

# Predictability in a simulated global prediction system that assimilates only surface pressure observations

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Special thanks to:

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Tim Hoar

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Data Assimilation Research Testbed (DART) update

Ensemble filter assimilation

Model description

Many assimilation system simulation experiments

Conclusions

## The Data Assimilation Research Testbed (DART)

### What is DART?

1. Allows combinations of assimilation algorithms, models, and observation sets
2. Diagnostic tools
3. Supports Data Assimilation R&D for NOAA/NCAR and external partners  
NOT for operational use or support

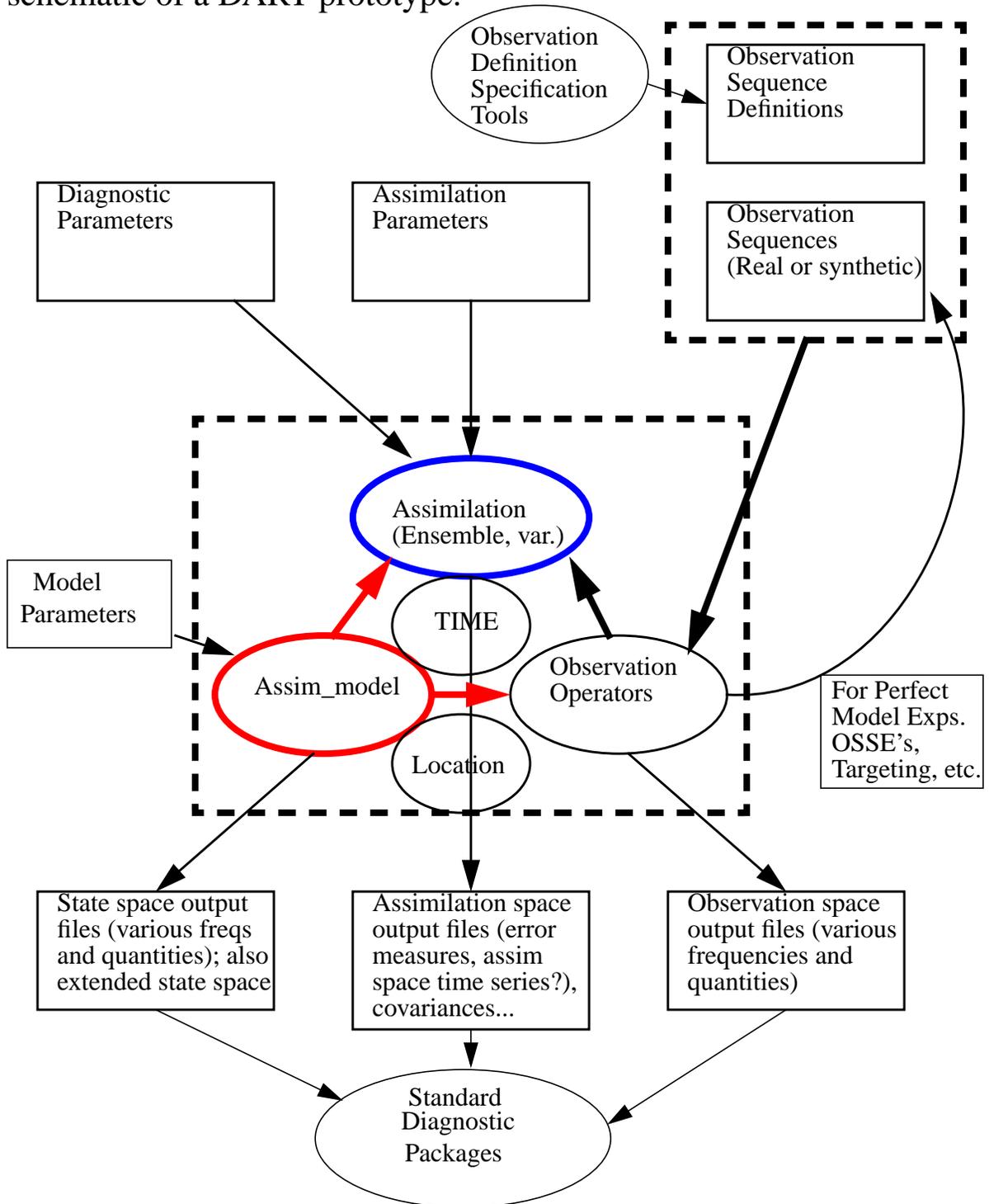
### Status of DART

1. Basic framework implemented
2. Currently using GFDL FMS infrastructure
3. Switch to ESMF infrastructure when available
4. Primarily implementing ensemble (Kalman) filters
5. Variational for low-order models only
6. Plans MAY include a variational (4D-Var) capability

### DART compliant models

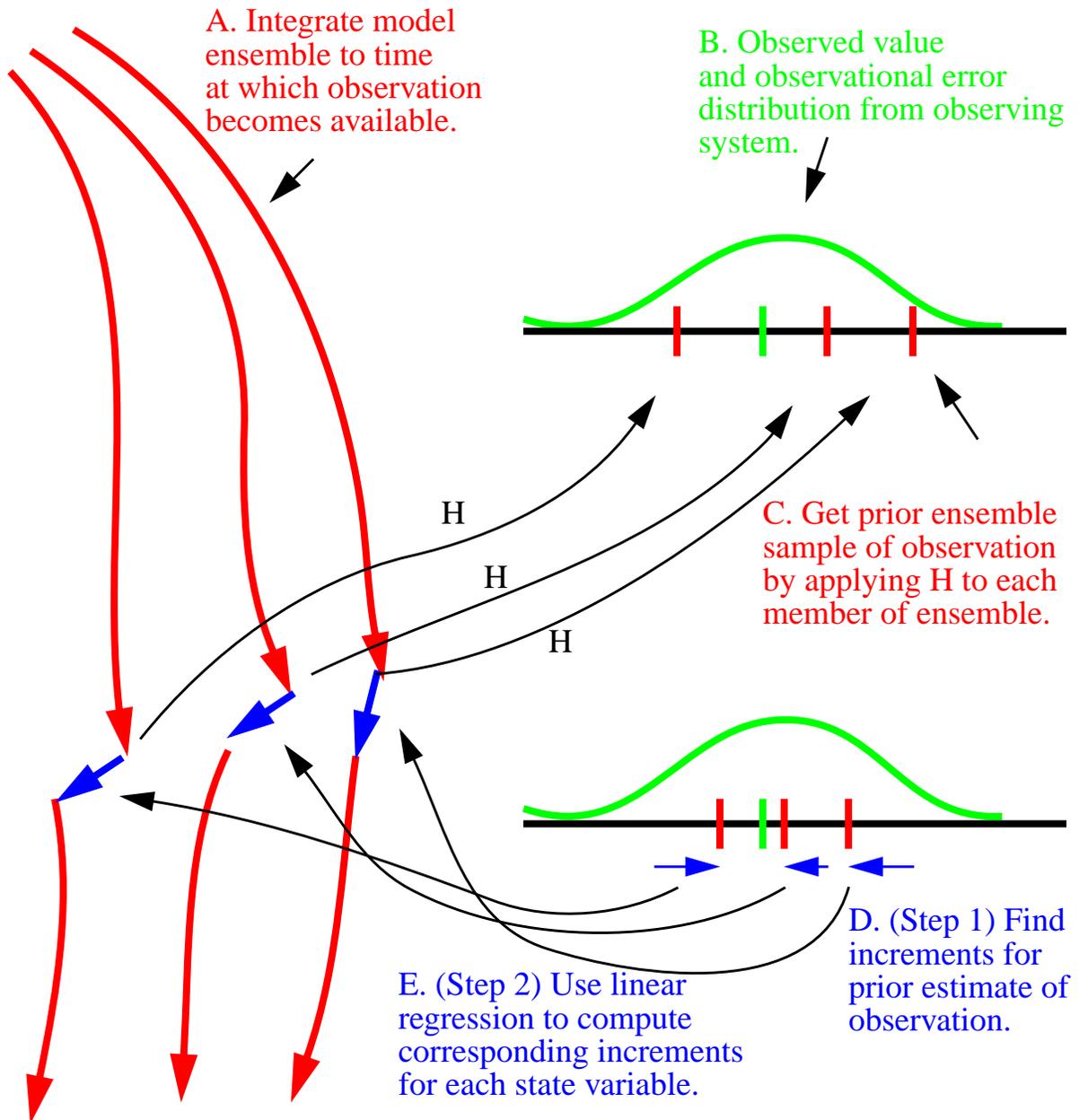
1. GFDL FMS B-grid GCM incorporated and in use
2. Many low-order models available
3. WRF model in process of being incorporated
4. NCEP MRF being tested quasi-operationally in partial implementation
5. GFDL MOM ocean model partially incorporated in earlier version
6. Initial work on incorporating CAM

A schematic of a DART prototype.



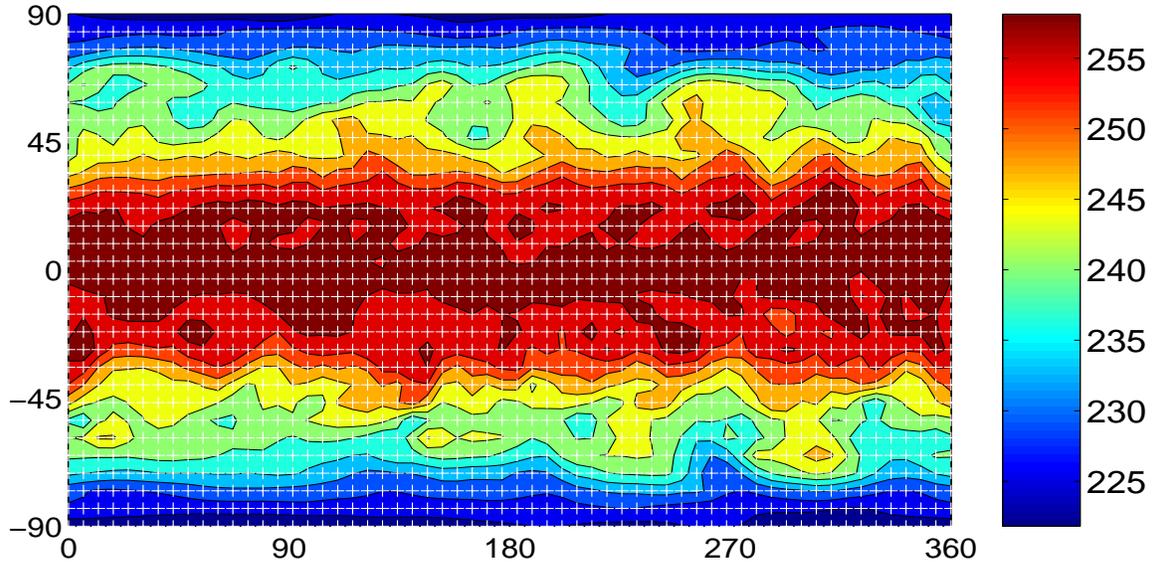
## How an Ensemble Filter Works

Theory: Impact of observations can be handled sequentially  
 Impact of observation on each state variable can be handled sequentially



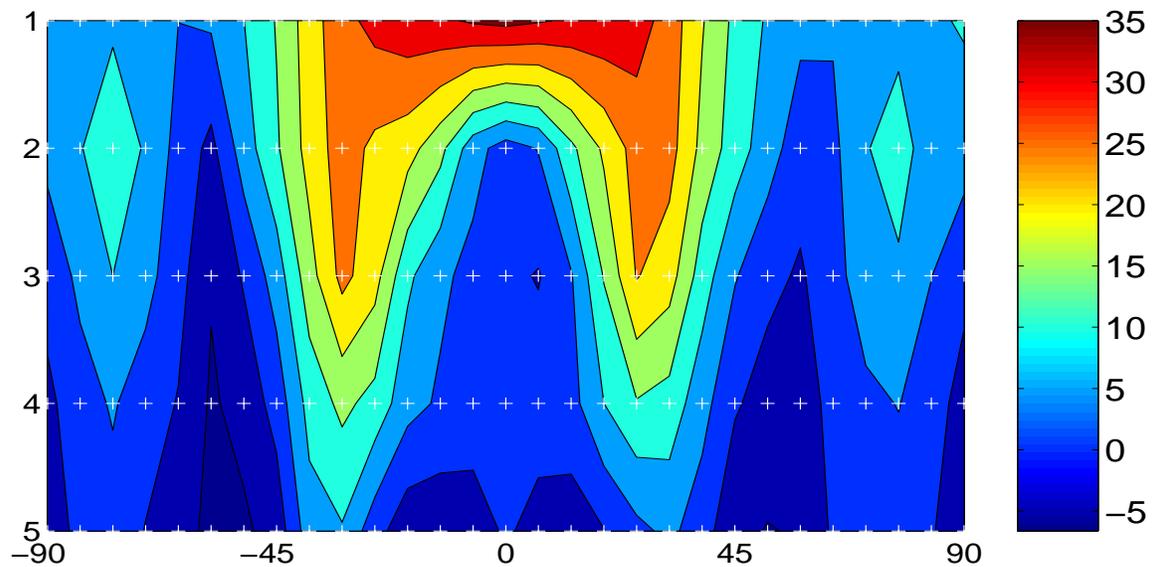
GFDL FMS B-Grid Dynamical Core (Havana)  
 Held-Suarez Configuration (no zonal variation, fixed forcing)

T at Level 3 (Middle of Atmosphere)

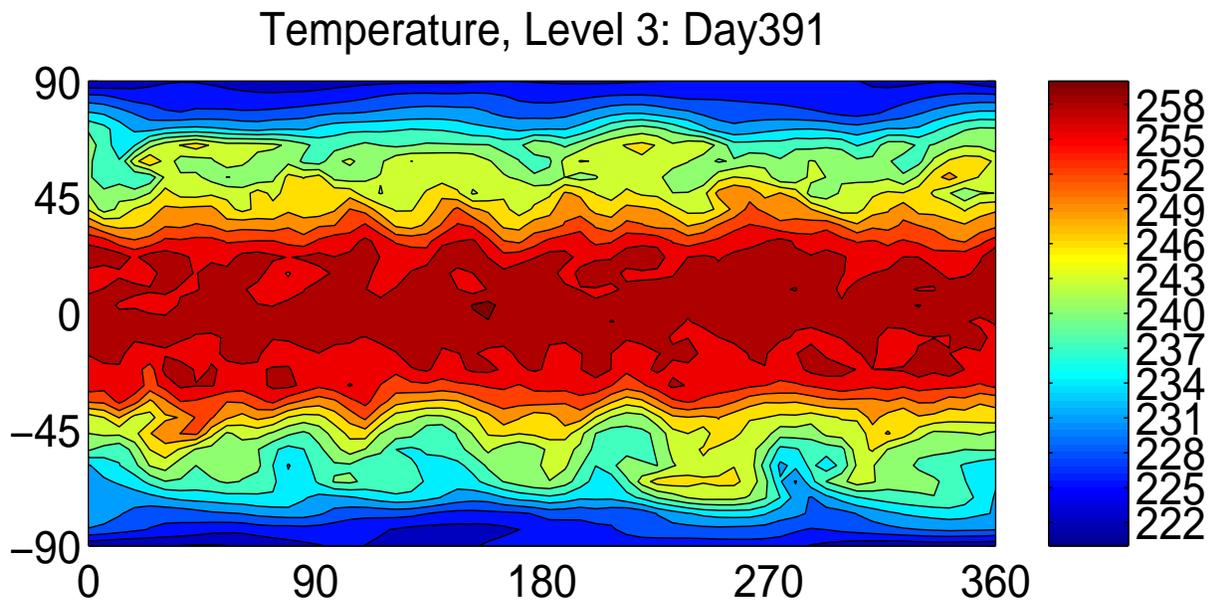
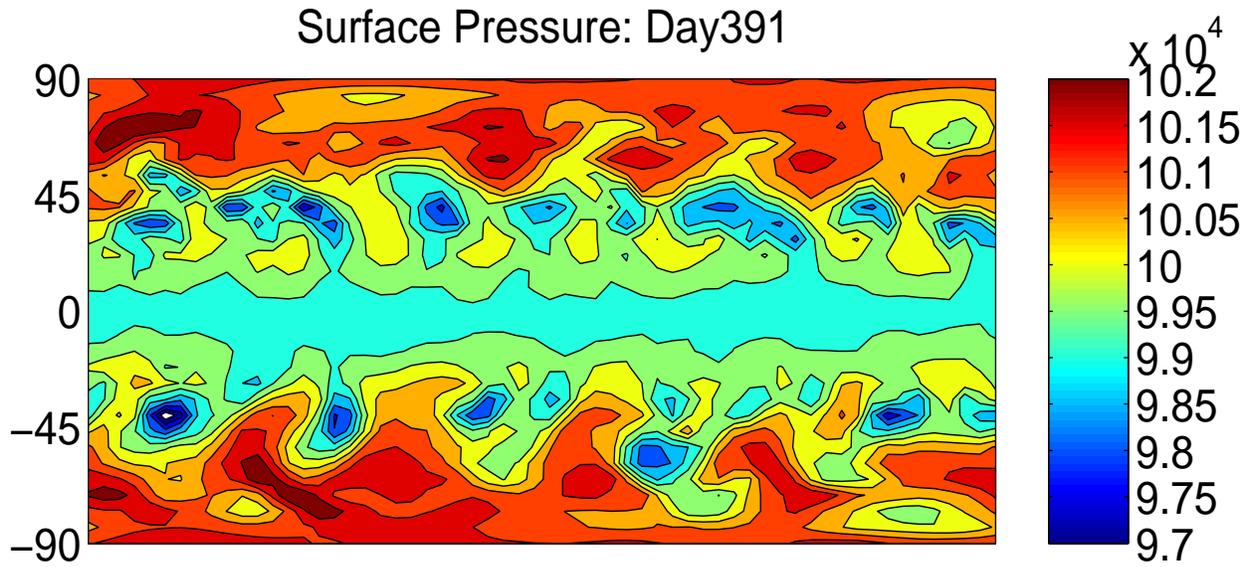


