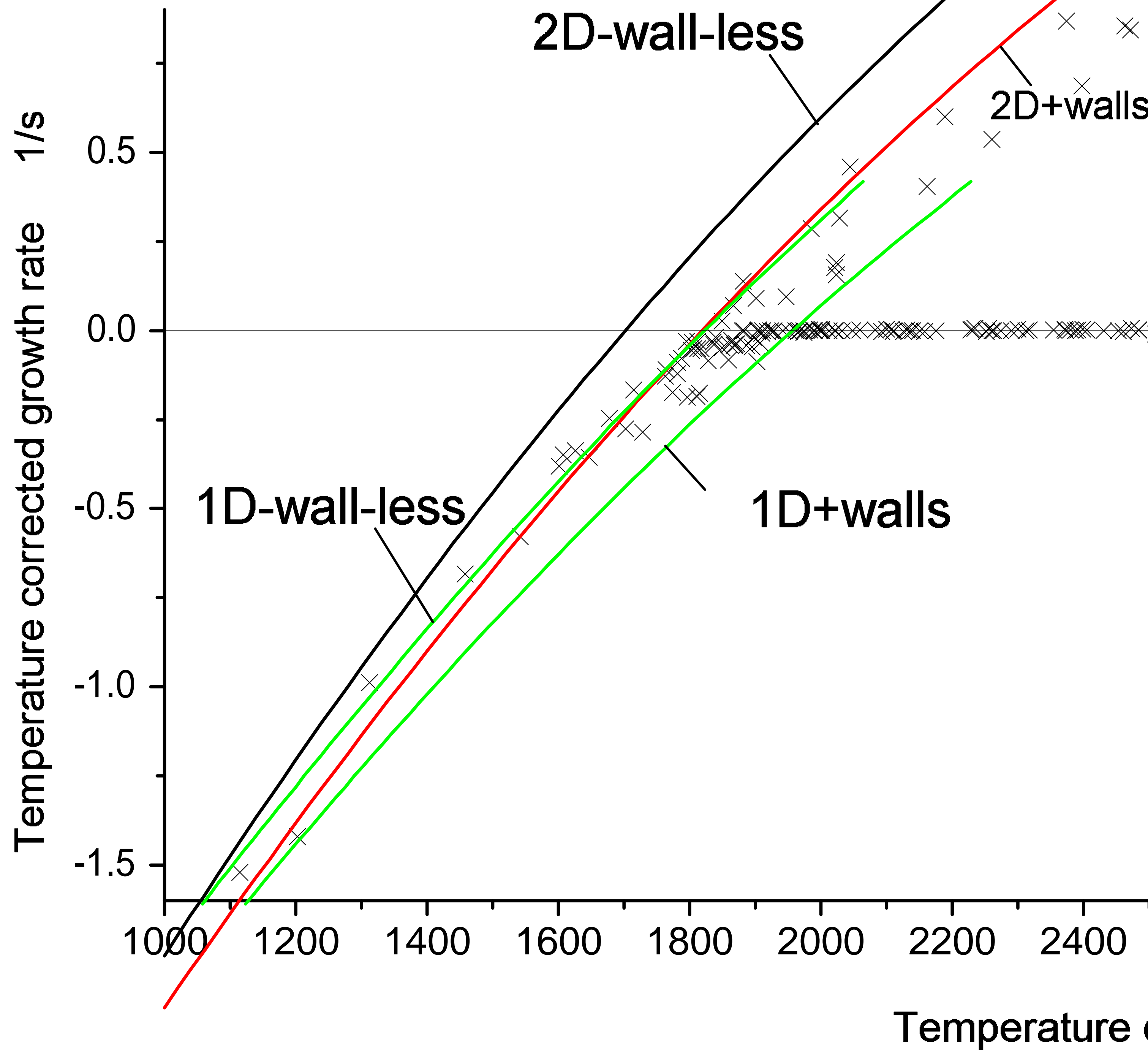
At measured level:

