## NonGaussian filters

Douglas Nychka Geophysical Statistics Project National Center for Atmospheric Research

- A mixture filter
- A local- local filter
- A hybrid
- Something completely different



IMAG.e

Supported by the National Science Foundation DMS

## Overview

#### Data Assimilation

Combining predictions made by a numerical model with observed data to estimate the state of a system, x. This is also called a *filter*.

The statistical foundation is Bayes Theorem and the uncertainty in the state of the system is represented by a probability distribution.

## Overview

#### Data Assimilation

Combining predictions made by a numerical model with observed data to estimate the state of a system, x. This is also called a *filter*.

The statistical foundation is Bayes Theorem and the uncertainty in the state of the system is represented by a probability distribution.

PRIOR for  $\boldsymbol{x}$  + observations  $\rightarrow$  POSTERIOR for  $\boldsymbol{x}$ 

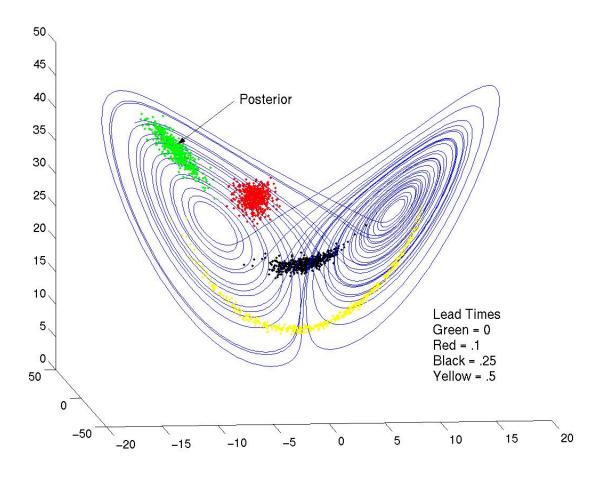
## Effects of nonlinear dynamics

Nonlinear systems can generate assimilation problems where distributions of priors are not well represented by multivariate normals.

Lorenz '63 a simple three dimensional system

Easy to visualize, large distortion occurs for trajectories near the origin. Ensembles are often distinctly non-Gaussian.

## Following an ensemble through state space



### The Bayes cycle

Observations at time t

 $\boldsymbol{y}_t = H \boldsymbol{x}_t + \text{measurement error}$ 

System dynamics:

 $\boldsymbol{x}_{t+1} = g(\boldsymbol{x}_t)$  (deterministic)

 $p(\mathbf{x}_{t}), \mathbf{y}_{t} \xrightarrow{\mathsf{Bayes}} p(\mathbf{x}_{t}|\mathbf{y}_{t}) \xrightarrow{g(.)} p(\mathbf{x}_{t+1}|\mathbf{y}_{t}) = p(\mathbf{x}_{t+1}), \mathbf{y}_{t+1}$ data update forecast new data

Yesterday's posterior becomes today's prior!

#### Some key ideas:

- Represent a continuous distribution by a discrete sample.
- Non-Gaussian update for components of the state vector *local* to the observation locations and close to the observation value.
- Parts of state vector far from the observation are updated using Ensemble Kalman Filter.

### Standard Kalman Filter/ conditional multivariate normal distributions

This is easy in closed form if everything is *multivariate normal* and *linear*.

Observation Model

$$oldsymbol{y} = Holdsymbol{x}_t + oldsymbol{e}$$
 with  $oldsymbol{e} \sim MN(0,R)$ 

Prior

$$\boldsymbol{x}_t \sim MN(\boldsymbol{\mu}_t, \boldsymbol{P})$$

### Standard Kalman Filter/ conditional multivariate normal distributions

This is easy in closed form if everything is *multivariate normal* and *linear*.

Observation Model

$$oldsymbol{y} = Holdsymbol{x}_t + oldsymbol{e} \quad ext{with} \quad oldsymbol{e} \sim MN(0,R)$$

Prior

$$\boldsymbol{x}_t \sim MN(\boldsymbol{\mu}_t, \boldsymbol{P})$$

Kalman update for state

 $\hat{\boldsymbol{x}}_t = E(\boldsymbol{x}_t | \boldsymbol{y}) = \boldsymbol{\mu}_t + \boldsymbol{P} H^T (H \boldsymbol{P} H^T + R)^{-1} (\boldsymbol{y} - H \boldsymbol{\mu}_t)$ 

## The Ensemble KF

- All means and covariances in Kalman Filter are replaced by sample quantities found from the ensemble.
- The sample covariance matrix from the ensemble is tapered spatially to regularized the estimate.
- Essentially the ensemble encodes a low rank approximation to the mean and covariance following the *exact* calculation under the assumption that everything is multivariate normal.