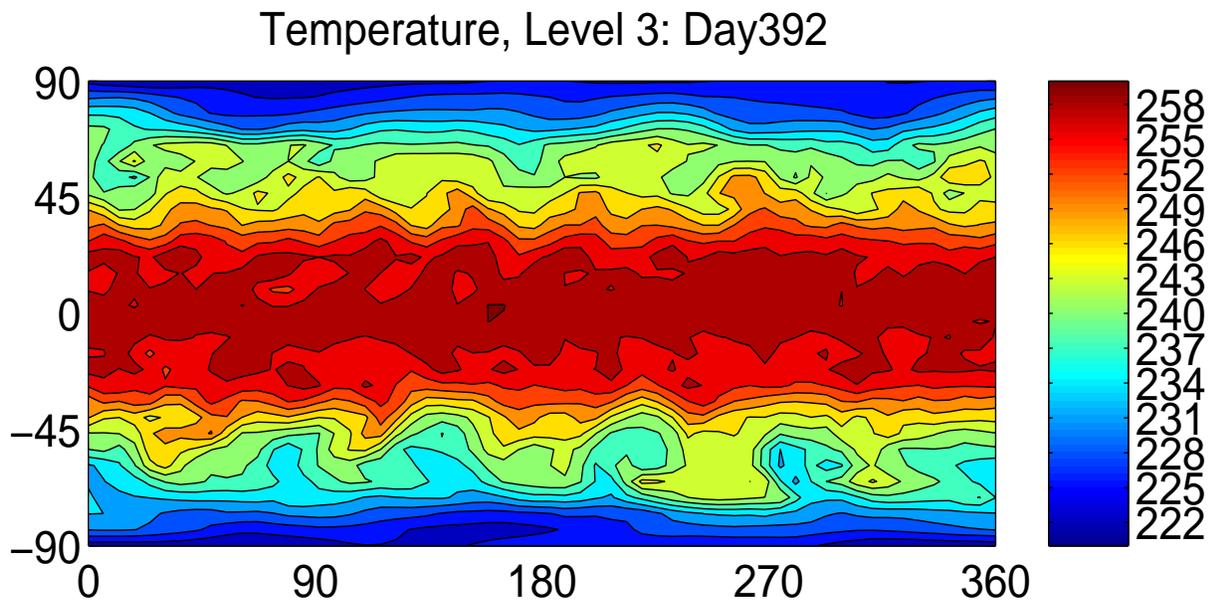
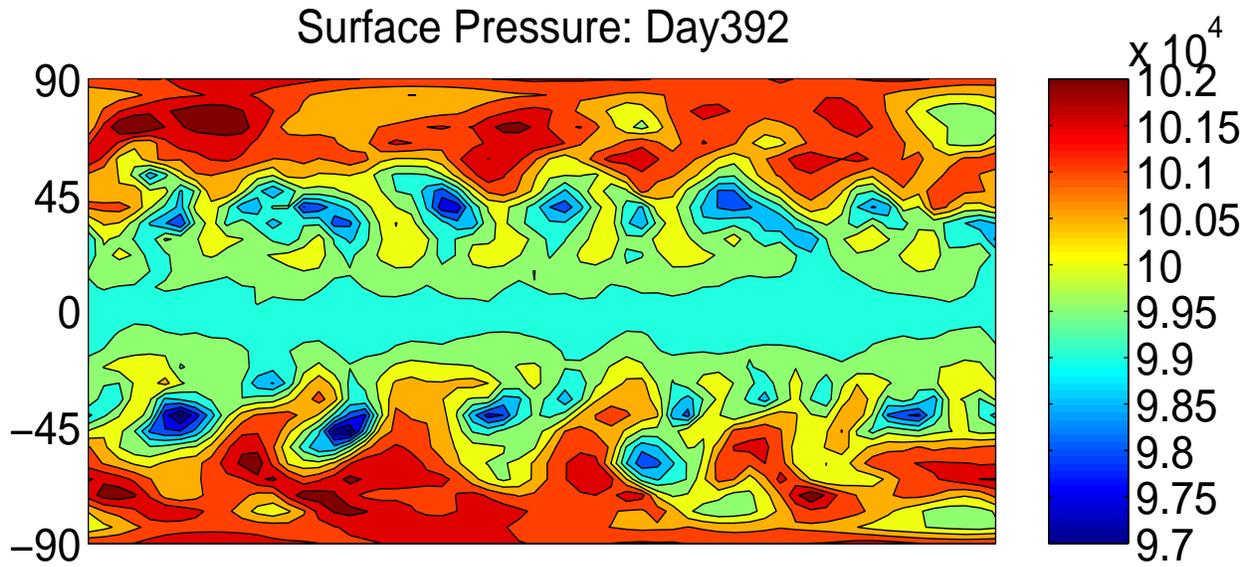
Low-Resolution (60 longitudes, 30 latitudes, 5 levels)  
 Damping coefficients reduced to 0.10 for error growth  
 Timestep 1 hour (or less for frequent observations)

Cross section of Zonal mean U

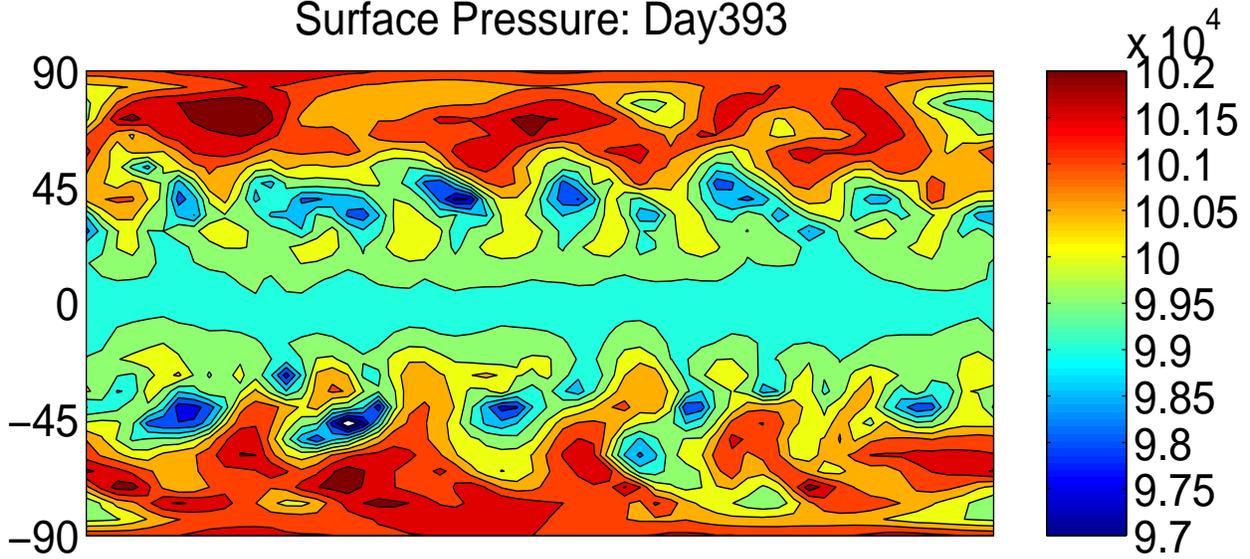


Has baroclinic instability

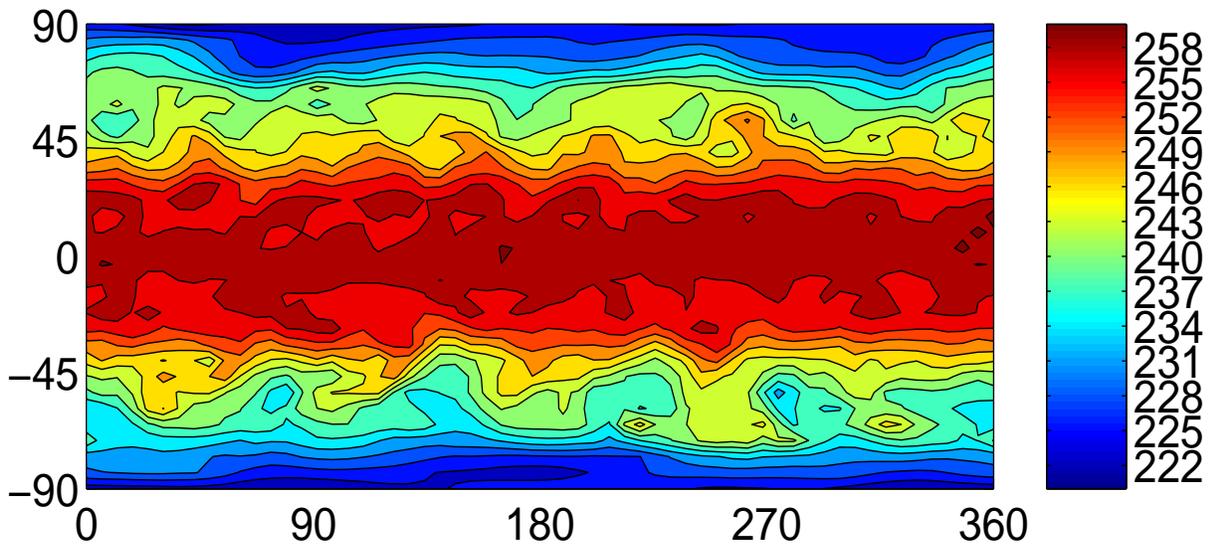




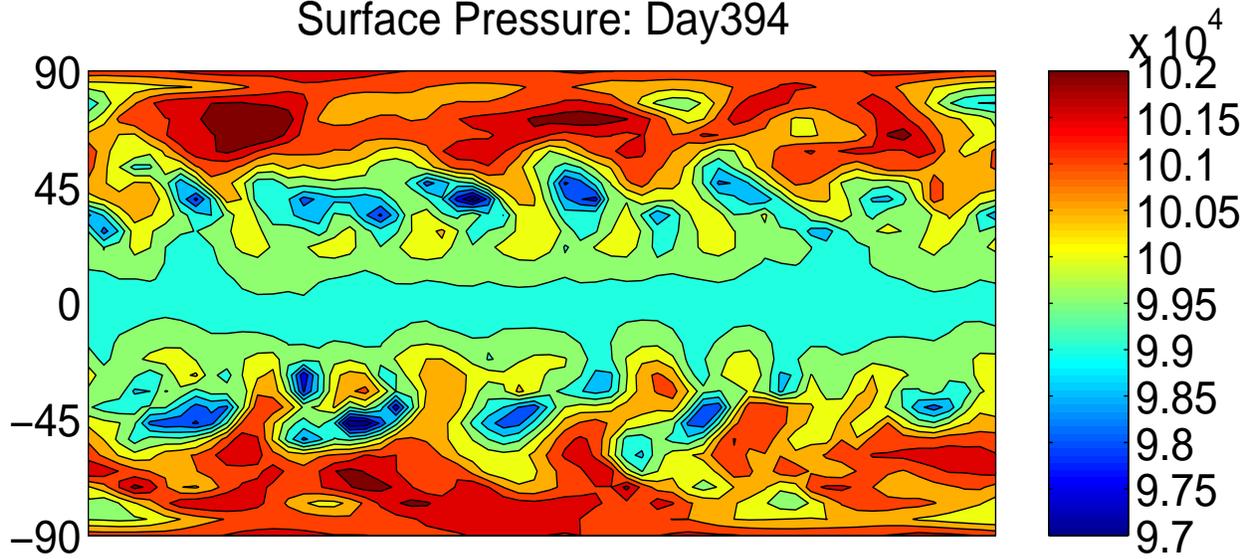
Surface Pressure: Day393



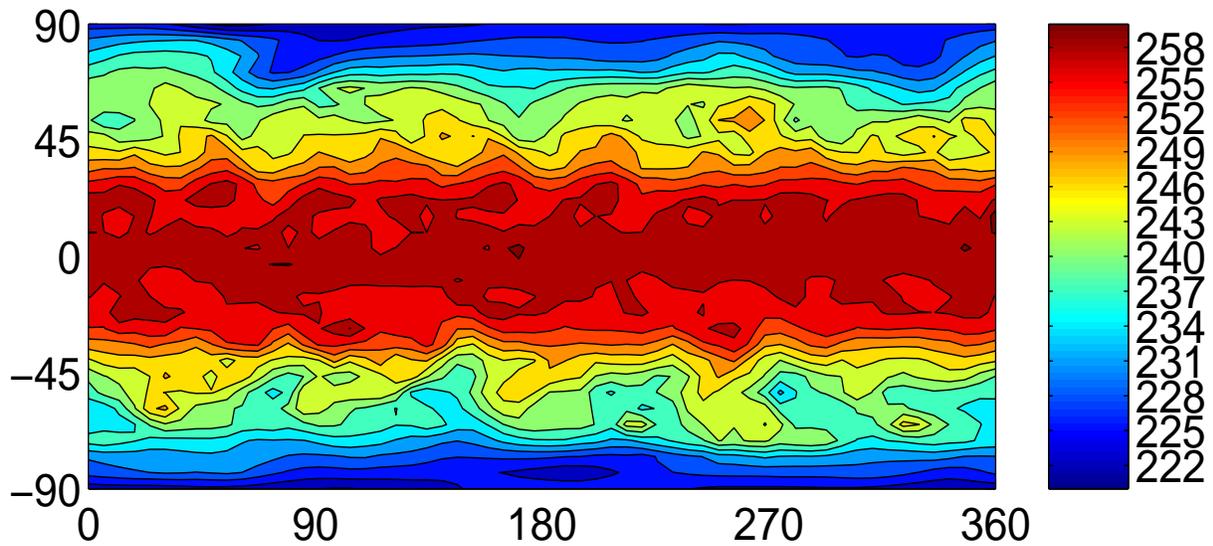
Temperature, Level 3: Day393

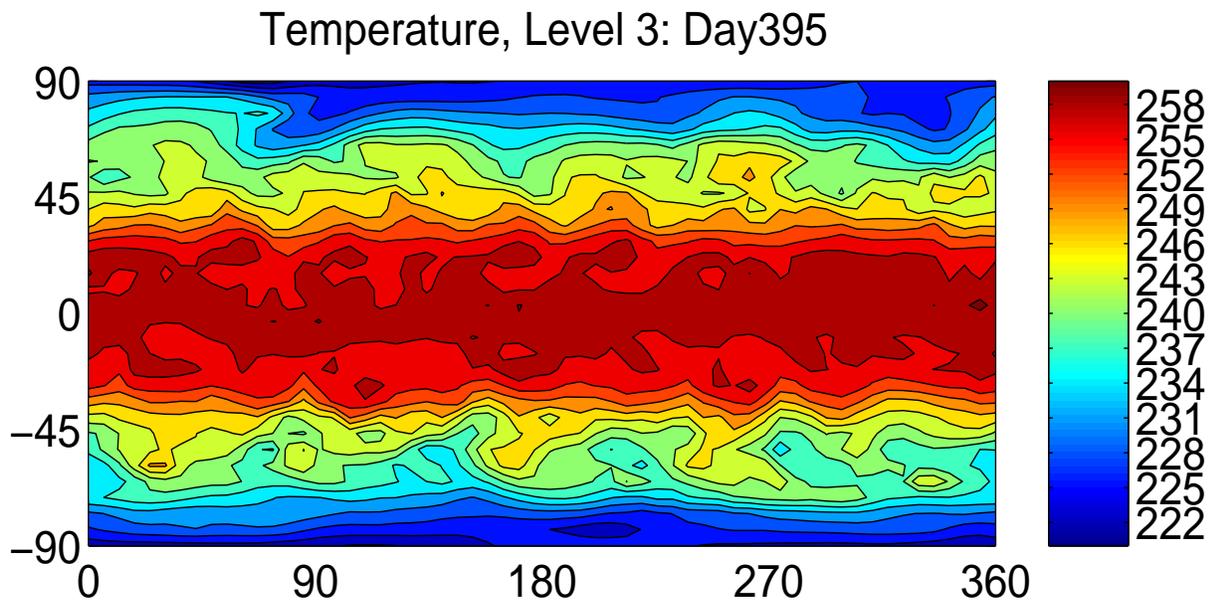
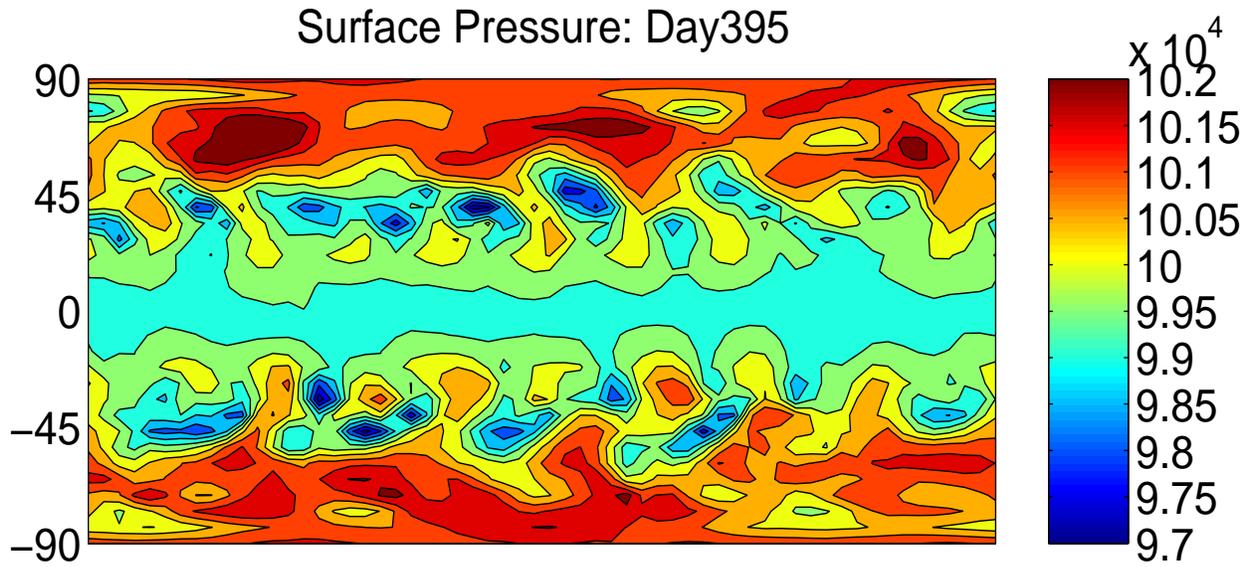


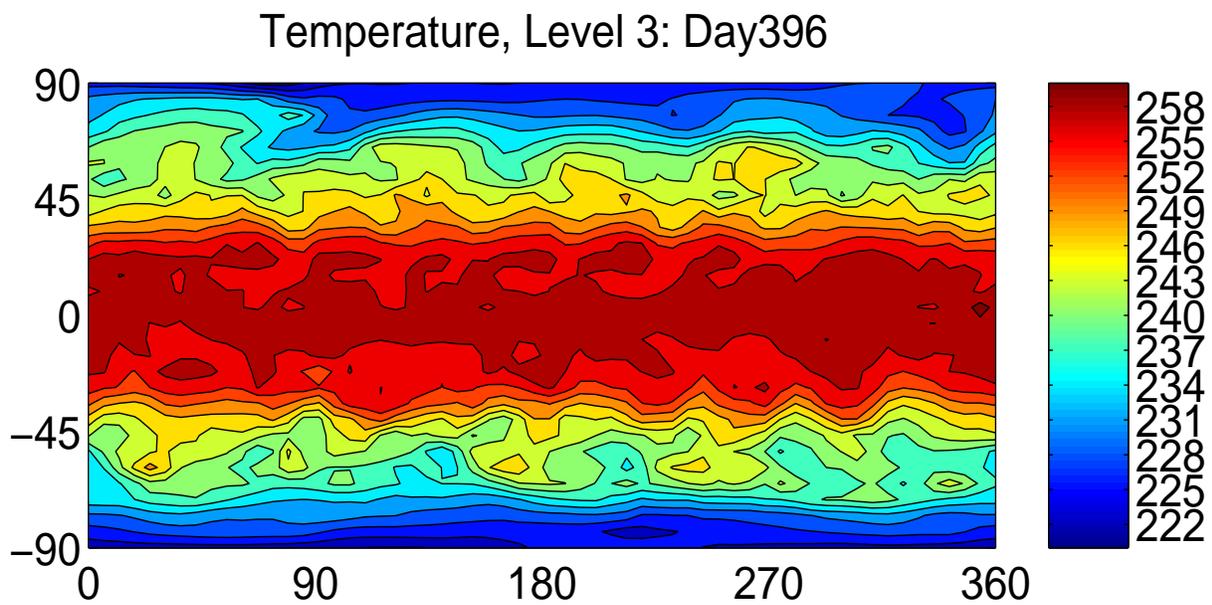
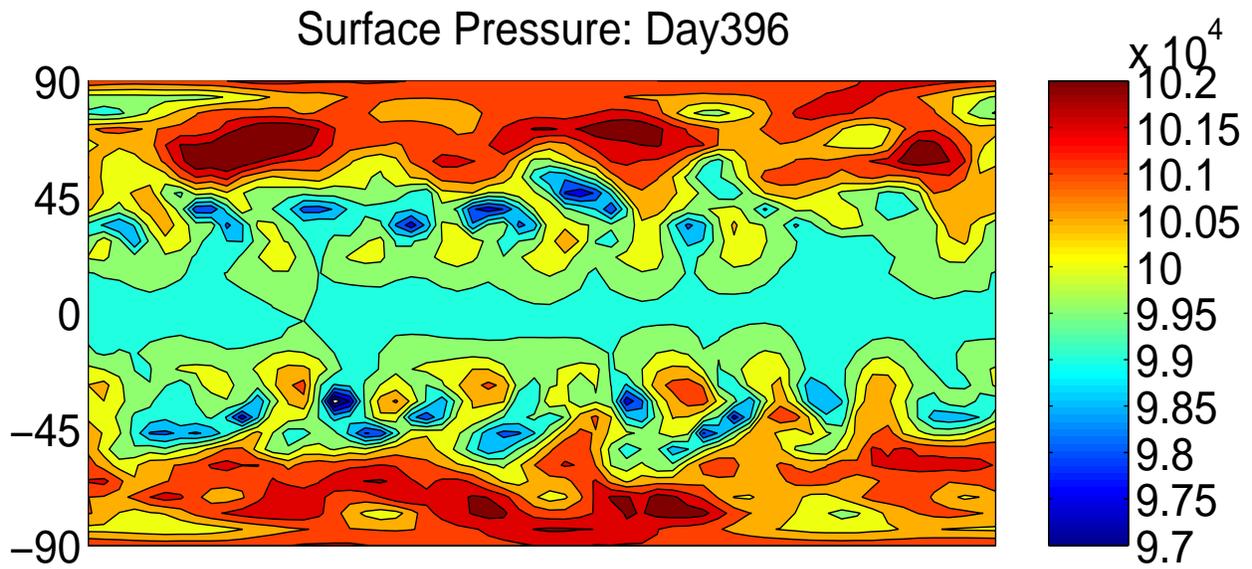
Surface Pressure: Day394