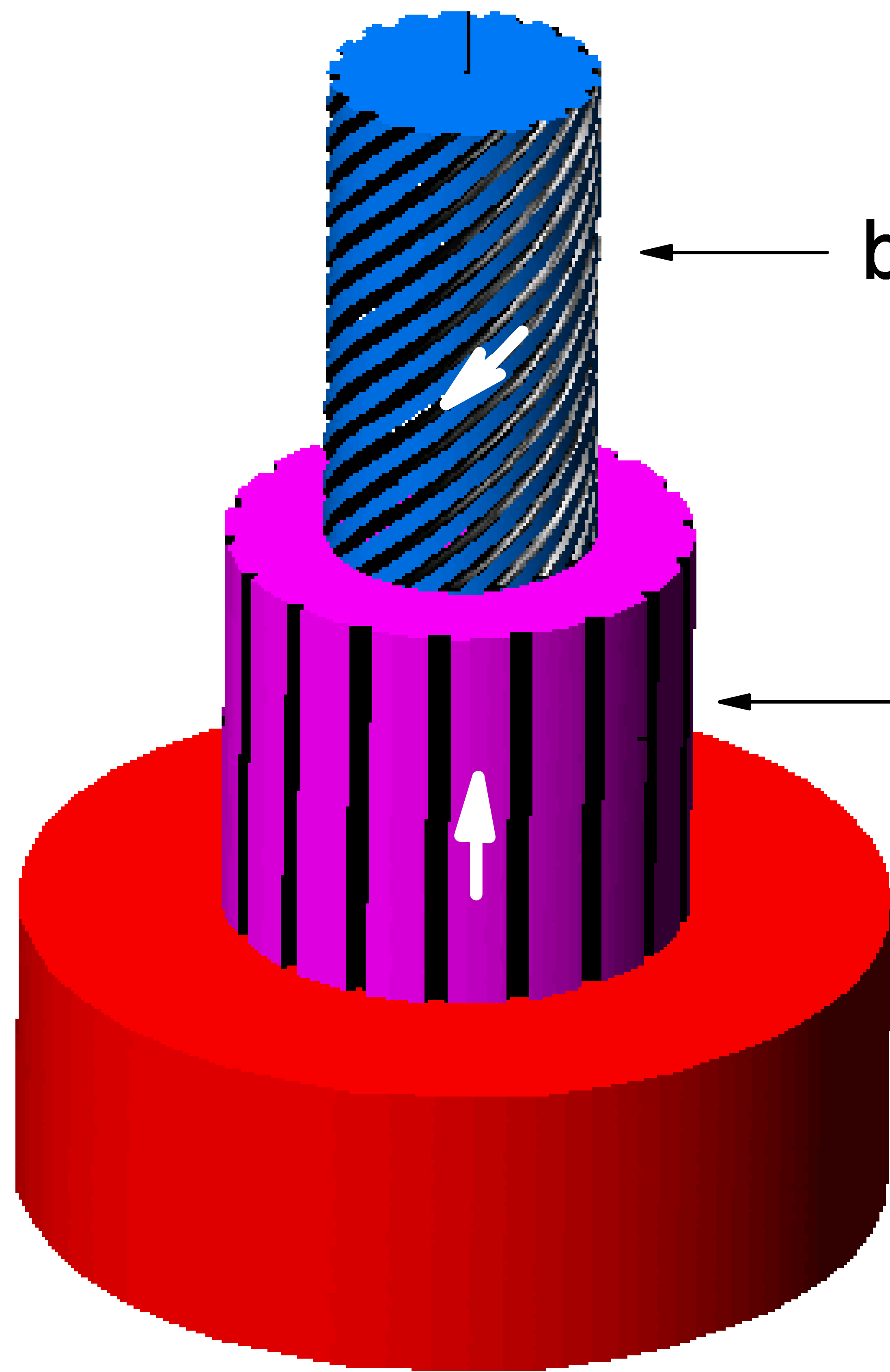
magnetic energy 175 J/m

Magnetic/kinetic 2%

$$\frac{\partial \vec{B}}{\partial t} = \nabla \times \vec{v} \times \vec{B} + \frac{1}{\mu_0 \sigma} \Delta \vec{B}$$

$$\vec{B}(\infty) \rightarrow 0$$





$$B_r \pm iB_\phi = b^\pm(r) \exp(i(kz + m\phi - \omega t))$$

← $b^\pm(r) = a_1^\pm I_{m\pm 1}(q_1 r)$ for solid velocity
or power series for
polynomial velocity profiles