### Forecast step with EnKF

In place of

$$p(\boldsymbol{x}_t|y_t) \rightarrow p(g(\boldsymbol{x}_t)|y_t)$$

propagate each ensemble member.

Note:

With no model error, the relationship among state vectors is preserved correctly. The dynamics generates information ...

Suppose we observe just the  $J^{th}$  state component with noise

$$Y = x_J + e$$

Update the  $i^{th}$  ensemble member:

$$\boldsymbol{x}_{i}^{u} = \boldsymbol{x}_{i}^{f} + COV(\boldsymbol{x}, x_{J})VAR(Y)^{-1}(Y - x_{i,J}^{f} - \text{perturbation})$$

Suppose we observe just the  $J^{th}$  state component with noise

$$Y = x_J + e$$

Update the  $i^{th}$  ensemble member:

$$\boldsymbol{x}_{i}^{u} = \boldsymbol{x}_{i}^{f} + COV(\boldsymbol{x}, x_{J})VAR(Y)^{-1}(Y - x_{i,J}^{f} - \text{perturbation})$$

$$m{x}_i^u = m{x}_i^f + rac{COV(m{x}, x_J)}{VAR(x_J)} rac{VAR(x_J)}{VAR(Y)} imes (Y - x_{i,J}^f - ext{perturbation})$$

Suppose we observe just the  $J^{th}$  state component with noise

$$Y = x_J + e$$

Update the  $i^{th}$  ensemble member:

$$\boldsymbol{x}_{i}^{u} = \boldsymbol{x}_{i}^{f} + COV(\boldsymbol{x}, x_{J})VAR(Y)^{-1}(Y - x_{i,J}^{f} - \text{perturbation})$$

$$m{x}_i^u = m{x}_i^f + rac{COV(m{x}, x_J)}{VAR(x_J)} rac{VAR(x_J)}{VAR(Y)} imes (Y - x_{i,J}^f - ext{perturbation})$$

$$oldsymbol{x}_i^u = oldsymbol{x}_i^f + rac{COV(oldsymbol{x}, x_J)}{VAR(x_J)} \quad imes ( ext{ draw from } p(x_J|Y) - x_{i,J}^f)$$

Suppose we observe just the  $J^{th}$  state component with noise

$$Y = x_J + e$$

Update the  $i^{th}$  ensemble member:

$$\boldsymbol{x}_{i}^{u} = \boldsymbol{x}_{i}^{f} + COV(\boldsymbol{x}, x_{J})VAR(Y)^{-1}(Y - x_{i,J}^{f} - \text{perturbation})$$

$$m{x}_i^u = m{x}_i^f + rac{COV(m{x}, x_J)}{VAR(x_J)} rac{VAR(x_J)}{VAR(Y)} imes (Y - x_{i,J}^f - ext{perturbation})$$

$$oldsymbol{x}_i^u = oldsymbol{x}_i^f + rac{COV(oldsymbol{x}, x_J)}{VAR(x_J)} \quad imes ( ext{ draw from } p(x_J|Y) - x_{i,J}^f)$$

Now substitute sample COV and sample VAR to get a linear regression.

Represent the prior distributions as mixtures of multivariate normals  $p(\mathbf{x}_t) = \sum_{i=1}^k p_i \mathsf{MN}(\boldsymbol{\mu}_i, \mathbf{P}_i)$ 

The posterior distribution is also a mixture:  $p(\mathbf{x}_t | \mathbf{y}_t) = \sum_{i=1}^k p_i^{\star} \mathsf{MN}(\boldsymbol{\mu}_i^{\star}, \mathbf{P}_i^{\star})$ 

Each component is just the usual KF or EnKF update!

Represent the prior distributions as mixtures of multivariate normals  $p(\mathbf{x}_t) = \sum_{i=1}^k p_i \mathsf{MN}(\boldsymbol{\mu}_i, \mathbf{P}_i)$ 

The posterior distribution is also a mixture:  $p(\mathbf{x}_t | \mathbf{y}_t) = \sum_{i=1}^k p_i^{\star} \mathsf{MN}(\boldsymbol{\mu}_i^{\star}, \mathbf{P}_i^{\star})$ 

Each component is just the usual KF or EnKF update!

Key aspect is the update of the mixture probability

$$p_i^{\star} \sim \frac{e^{-(y-H\mu_i)^T(H\mathbf{P}_iH^T+R)^{-1}(y-H\mu_i)}}{|(H\mathbf{P}_iH^T+R)|^{1/2}}$$

 $p_i$  is large if  $H\mu_i$  is close to the observed value.