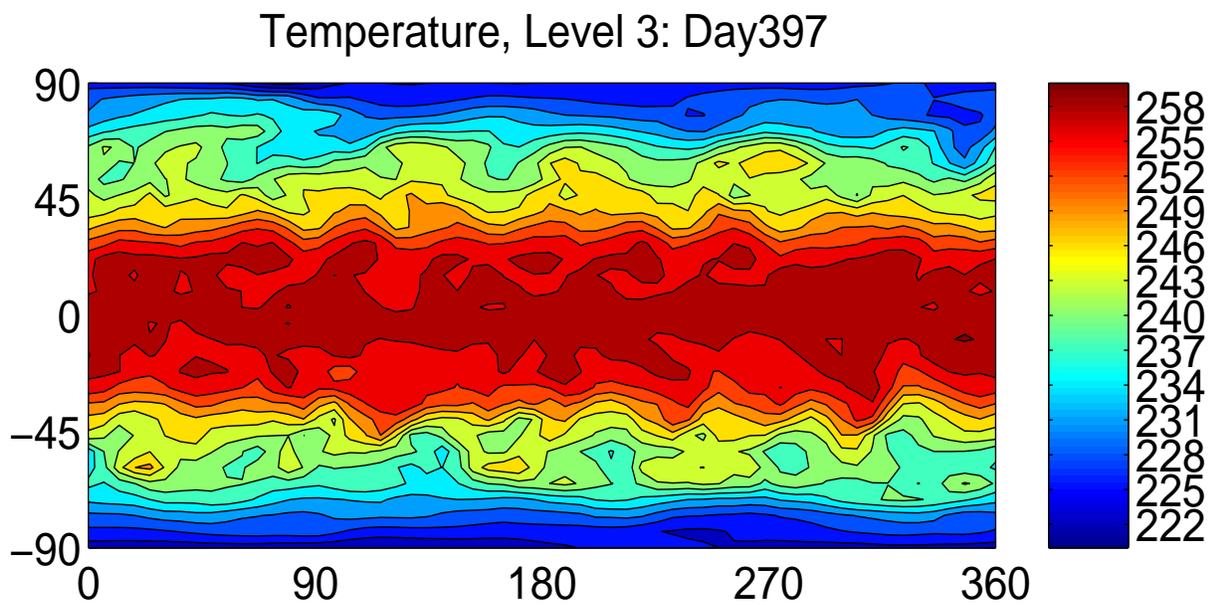
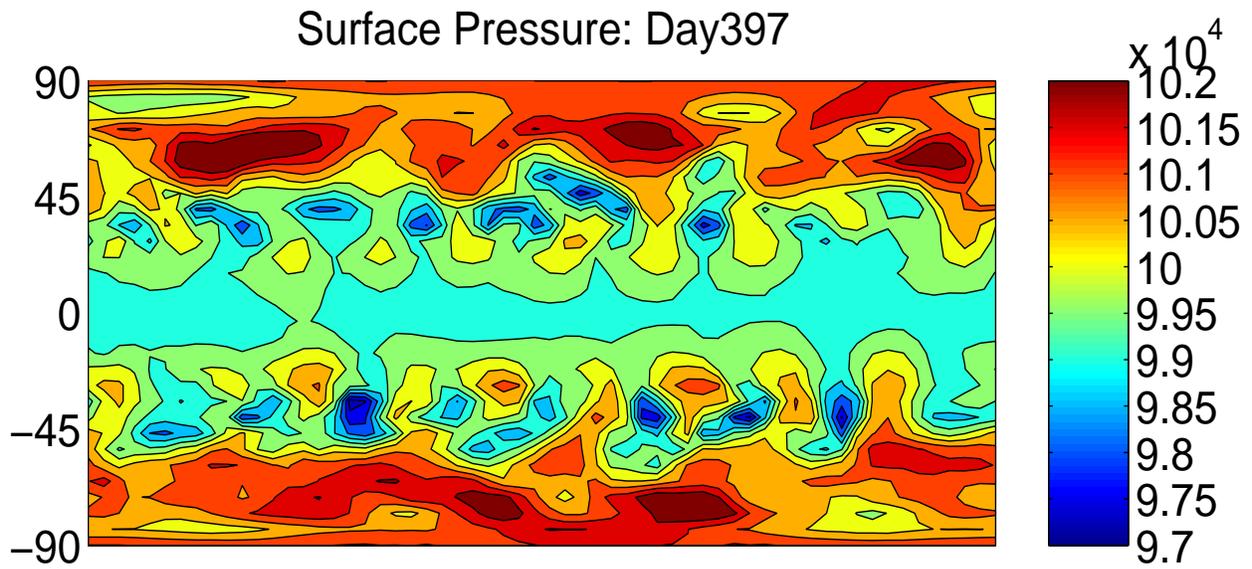


Temperature, Level 3: Day394

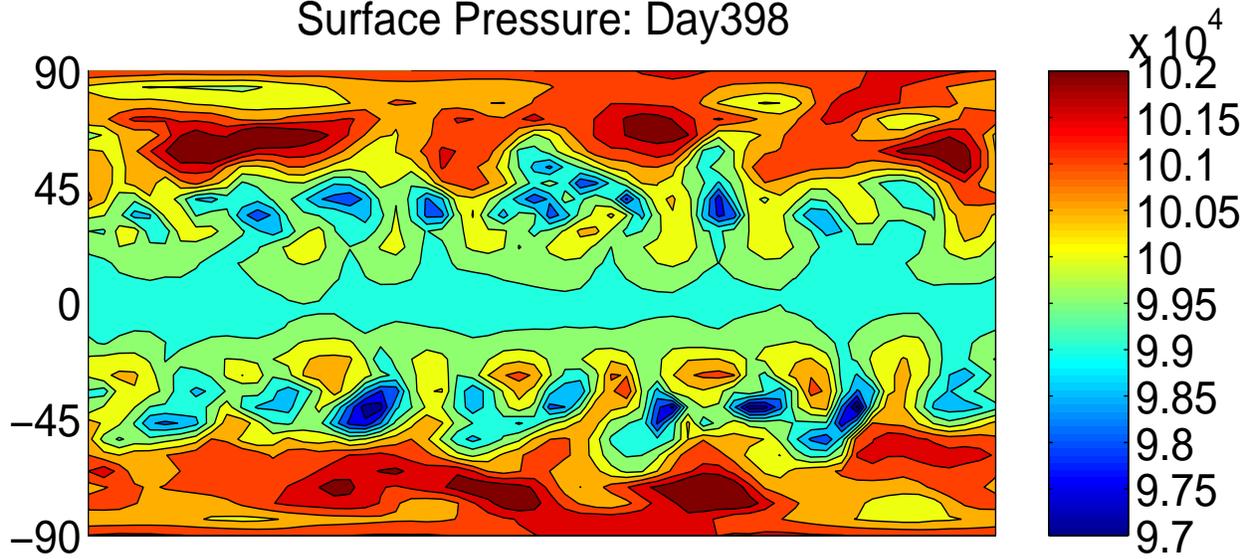




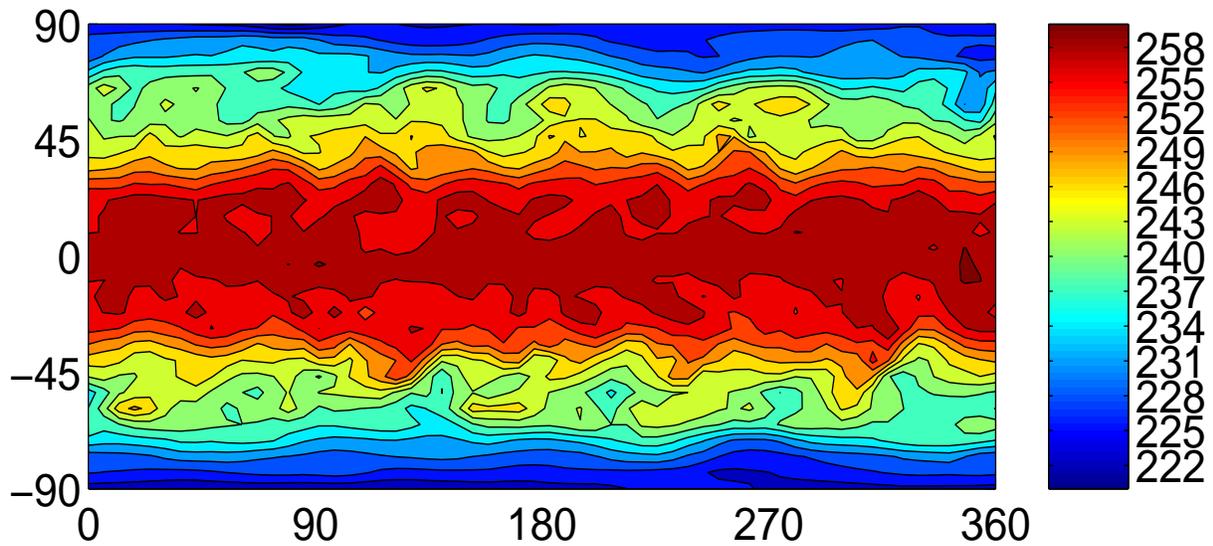


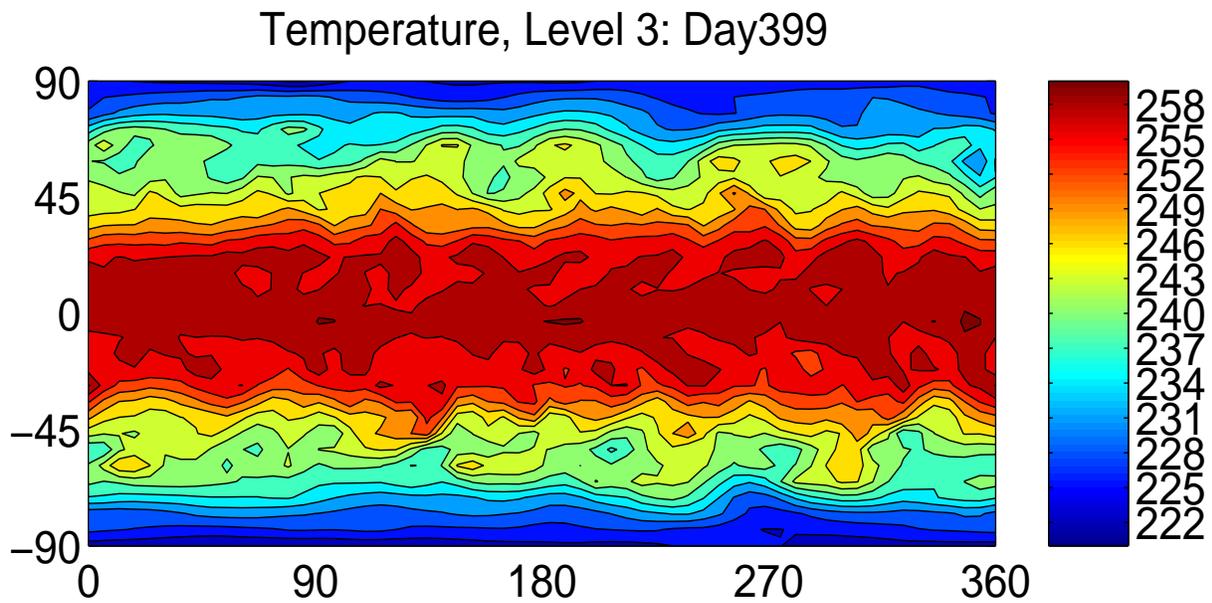
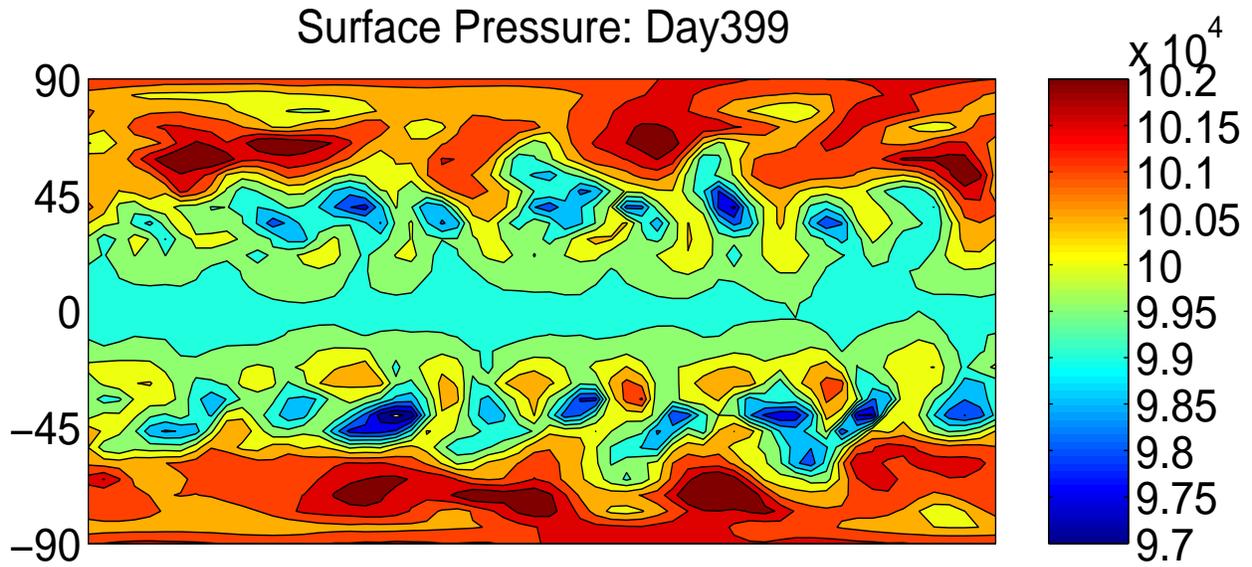


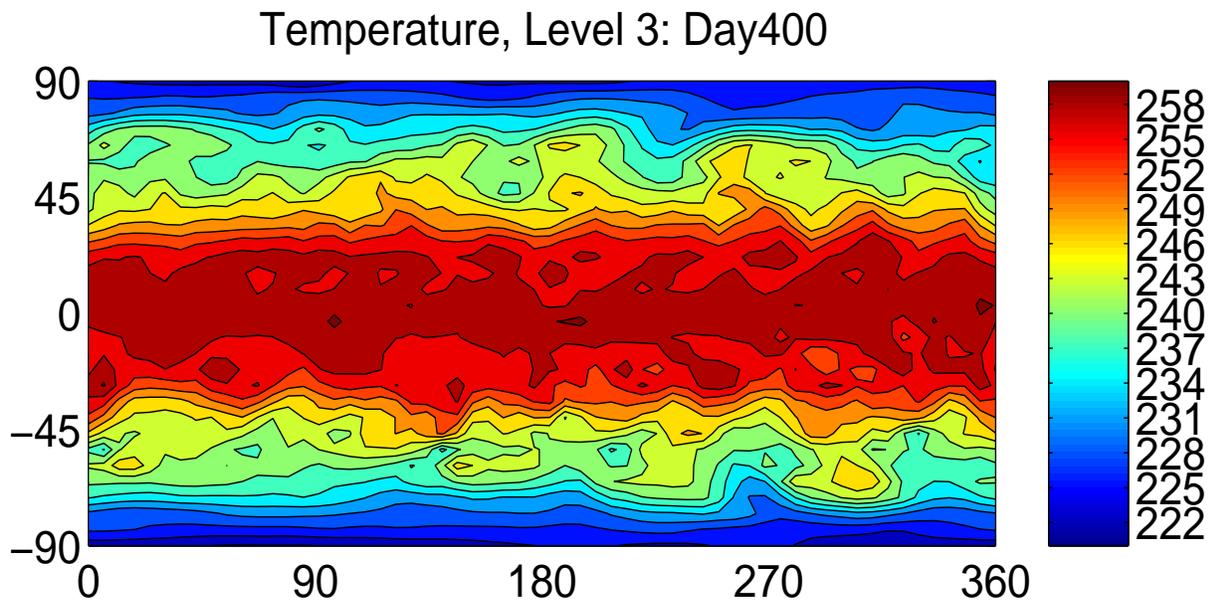
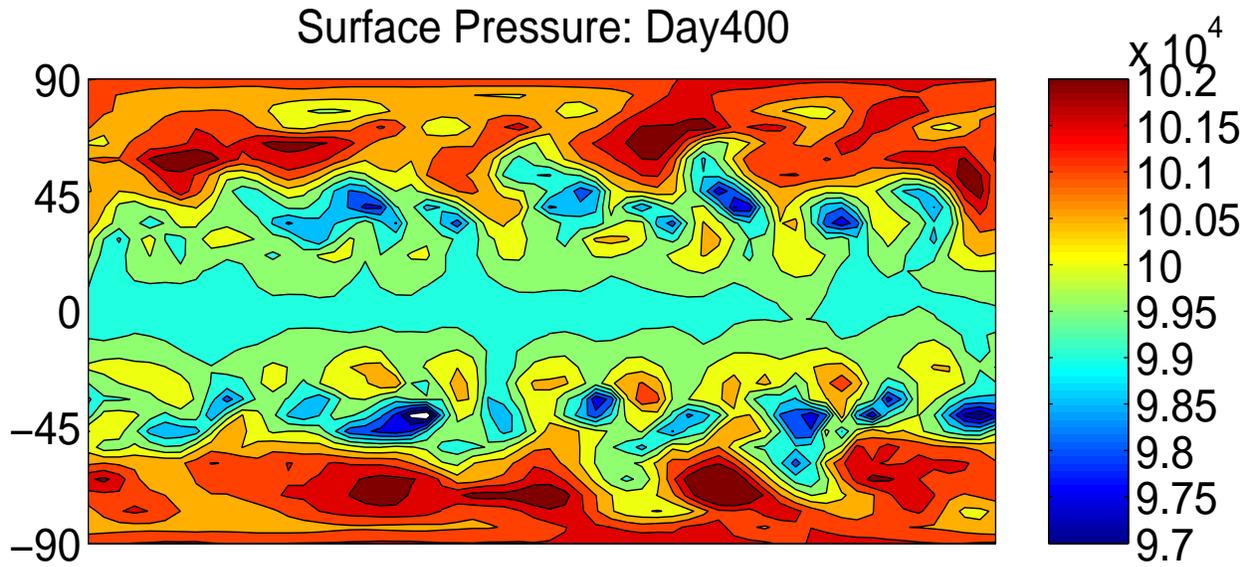
Surface Pressure: Day398