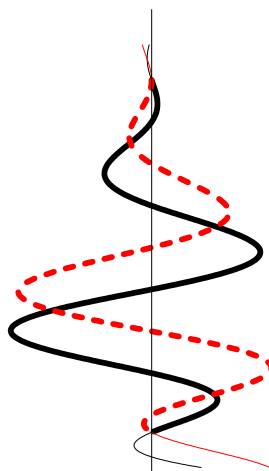
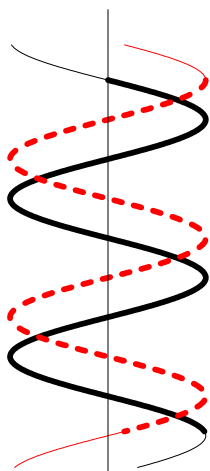
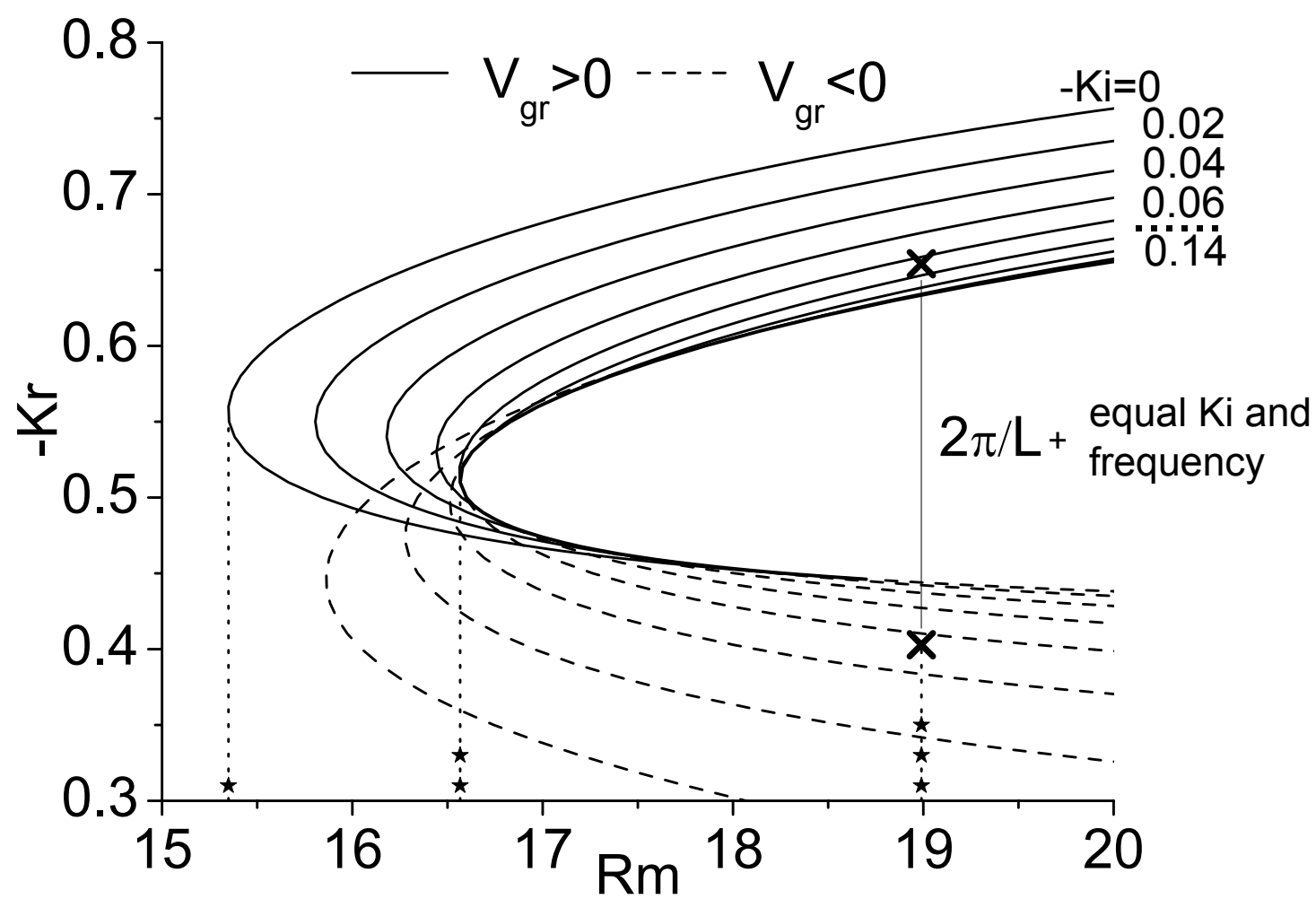
← $= a_2^\pm I_{m\pm 1}(q_2 r) + c_2^\pm K_{m\pm 1}(q_2 r)$