Temperature, Level 3: Day398



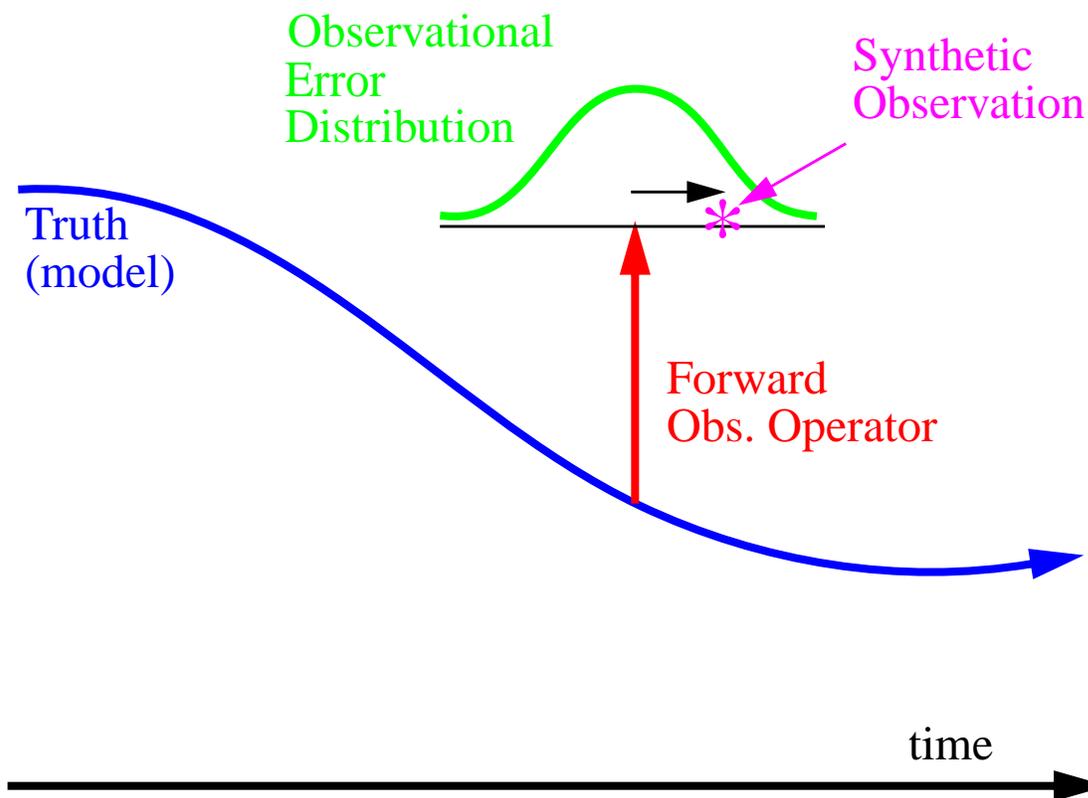




## Perfect model experiments

‘Truth’ is generated by integrating model

B-grid, integrated for 100 years from state of rest before starting  
(Multi-year spin-up for upper level temperatures)



‘Synthetic’ obs. by applying ‘forward observation’ operator to truth  
(Here, this is just interpolating to a random horizontal location)

Instrument error simulated by adding random draw from a specified  
Gaussian distribution to the interpolated observation

All the assimilation algorithm ever sees is these simulated observations

Result of assimilation can be compared to ‘truth’

## Experimental Design Details

Base case assimilation starts from 'climatological' ensemble

Add tiny perturbations to control integration (truth)  
Integrate this ensemble for several years

Ensemble size is 20 for ALL cases here

Each assimilation case is run for 400 days

Summary results are from last 200 days

No bias correction steps taken (no covariance inflation)

Single tuning parameter controls distance dependent correlation mask

Gives less weight to distant observations

This was tuned to give best RMS results in base case

Not changed for any other experiments

Note: Level 1 temperature in Held-Suarez configuration has very low frequency adjustment,

## Experimental Sets

1. Impact of spatial density of observations:  
150, 300, 450, 900, 1800, 3600, 7200, 14400, 28800 PS obs  
Every 24 hours  
PS observational error standard deviation 1.0 mb
2. Impact of frequency of observations  
1800 PS observations  
Every 24, 12, 6, 4, 3, 2, and 1 hours, 30, 15, and 5 minutes  
PS observational error standard deviation 1.0 mb
3. Information content of different observation types  
1800 observations of PS, or low-level T, or low-level U/V  
Every 24 hours  
PS observational error SD 2.0 and 1.0 mb  
T observational error SD 1.0 and 0.5 K  
U/V observational error SD 2.0 and 1.0 m/s, U, V errors independent
4. What happens if observations are confined to limited spatial domain  
450 PS obs, only in N. Hemisphere between 90 and 270 deg. longitude  
Every 24 hours  
PS observational error standard deviation 1.0 mb
5. Impact of increased vertical resolution  
1800 PS obs  
Every 24 hours  
PS observational error standard deviation 1.0 mb  
5 and 18 vertical levels
6. Impact of adding stochastic 'sub-grid scale' noise  
1800 PS obs, Every 24 hours  
PS observational error standard deviation 1.0  
Temperature time tendency noise standard deviation 0, 10%, 40%

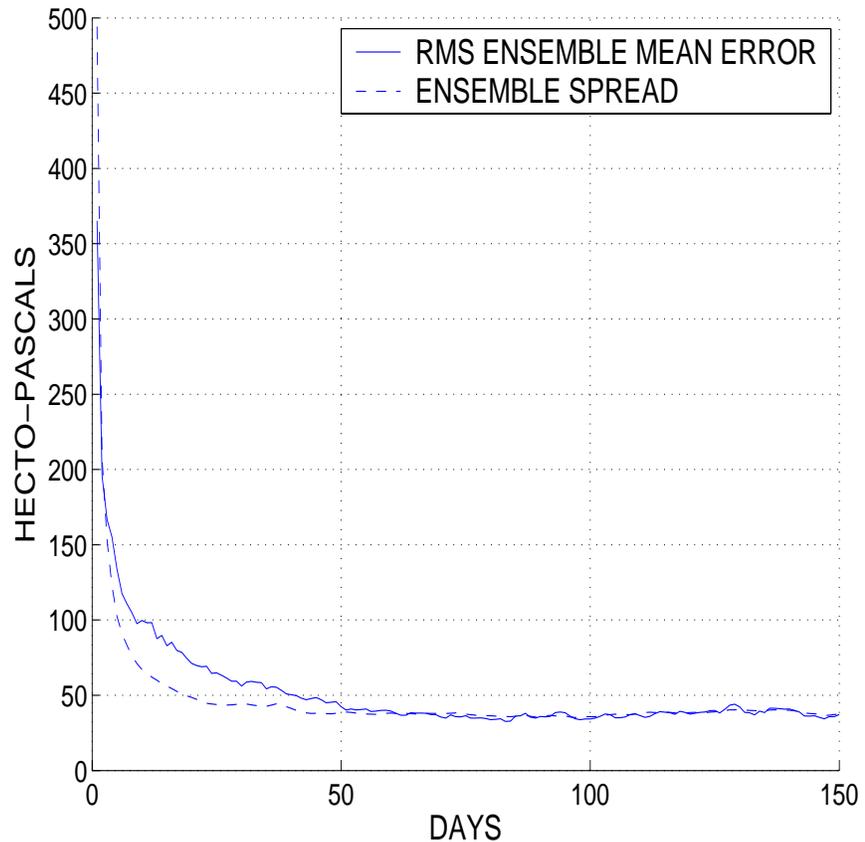
## Baseline Case: 1800 PS Obs every 24 hours

PS Error Reduces  
by about factor of  
10

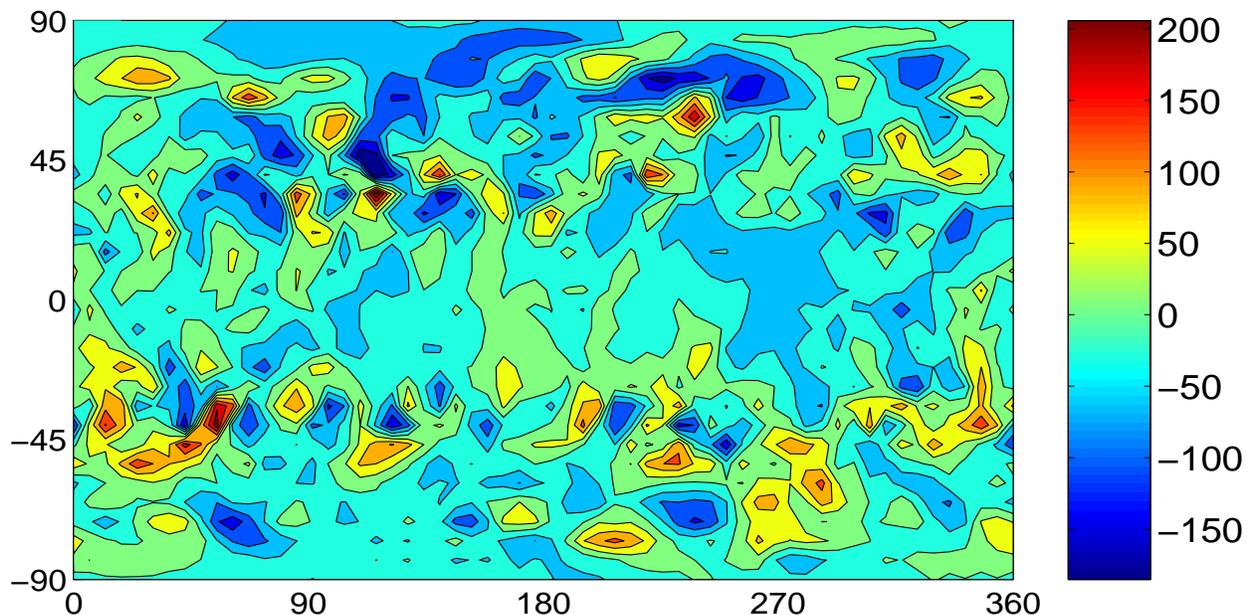
Asymptotes after  
about 50 days

Ensemble spread  
is approximately  
correlated with  
RMS error

Final Error is  
about 0.4 mb



Largest error in mid-latitudes, 'synoptic' scales after 400 days

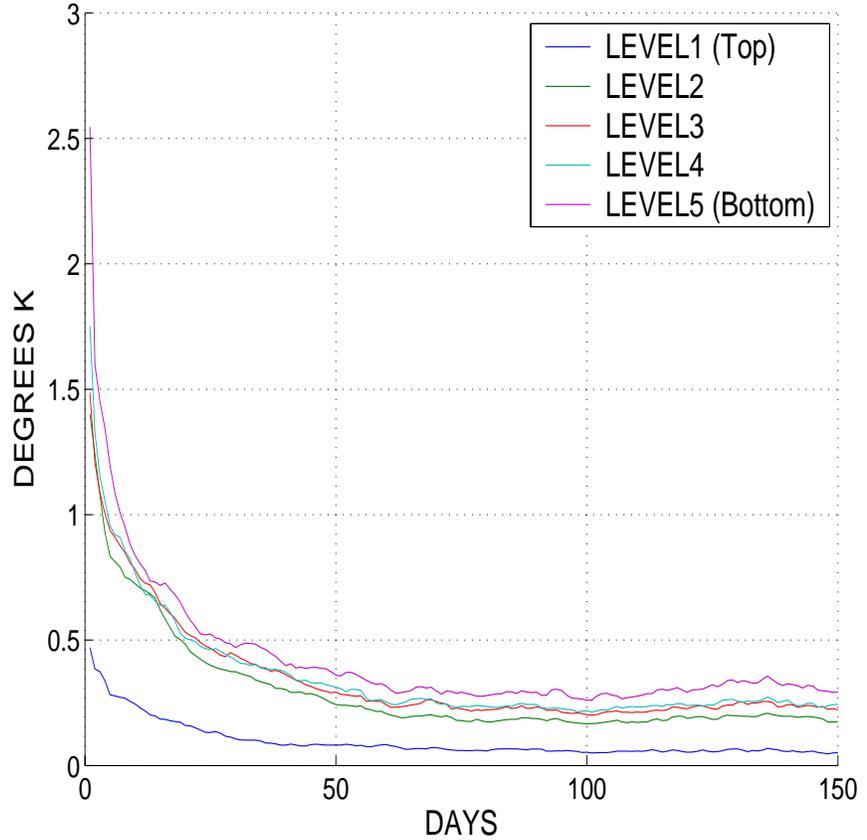


## Baseline Case: 1800 PS Obs every 24 hours

T Error also reduced  
by about factor of 10

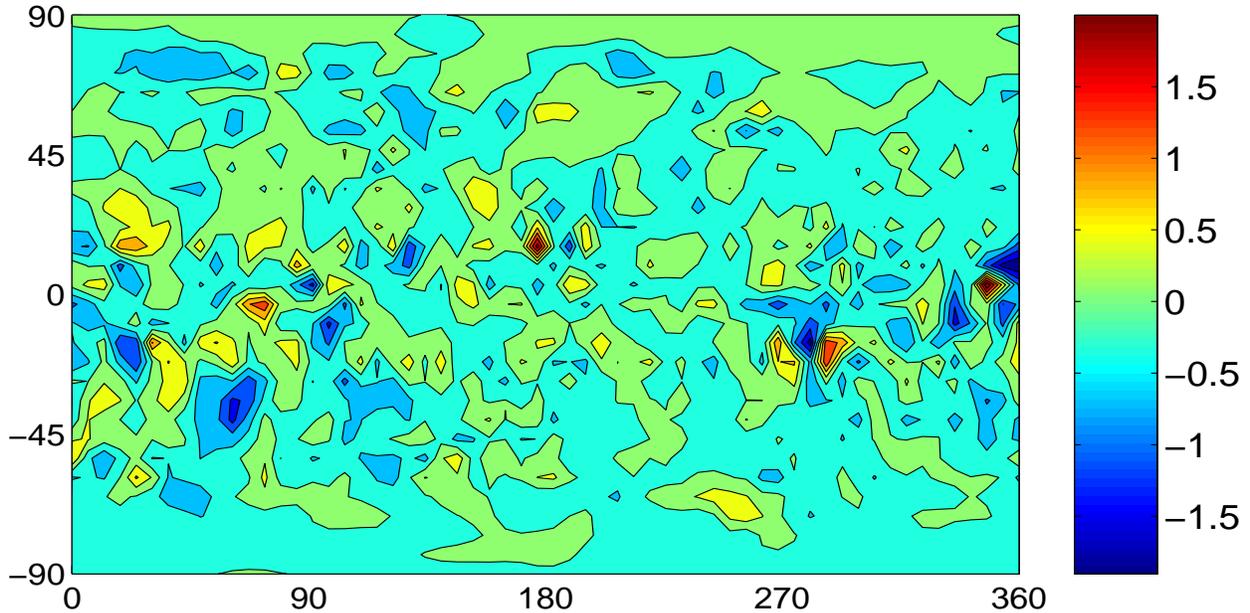
Asymptotes about  
day 70

Final error about  
0.25 K for interior  
levels



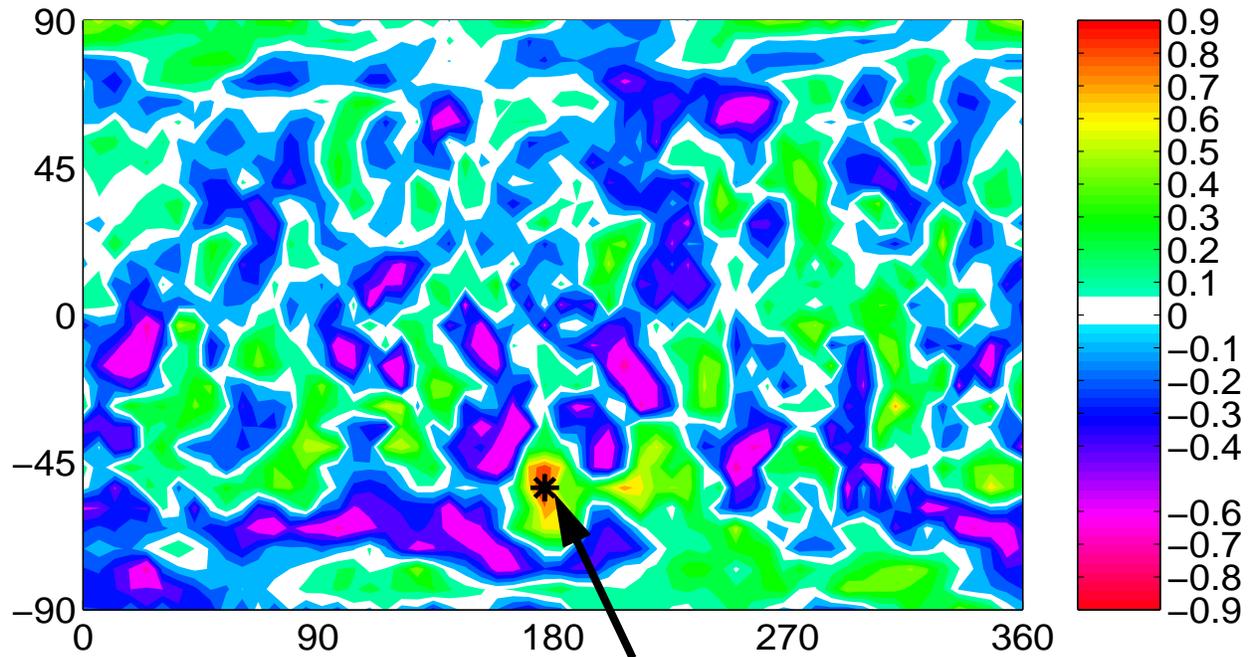
Largest T error in tropics for interior levels (level 3, day 400 shown)

Min = -1.9012 Max = 2.3893 RMS ERROR = 0.24305

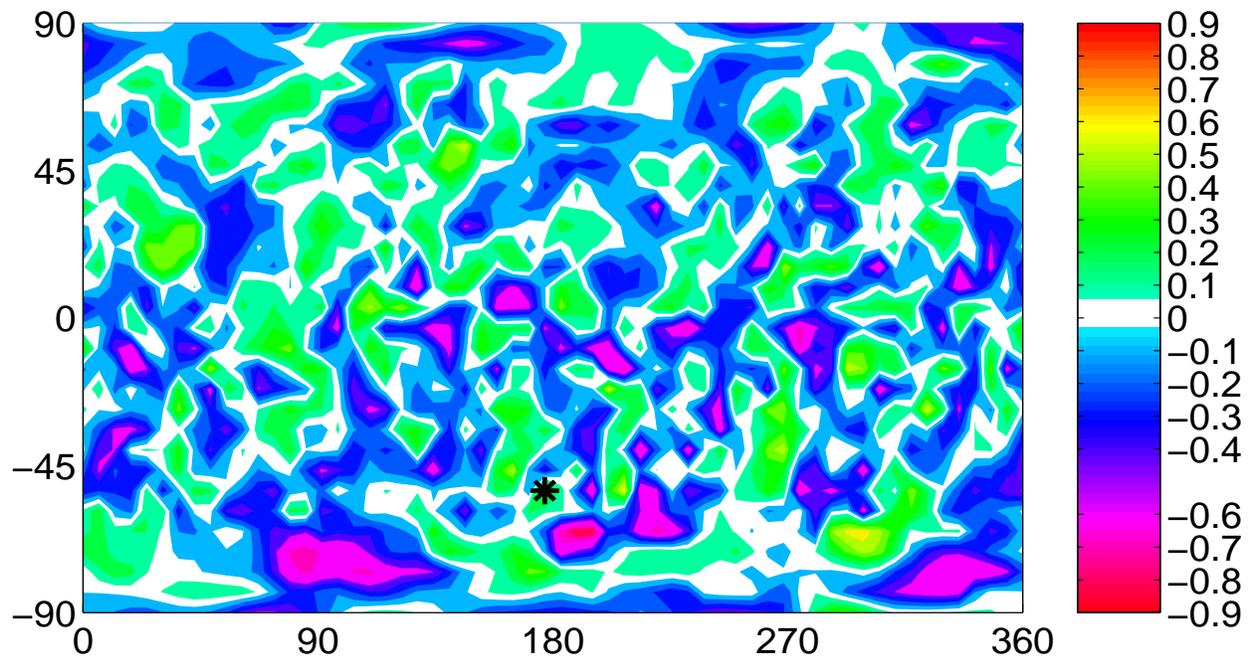


## Sample Correlation: Baseline Case

Sample correlations reflect how observations can impact state variables:



Correlation of PS with PS at (180, 50S): largest values local but noisy



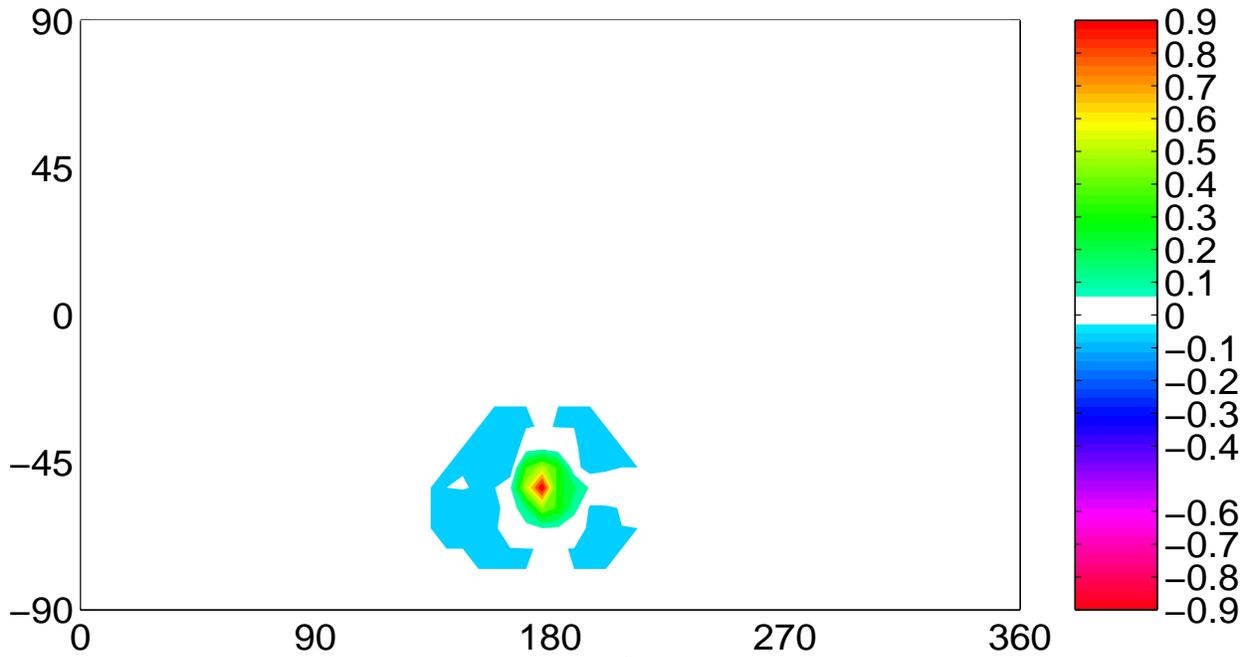
Correlation of T at level 3 with PS at (180, 50S);

Lots of noise, limited local signal

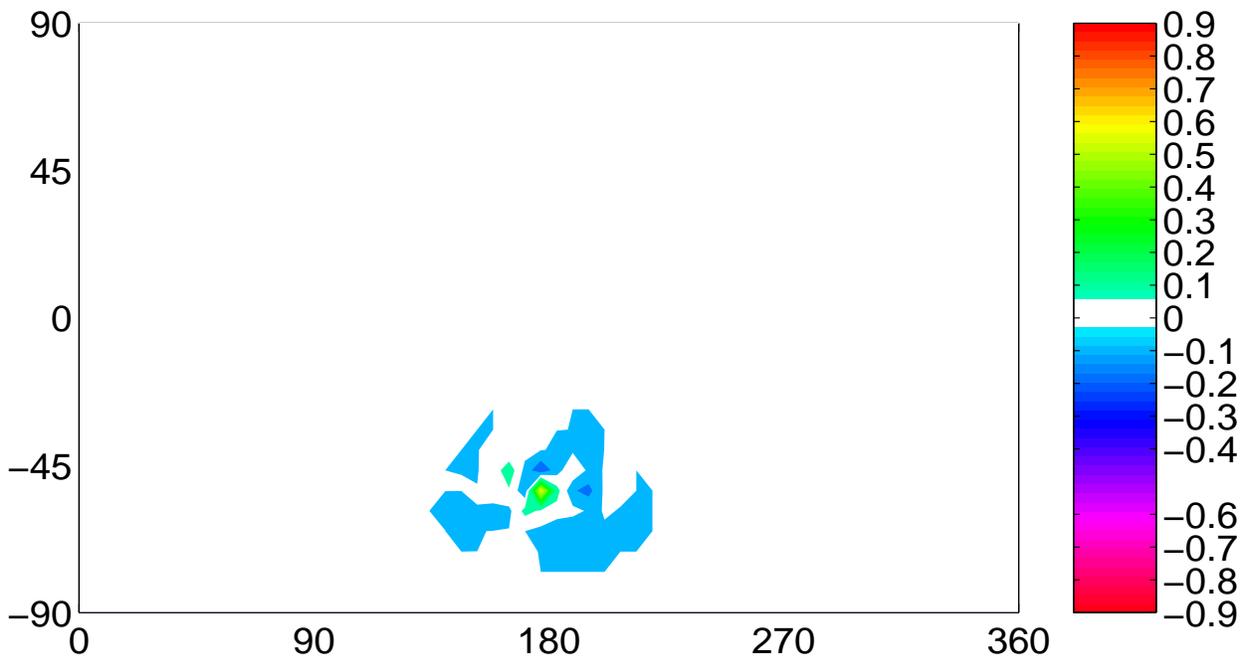
Filter must be able to extract limited signal from lots of noise

## Sample Correlation with Envelope: Baseline Case

Sample correlations reflect how observations can impact state variables:

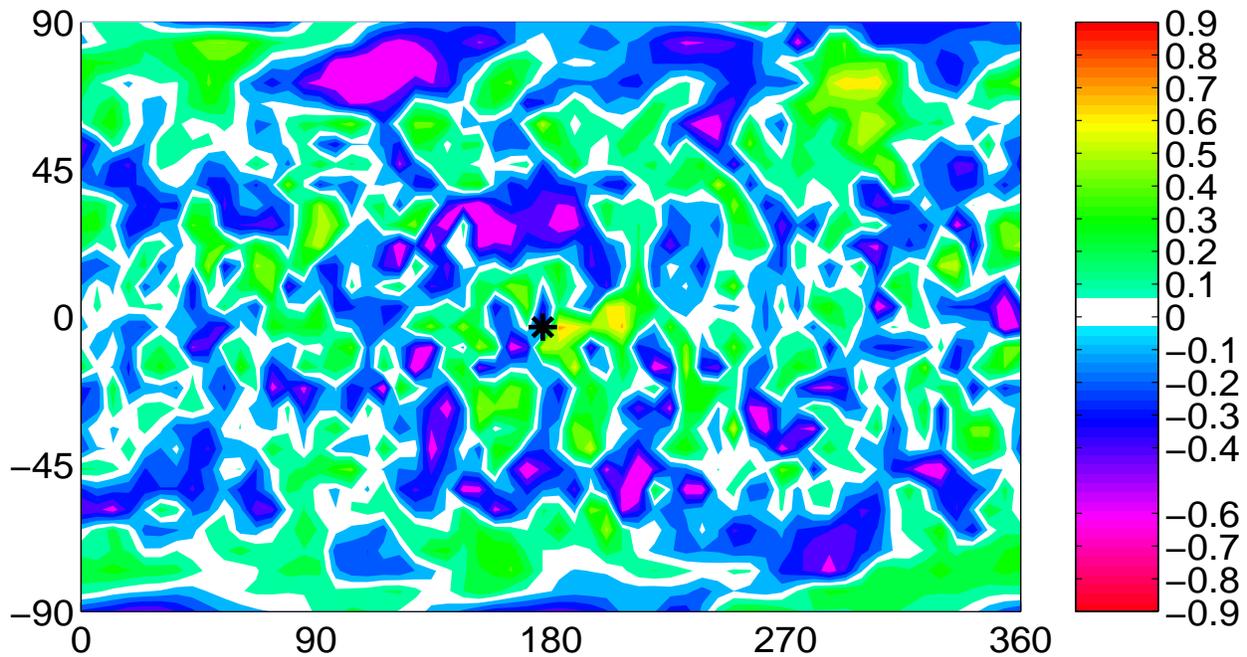
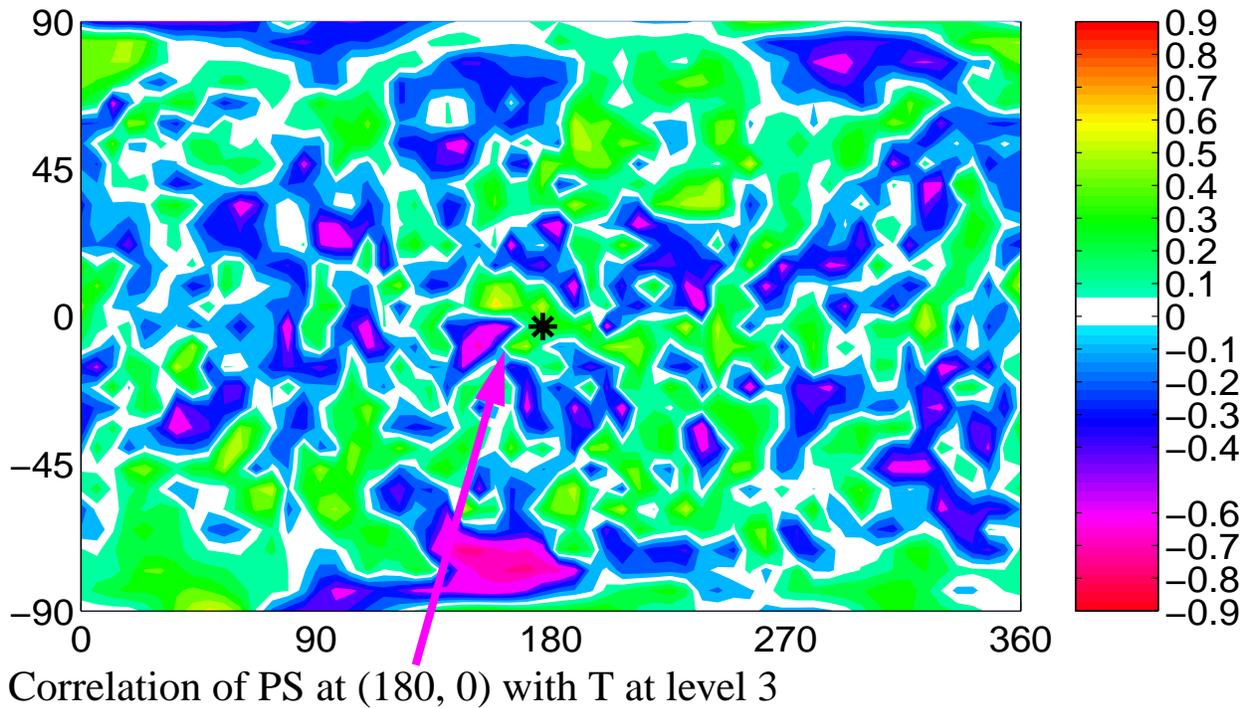


Correlation of PS with PS at (180, 50S):



Correlation of T at level 3 with PS at (180, 50S);

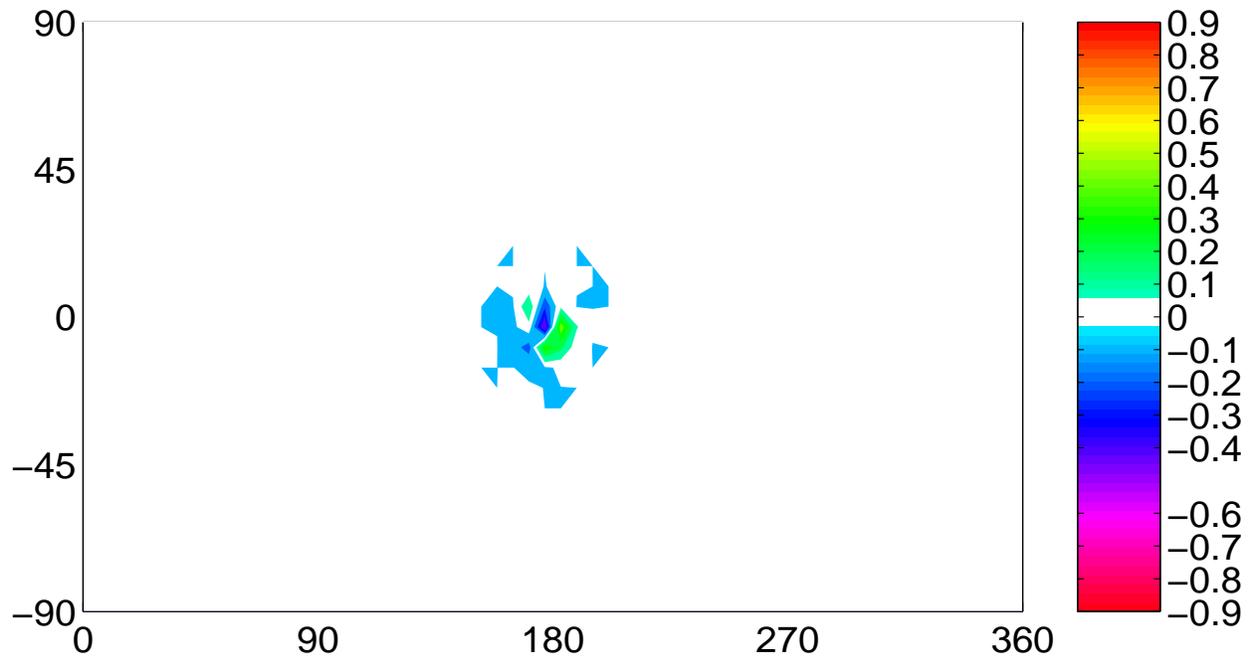
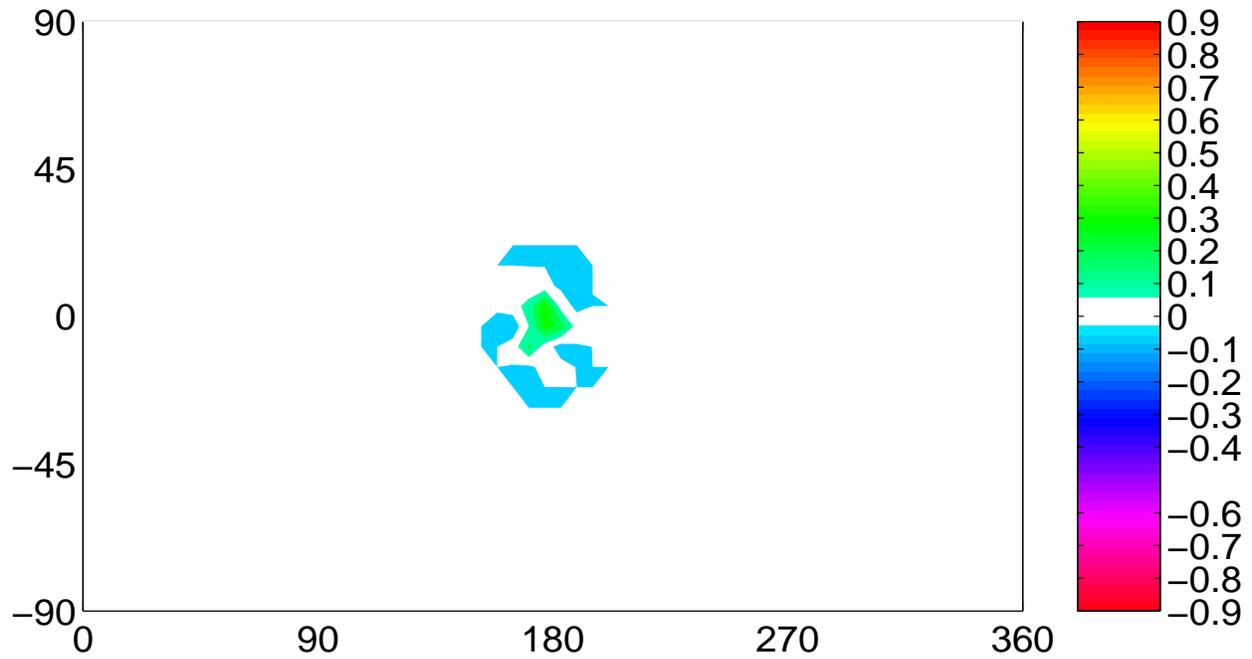
## Sample correlations vary significantly in time and space



Same field, but 10 days later; Local structure is somewhat similar  
 Noise at a distance has moved around randomly