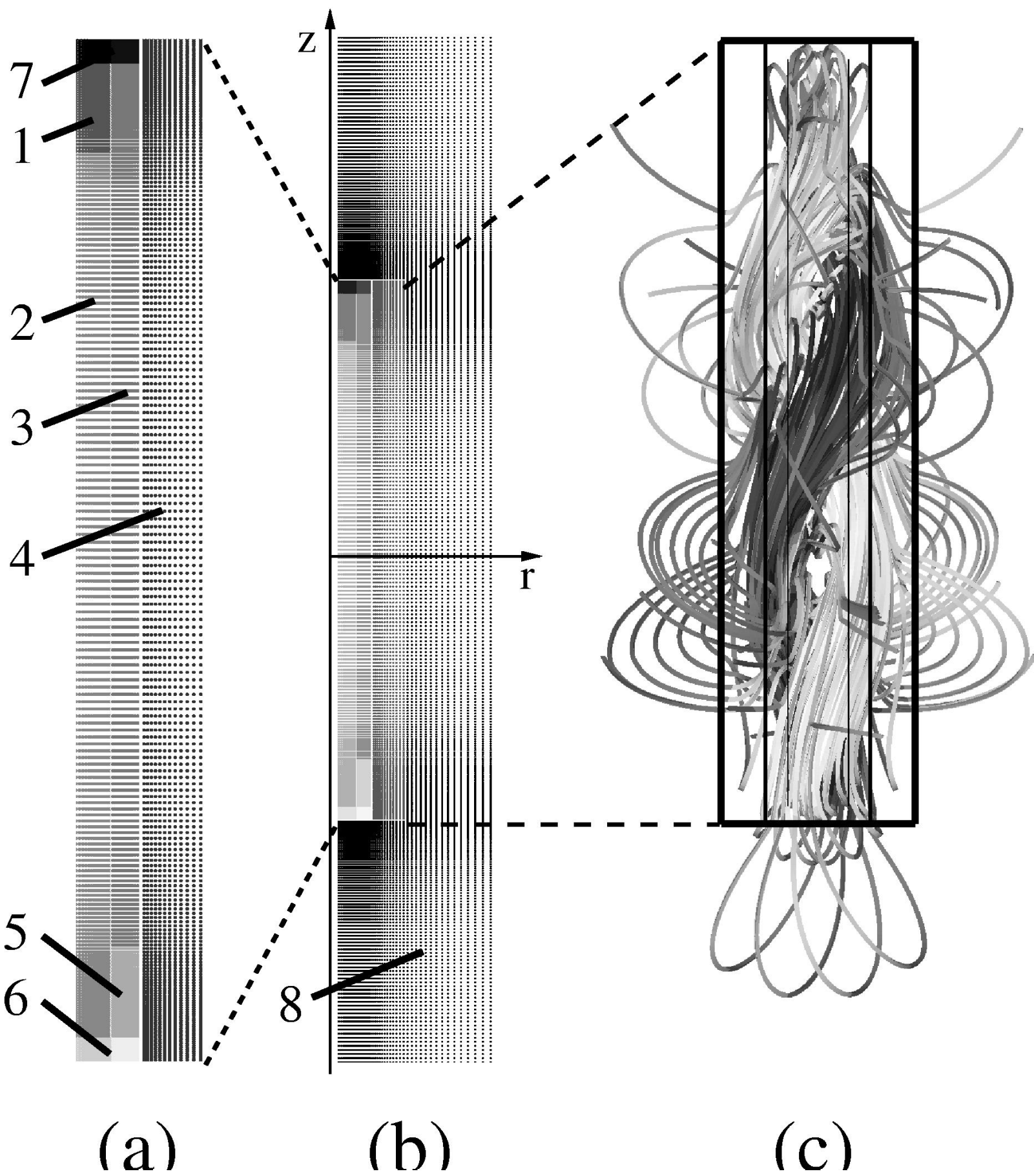
← $= a_3^\pm I_{m\pm 1}(q_3 r) + c_3^\pm K_{m\pm 1}(q_3 r)$

← $= c_4^\pm K_{m\pm 1}(kr)$ in outer space

$$q_n^2 = k^2 + i\mu_0 \sigma \omega_n$$

ω_n - frequency in co-moving frame





END