Must take actions to avoid impact from remote noise

## Sample correlations vary significantly in time and space



### Impacts of spatial density of PS obs

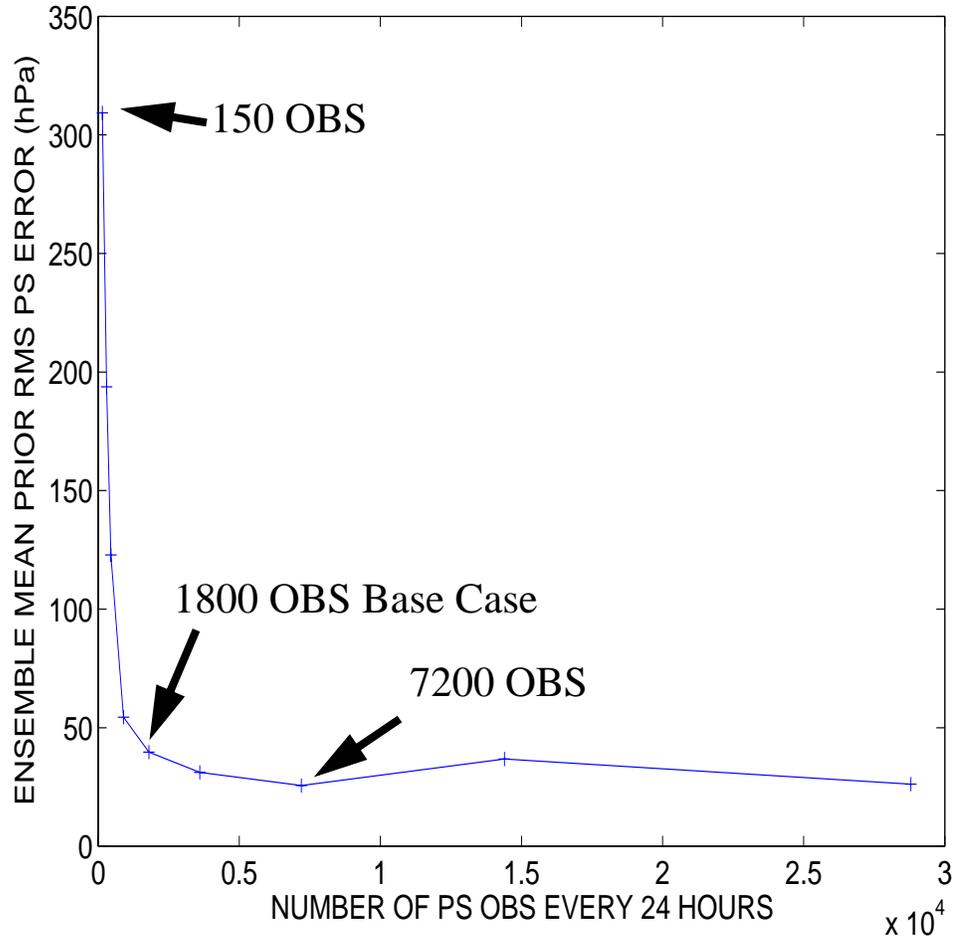
150, 300, 450, 900, 1800, 3600, 7200, 14400 and 28,800 every 24 hours

PS error  
reduces to  
about 0.3 mb

150 obs  
reduces clima-  
tological error  
by less than  
half

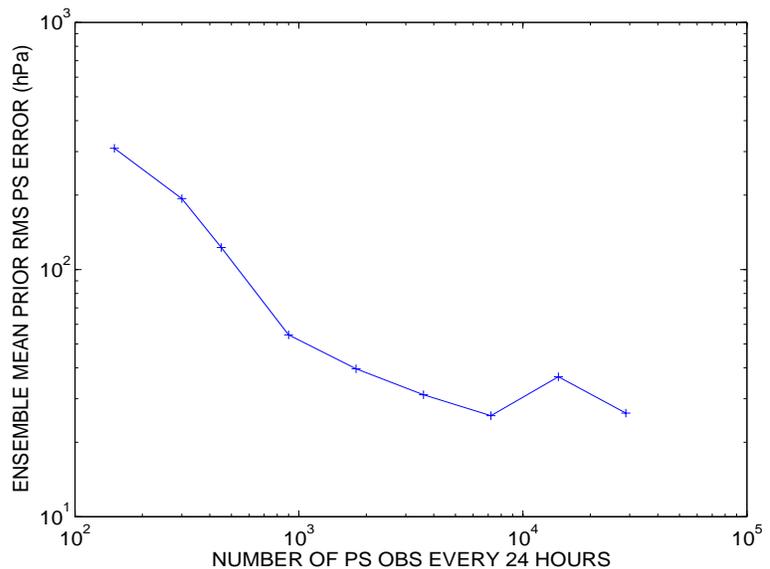
More than 7200  
obs appears  
superfluous

Why is 14,400  
worse? No  
clue.

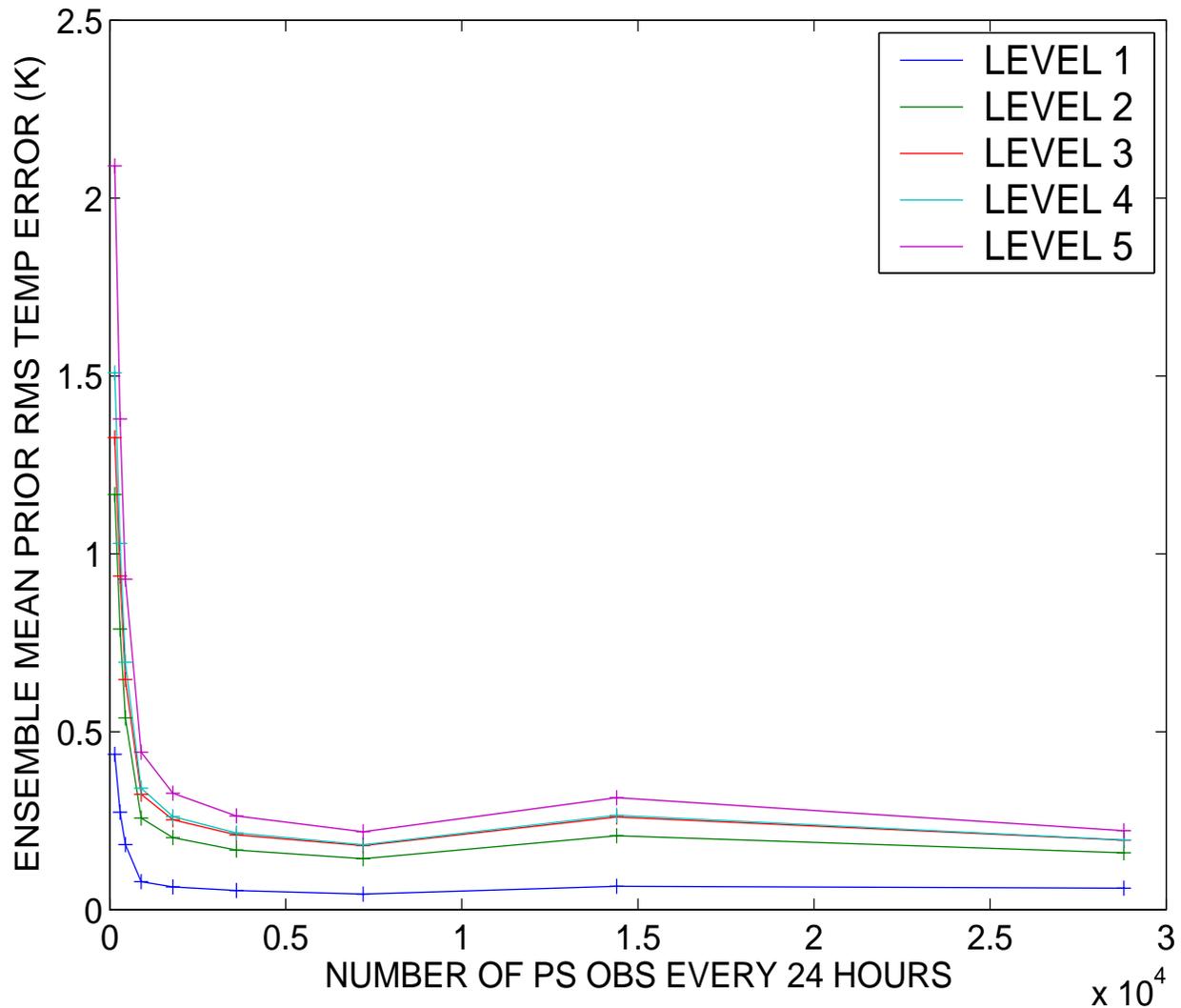


Plotting log /log of RMS  
shows approx. linear  
decrease from 150 to  
7200 obs

Behavior for very large  
numbers of obs clearly  
different



Impacts of spatial density of PS obs on Temperature RMS  
 150, 300, 450, 900, 1800, 3600, 7200, 14400 and 28,800 every 24 hours



Behavior for Temperature (and U, V not shown) similar to that for PS  
 Best results for 7200 PS observations  
 Interior level mean T RMS of about 0.25 K for best case

## Impacts of frequency of PS obs

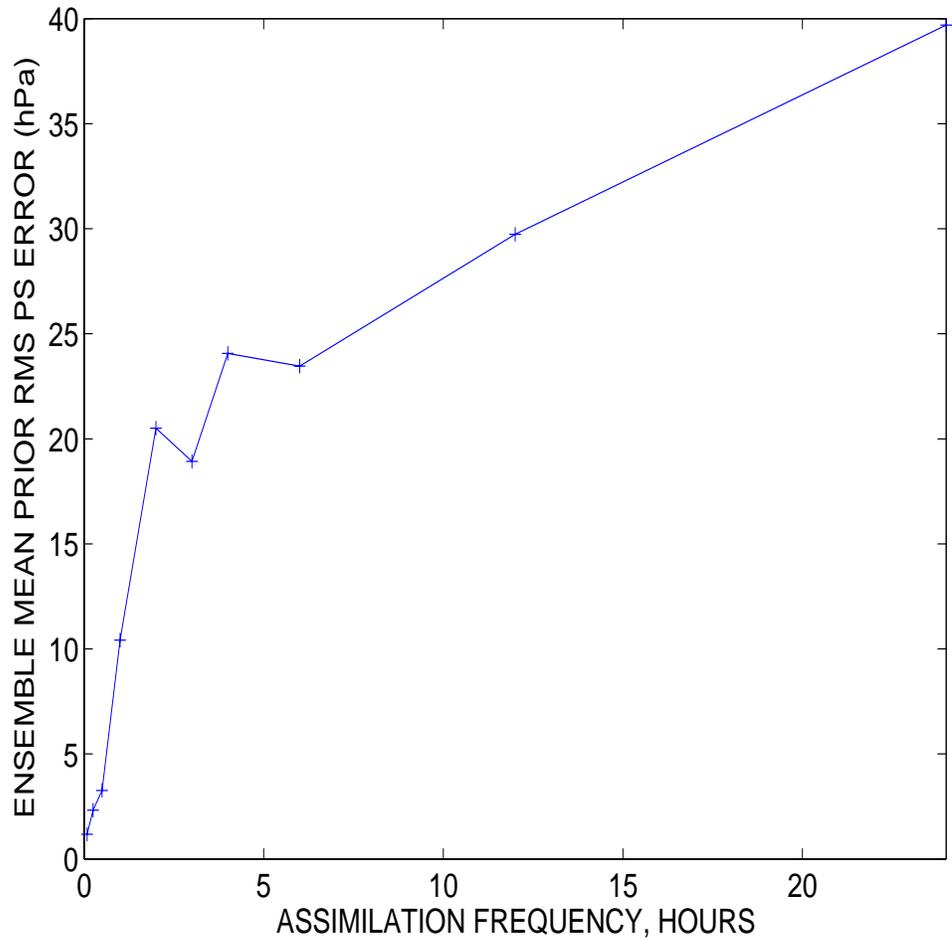
24, 12, 6, 4, 3, 2, 1 hours; 30, 15, 5 minutes; 1800 obs.

Steady RMS decrease as frequency increases

Much smaller RMS than for high density low frequency obs

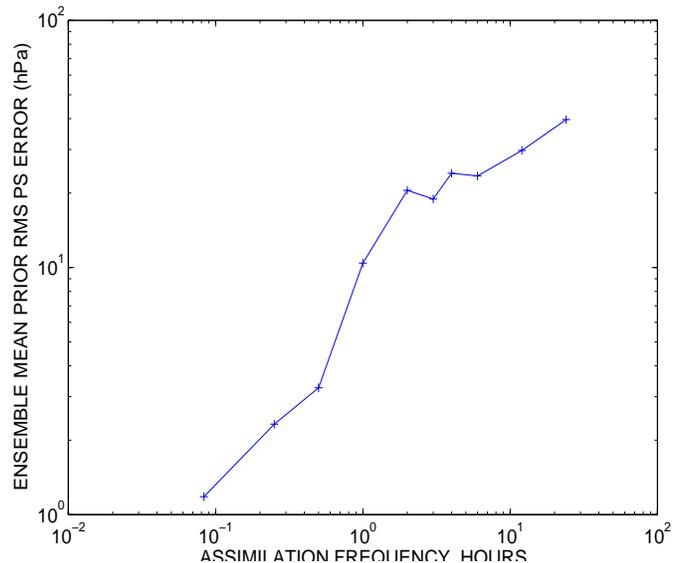
RMS < 0.02mb for 5 minutes

Strange behavior between 1 and 6 hour frequency



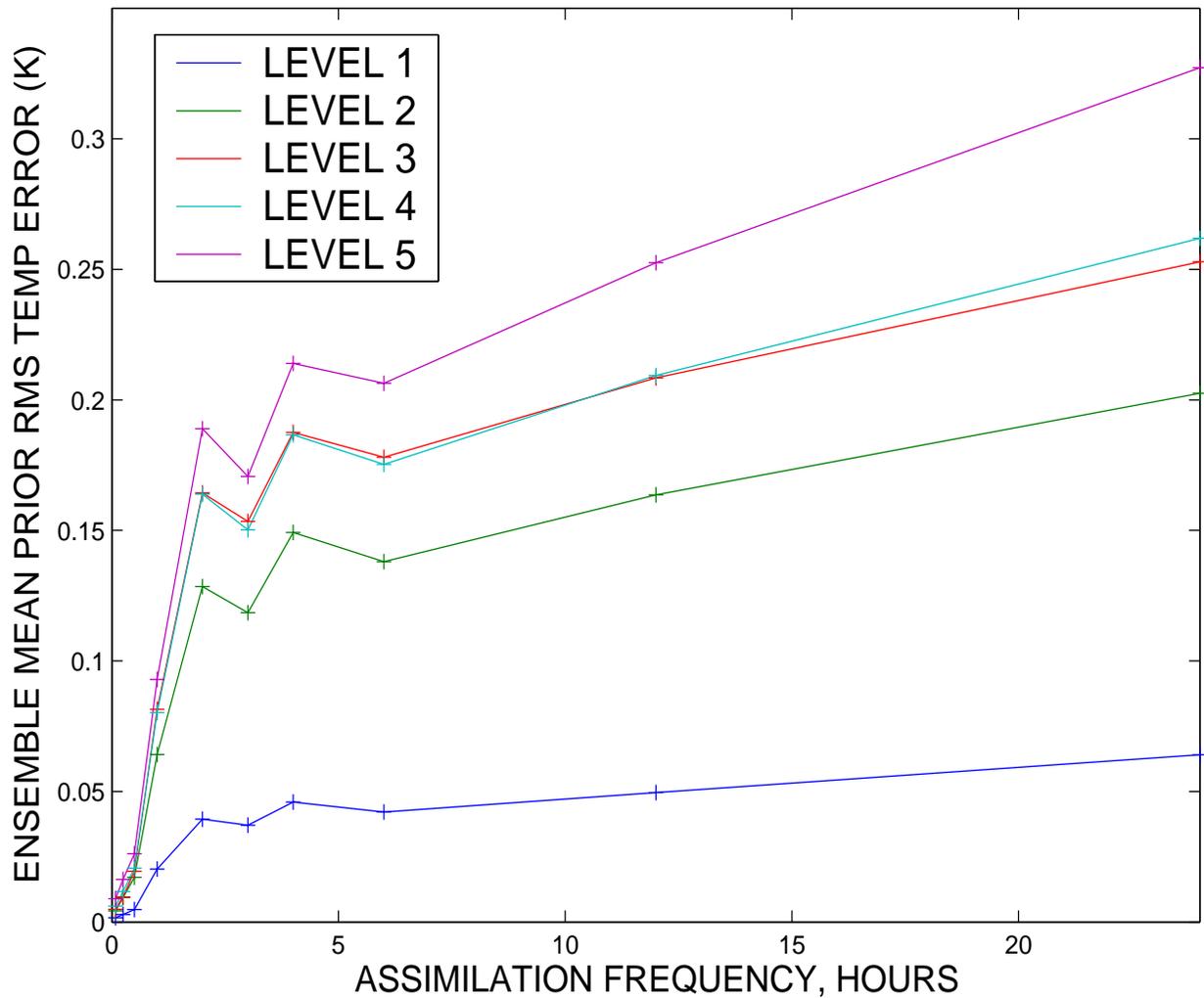
Plotting log /log of RMS shows approx. linear increase with a bump

What's going on in the middle?



## Impacts of frequency of PS obs

24, 12, 6, 4, 3, 2, 1 hours; 30, 15, 5 minutes; 1800 obs.



Temperature (and U and V, not shown) similar to PS  
 Consistent decrease in RMS with increased obs frequency

**Errors at 5 minute frequency less than 0.01 K !!!**

How low can you go?

## What's going on at moderate obs frequencies?

Equilibrated model has very low gravity wave amplitude  
 When perturbed, 'off-attractor' gravity waves can result  
 Noise in observations can project off attractor

Ensemble members pulled in same direction; get phased gravity waves

Gravity wave period varies: approximately 4 hours  
 Gravity waves heavily damped; quickly reduced in amplitude

Low frequency ( $> 12$  hours): gravity waves damped before next obs time

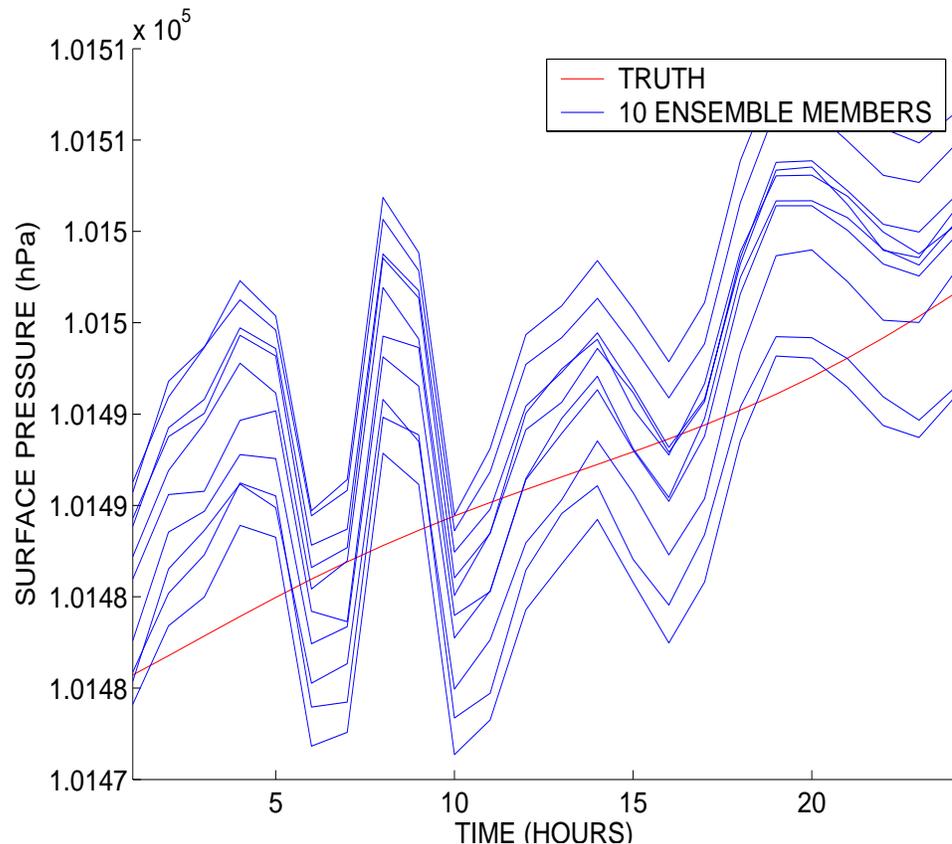
High frequency ( $< 1$  hour): enough obs per period to control amplitude

Moderate frequency (~ 4 hours): get phased gravity waves in ensemble;  
large bias; increased assimilation error

Time series of  
 10 out of 20  
 ensemble  
 members at  
 mid-latitude PS  
 point.

Forecast initi-  
 ated from end  
 of 1800 PS obs  
 every 4 hours

(Extreme  
 example)



## Why does increasing frequency do more than increasing density?

>>1. Temporal has more 'independent' correlation estimates

Can better eliminate sampling noise

>>2. Temporal sees observations at more 'phases' of wavelike structures

>>3. Large ensemble size could help to distinguish this by reducing sampling noise

These are yet to be done

>>4. Historically, high frequency obs were hard to acquire

Modern technology changes this

Exploring use of high frequency obs is planned

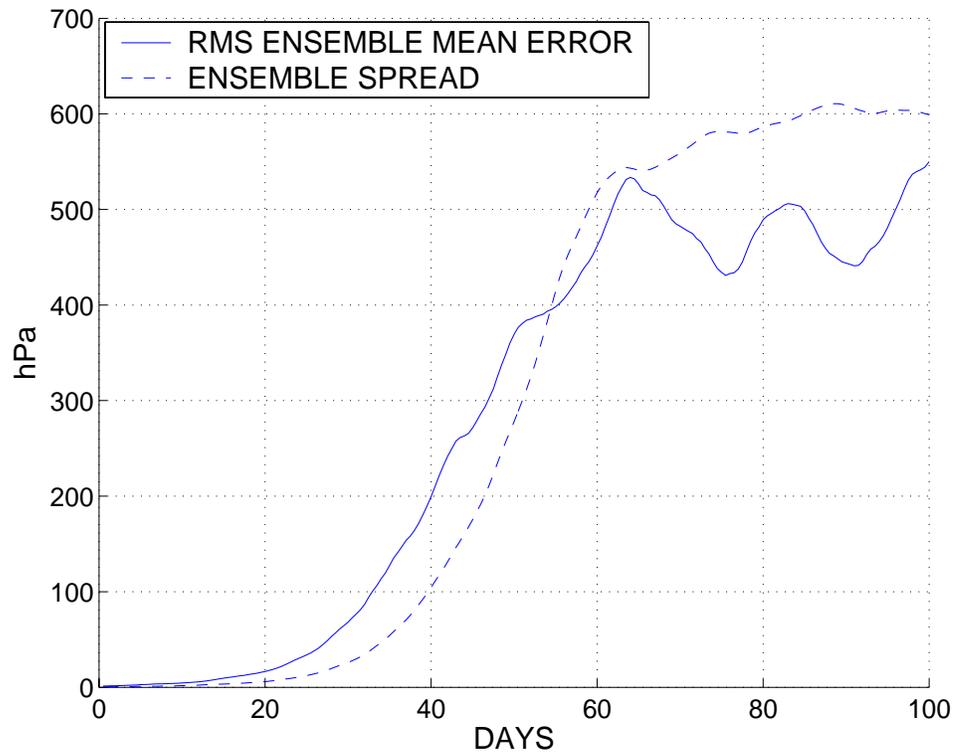
## Need to demonstrate model has error growth

Free integration (forecast) at end of 1800 PS obs every 5 minutes

Error saturates  
at climatologi-  
cal values after  
about 60 days

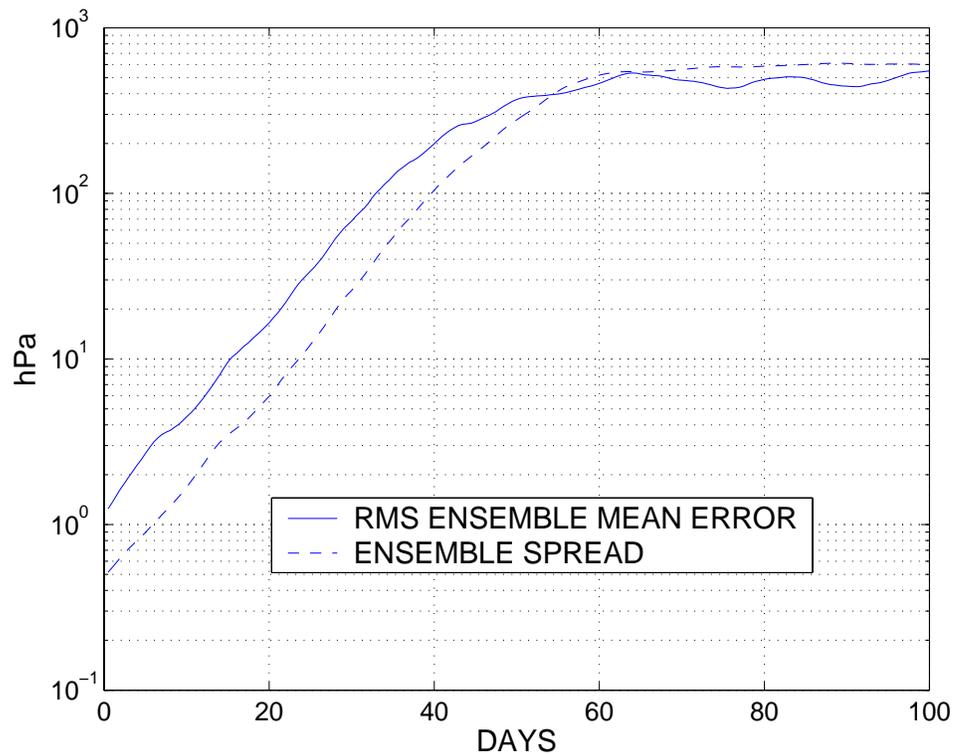
Error doubling  
time about 7  
days

Considerably  
slower than real  
atmosphere



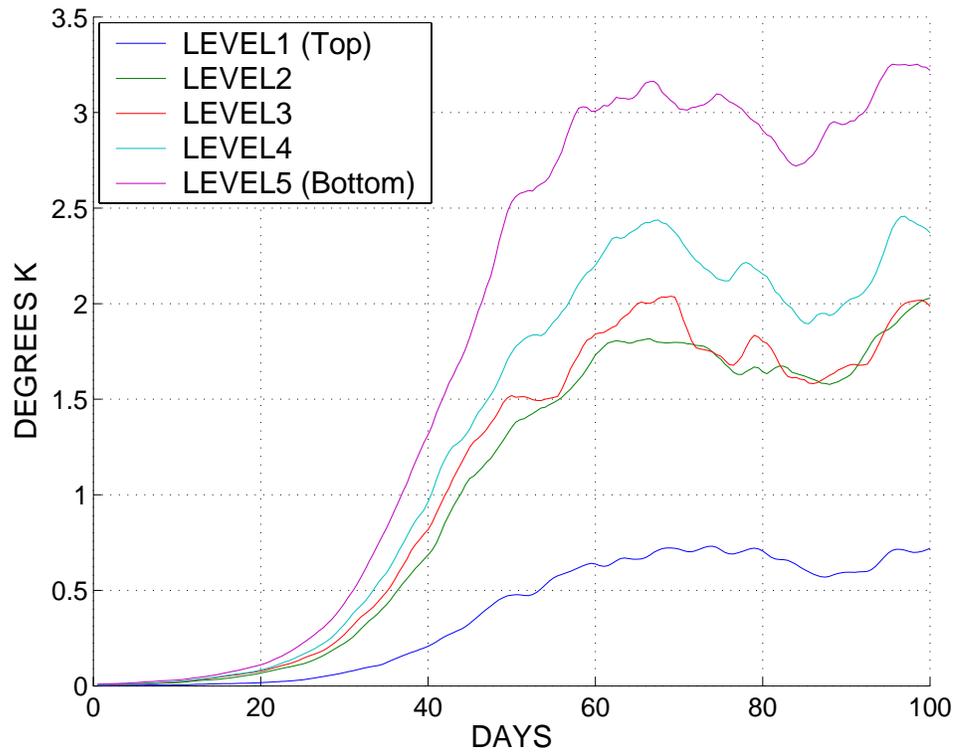
Error growth  
very nearly lin-  
ear in log plot

Growth is  
almost purely  
exponential to  
saturation

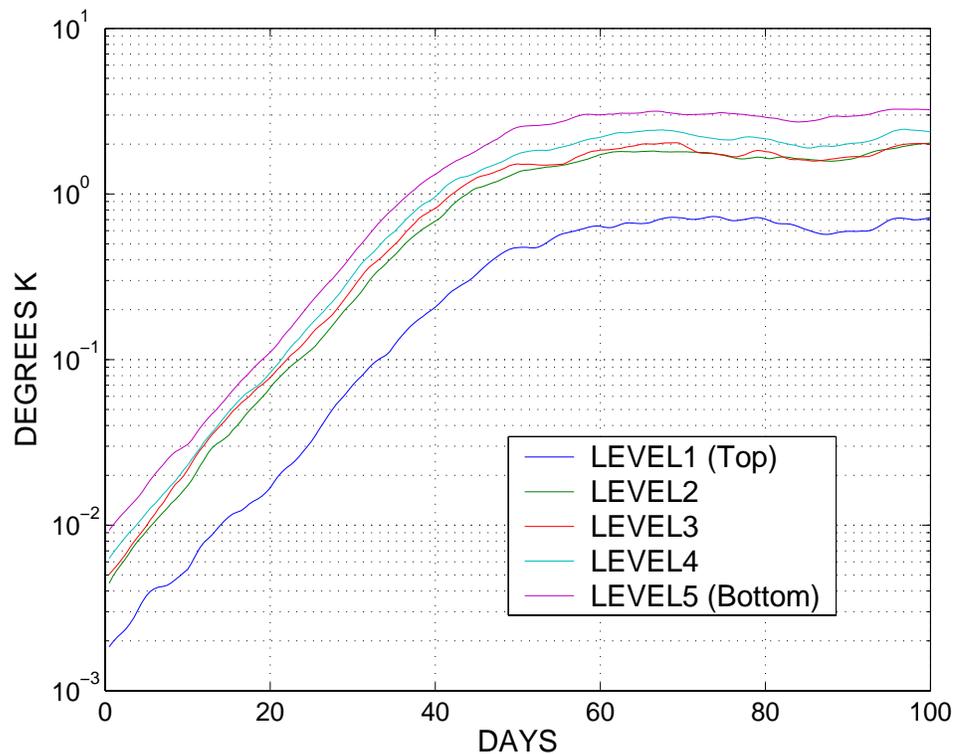


## Error growth of other fields similar to PS

T (and U and V not shown) also saturate at about 60 days



Growth is very nearly exponential throughout

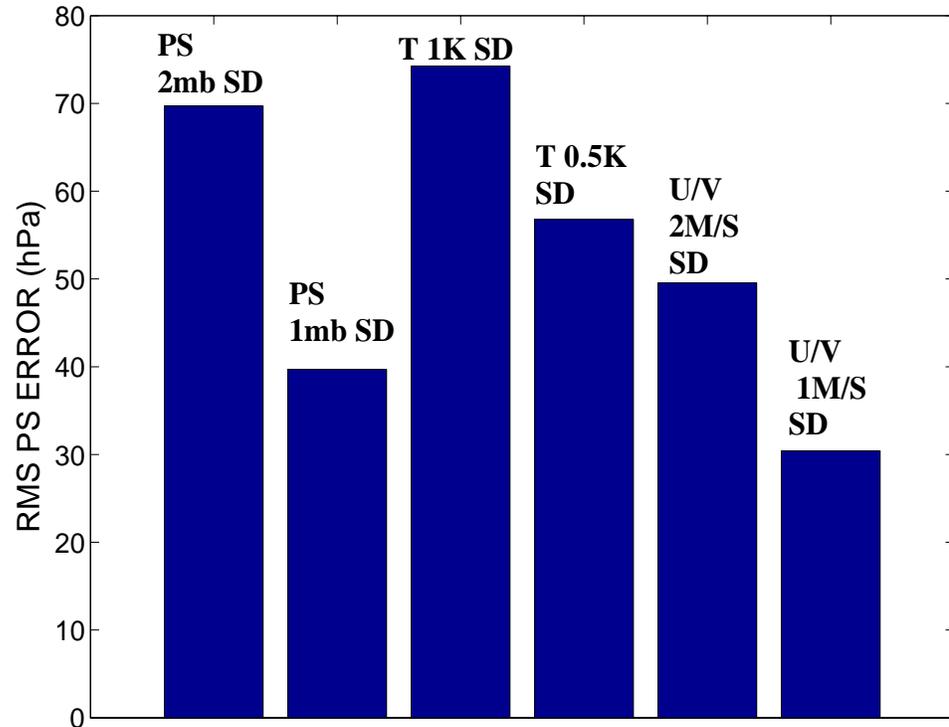


## Relative Information Content of Various Surface Obs

Compare PS with T and U/V obs from lowest level

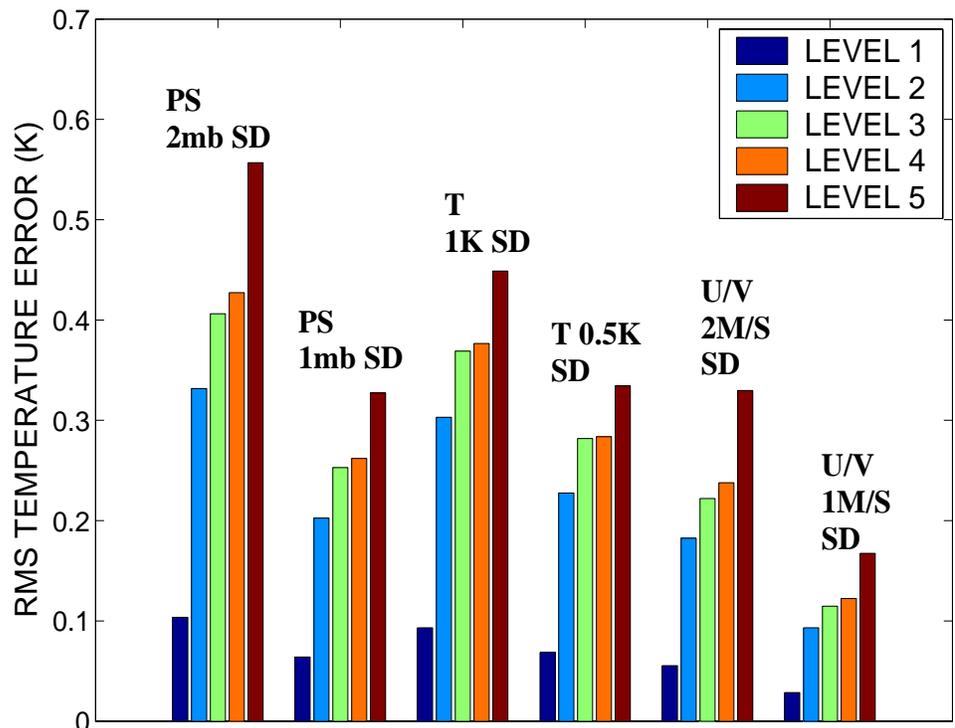
RMS error of PS prior assimilation when assimilating 1800 PS, T, or U and V wind components every 24 hours.

Two specified error SDs are checked for each.



RMS error of T for same cases.

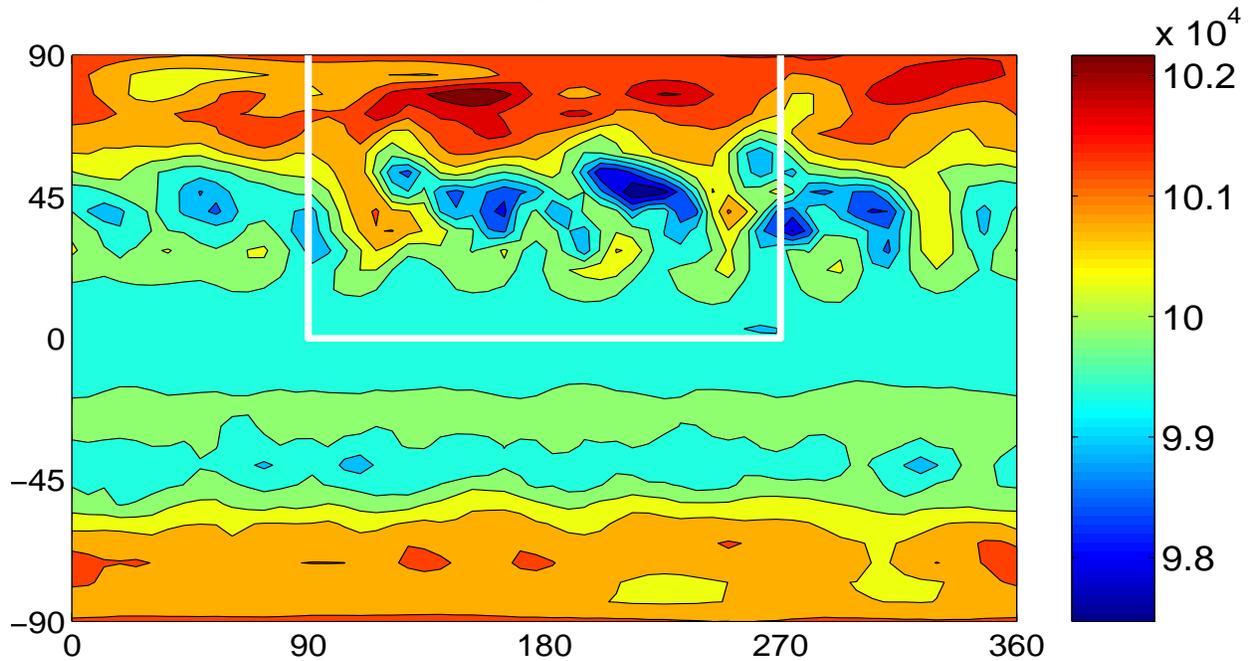
Very roughly, U/V obs with 2M/S SD have same information content as PS with 1mb SD or T with 0.5K SD.



## Assimilating PS over limited domains

450 PS obs every 24 hours over 1/4 of surface

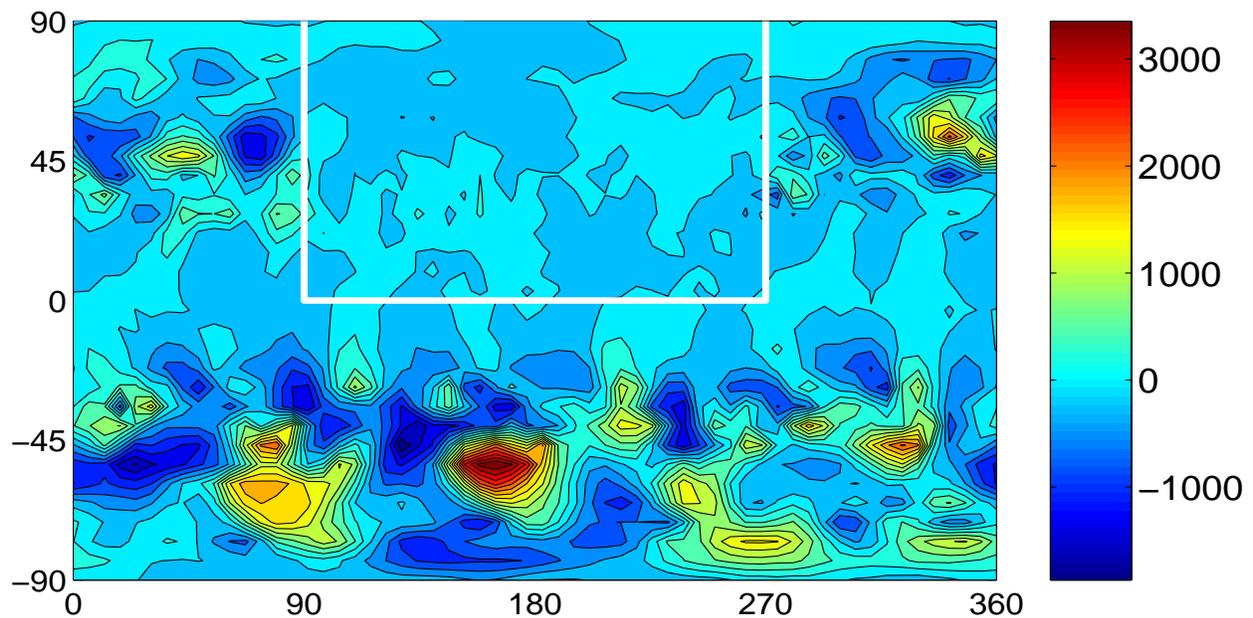
Ensemble mean prior assimilation for PS at 400 days  
Approaches zonal climatology with no obs information



RMS Error for PS at 400 days

Error in box about twice the value for 1800 global obs

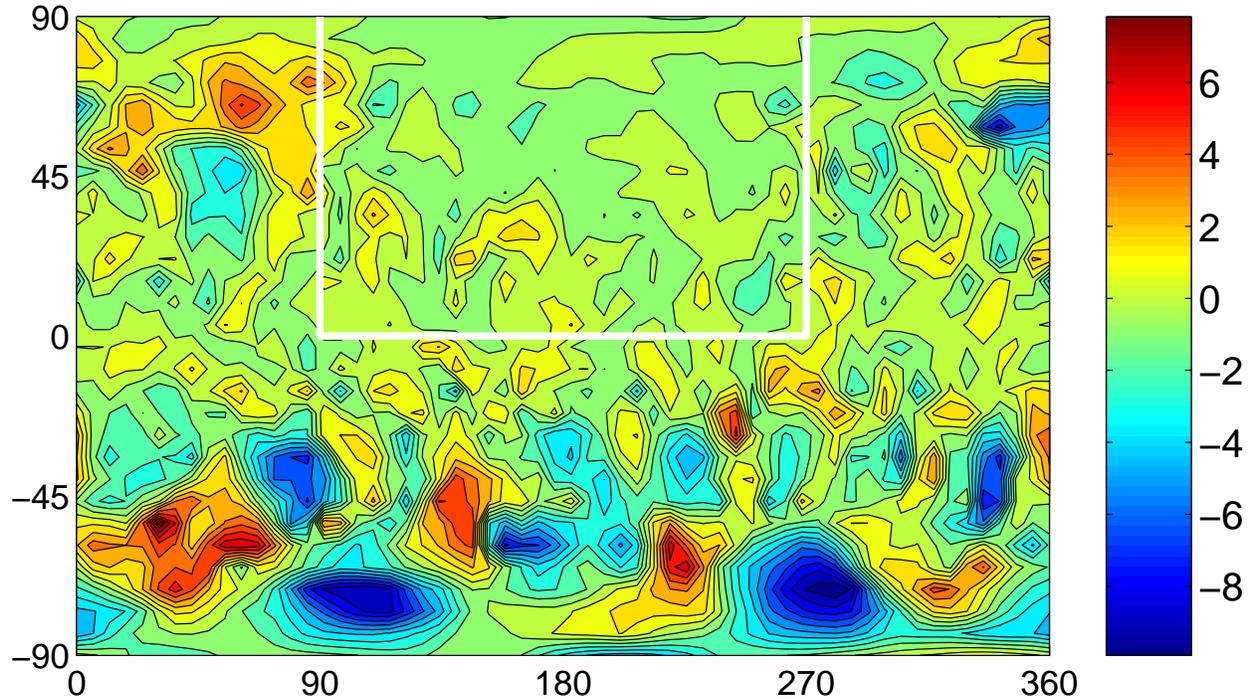
RMS ERROR = 320.2063 ERROR IN BOX = 84.584



## Assimilating PS over limited domain

RMS error for T at day 400;

RMS ERROR = 1.4118 ERROR IN BOX = 0.59219



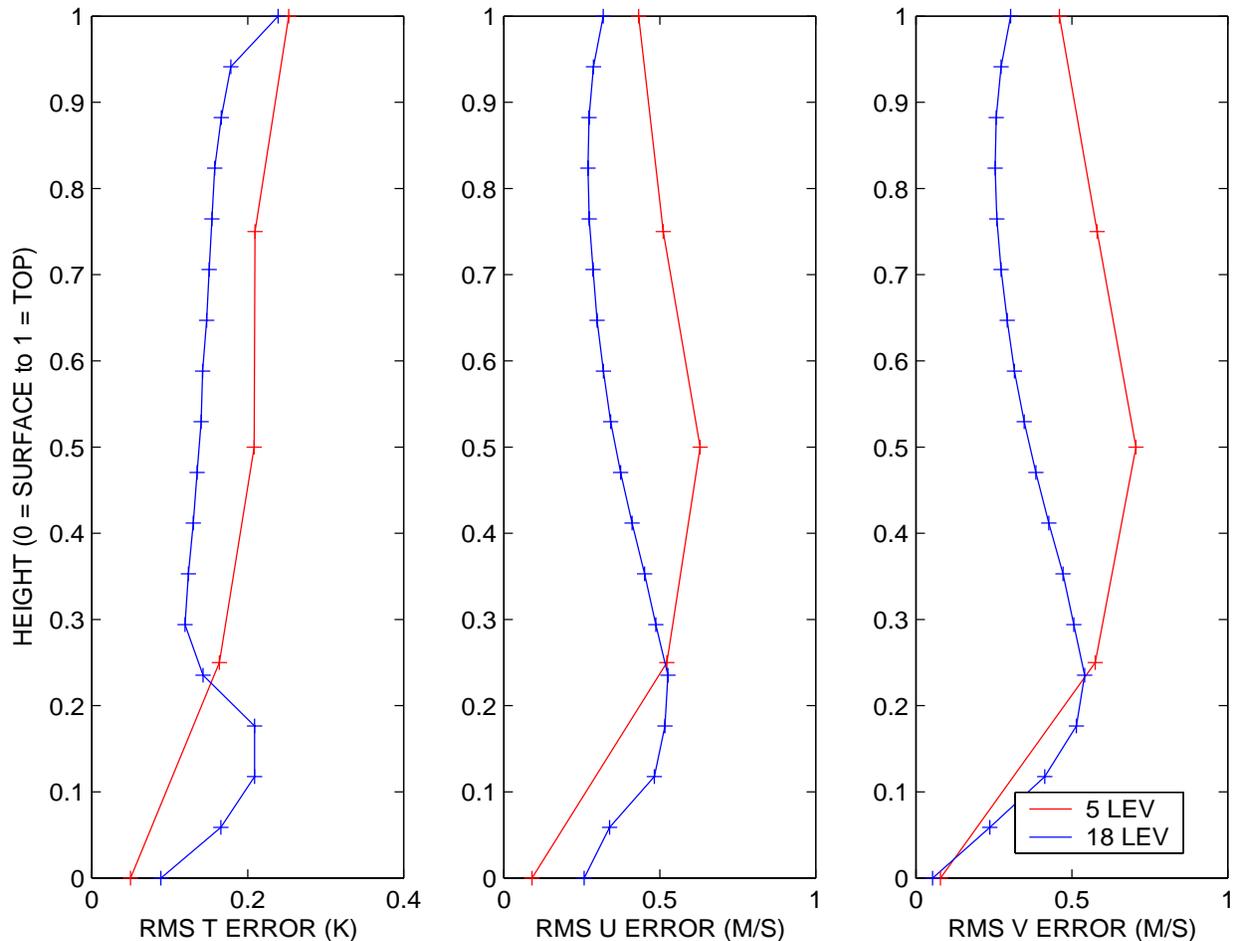
Error in box about twice that for 1800 global obs

Information is advected out of the box (to the east in mid-latitudes)

Method handles low information propagating in from upstream

Implications for regional and nested model filter data assimilation

## What happens with increased resolution?



### Comparison of 1800 PS obs for 5 and 18 level model

Tricky comparison, diffusion, etc. are identical

Error in upper levels of 18-level actually less

### Horizontal resolution, water vapor, and more comprehensive physics:

First results in NCAR CAM at 2 degree resolution appear consistent

Results by Whitaker and Hamill with PS obs in NCEP model are good

## Predictability and stochastic sub-grid scale parameterizations

Models don't resolve all spatial scales and processes

Normally parameterized (usually by column physics)

In prediction models, physics is usually deterministic

**In reality, best we can hope for is to know probability distribution for impact of unresolved processes**

Can simulate this in perfect model by adding random noise to model

Here, add noise factor to temperature tendency computation

At each gridpoint, let  $dT/dt = \text{MODEL} * (1 + N(0, R))$

$N(0, R)$  is random number with mean 0 and standard deviation  $R$

Independent noise at each point in current implementation

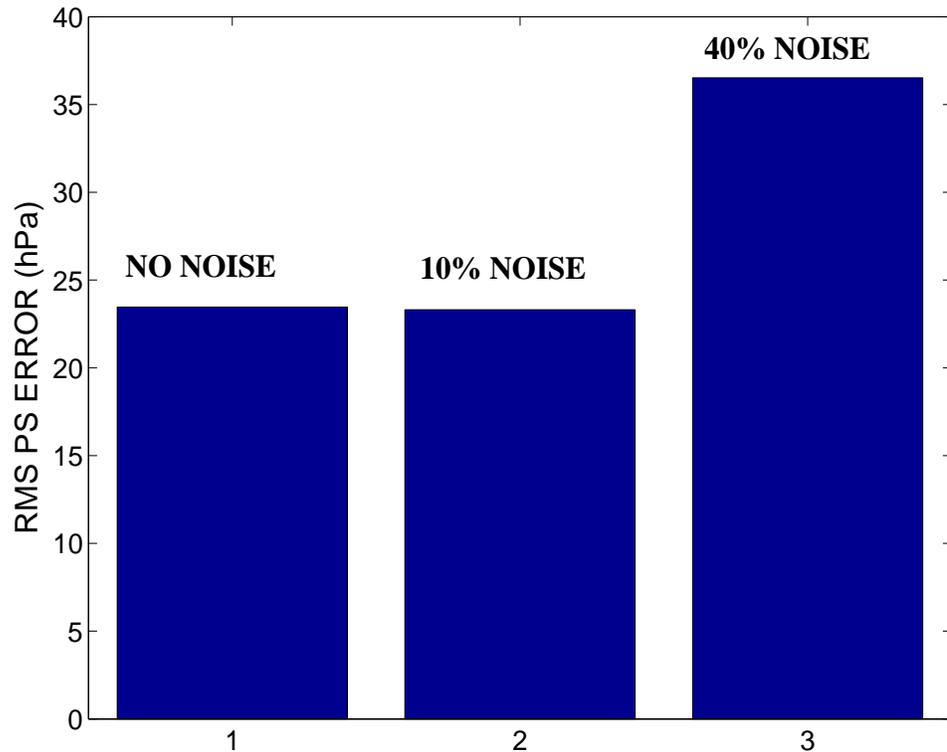
Ran cases with  $R = 0.1, 0.4$

1800 PS obs every 6 hours (moderate gravity wave amplitude)

## Impacts of sub-grid noise on Assimilation Error

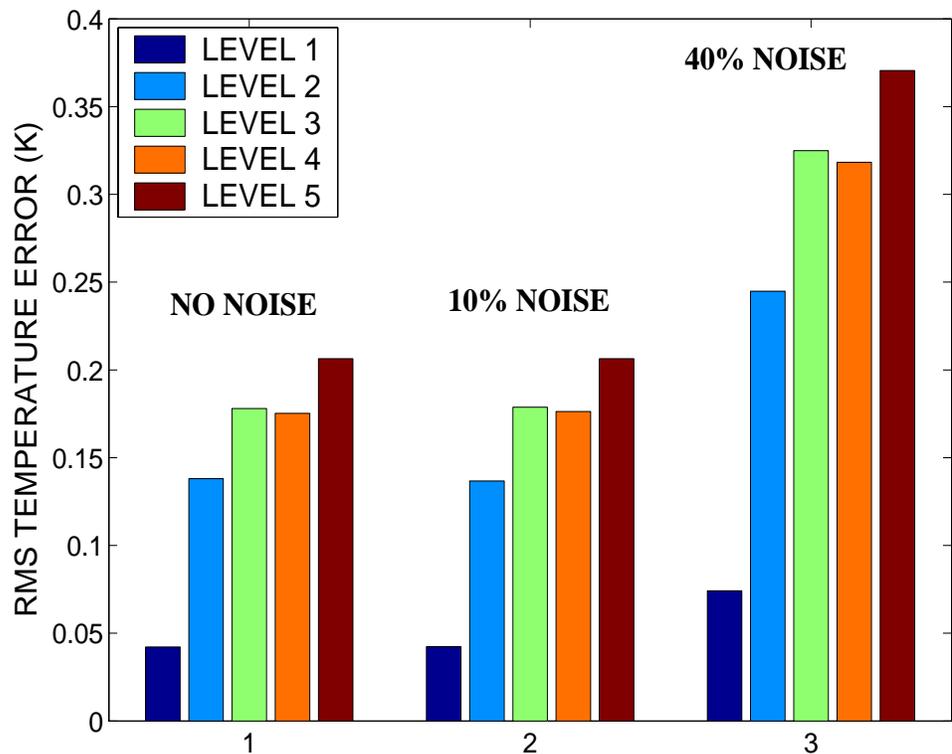
For 10%, error is slightly reduced

Adding noise in right proportions can solve gravity wave problem?



T results similar.

40% case has much increased error.

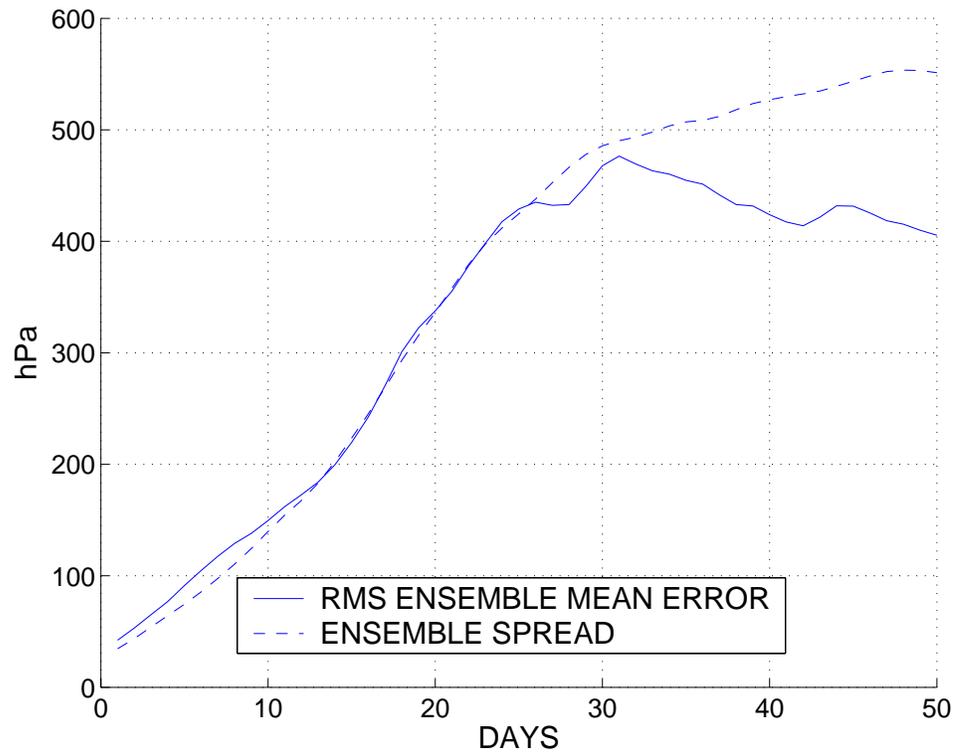


## Error growth and predictability with sub-grid scale noise

PS error growth is mostly linear at first.

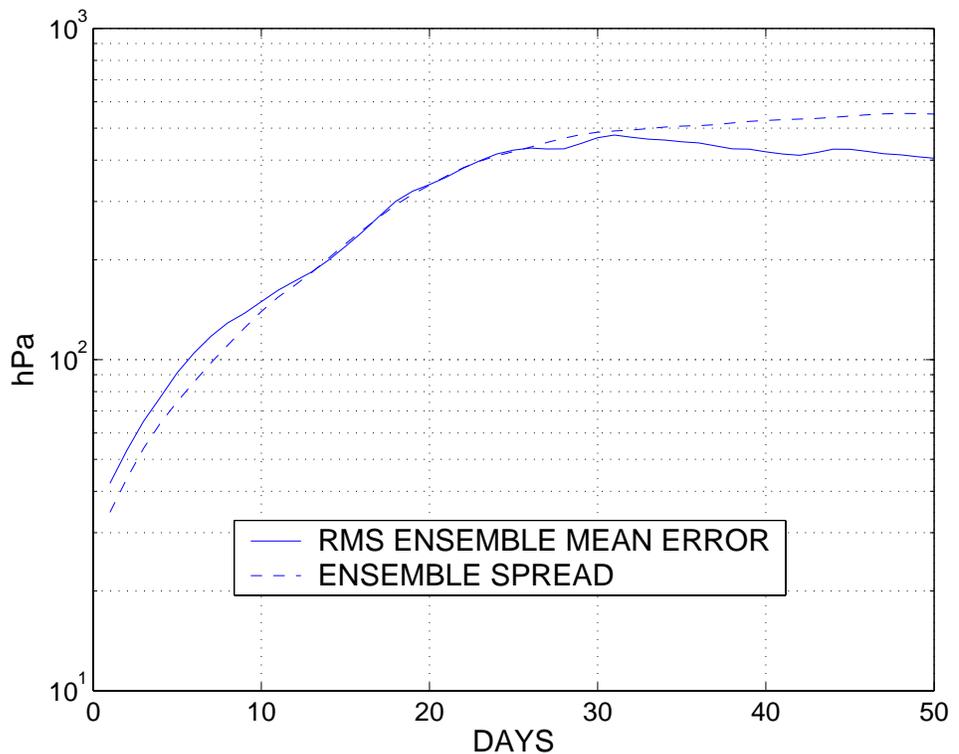
Then hint of exponential after day 15?

Saturates by day 25.



Should expect error growth in real systems to look like this.

Operational systems do not at this point?



## Conclusions and future work

1. (Small) ensemble filter can extract lots of information
2. Increasing temporal density of obs may be very effective
3. In perfect world, surface obs deliver accurate assimilations
4. What can high frequency surface obs do in real assimilation / prediction problems?
5. Bias, bias, and bias are key remaining problems
6. Predictability studies must be done in assimilation / prediction context with stochastic sub-grid scale parameterizations

Next step: Moderate resolution GCM with physics:  
(B-Grid I release?, CAM?, NCEP?)

## Dealing with bias in ensembles is remaining problem

Bayesian Theory supporting filters excludes bias

But, we know there are many violations of the Gaussian assumptions we make for implementation

Need to build an additional a priori model of bias

Covariance inflation and related tricks are one simple model

Have some advantage by retaining correlation structure

Simply States that there is an additional Gaussian component of error that is not accounted for by the model

Can we do more sophisticated, adaptive models?

With ensemble and known observation error distribution, can determine expected value of sum of model and observation bias for any observation

In other words, is the distance between the prior obs estimates and the obs inconsistent?

Can aggregate these statistics in time, or space or both

Need to partition unaccounted error into one of three bins:

1. Model first moment bias (error)
2. Model second moment bias (error)
3. Observation bias (error)

## Dealing with bias in ensemble filters (cont.)

May be easy to partition between 3 and combined (1, 2)

Similar to buddy checks

Are observations in same 'area' not consistently inconsistent

If so, much more inconsistent obs should have large bias

Tricky problem, how to partition bias between first and second moment in model

If it's first moment, just let observation be more compelling

If it's second, need to reduce decrement in spread

Initial results playing with this have been very successful in very large bias systems

Need to try out in a real setting

Note: this should eventually replace a part of quality control