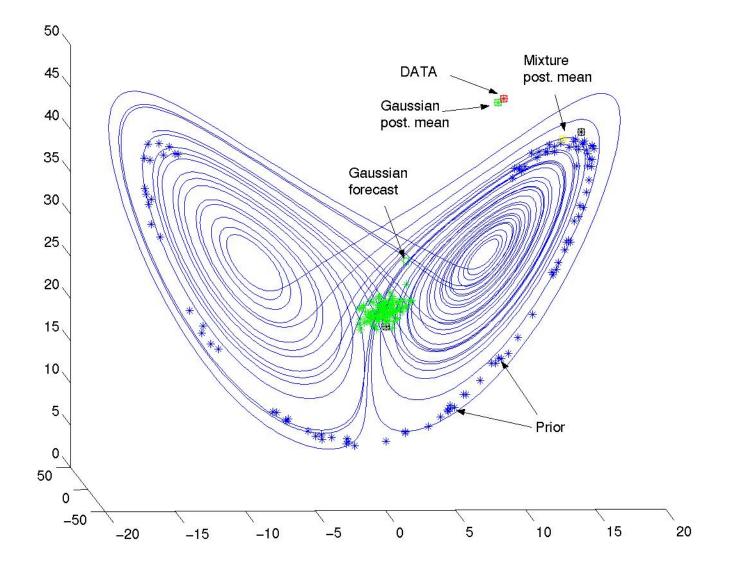
Ensembles as a mixture distribution:

A random subset of ensemble members are the centers of the mixture.  $(\mu_k)$ .

The sample covariance of nearest neighbors about the center is the covariance  $(P_k)$ .

The posterior probabilities look like weights based on a normal kernel. The use of neighborhoods to find the covariance results in a local linear regression.

#### A non-Gaussian example



Results for Lorenz '63

With a time step of .5 and observation standard deviation of 2 in the saddle region of the state space (400 ensemble members, 100 centers, 25 Nearest neighbors.)

RMSE Ensemble KF filter = 1.64RMSE Mixture filter = .73

### Extensions to larger state spaces

The mixture filter breaks down as the dimension of the state vector increases.

e.g. the posterior probability concentrates on a single member or is small

## The Local-Local Filter

LOCAL in physical space:

Only update components close to the observation location.

Call these local state components  $\boldsymbol{x}_L$ 

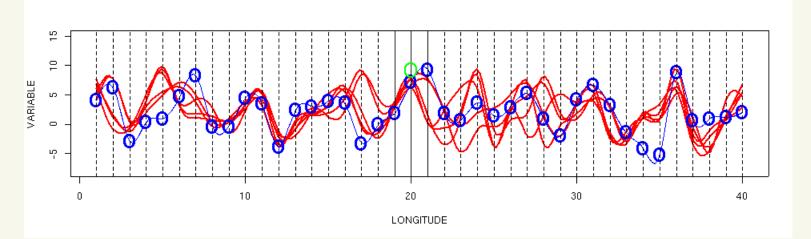
LOCAL in state space:

only use ensemble members that are "close" to the observed value.

Hybrid filter Update remaining components using EnKF.

Call these remaining (global) components  $oldsymbol{x}_G$ 

### A 40 dimensional system: Lorenz '96



True state Single Observation at 20 5 ensemble members (as lines)

Local state components: 19, 20, 21 Global components of state: the rest!

### How do we splice two types of solutions together?

 $Our \ first \ idea$ 

Find the global part conditional on first finding the local posterior.

 $p(\boldsymbol{x}|y) = p(\boldsymbol{x}_G|\boldsymbol{x}_L)p(\boldsymbol{x}_L|y)$ 

 $p(\boldsymbol{x}_L|y)$  from local-local (non-Gaussian) filter  $p(\boldsymbol{x}_G|\boldsymbol{x}_L)$  assuming Gaussian distributions.

### How do we splice two types of solutions together?

 $Our \ first \ idea$ 

Find the global part conditional on first finding the local posterior.

 $p(\boldsymbol{x}|y) = p(\boldsymbol{x}_G|\boldsymbol{x}_L)p(\boldsymbol{x}_L|y)$ 

 $p(\boldsymbol{x}_L|y)$  from local-local (non-Gaussian) filter  $p(\boldsymbol{x}_G|\boldsymbol{x}_L)$  assuming Gaussian distributions.

This did not work ...

What worked as a hybrid filter.

Use posterior means from  $p(\boldsymbol{x}_G | \boldsymbol{x}_L)$  and  $p(\boldsymbol{x}_L | y)$ The non Gaussian gives good point predictions. What worked as a hybrid filter.

Use posterior means from  $p(\boldsymbol{x}_G | \boldsymbol{x}_L)$  and  $p(\boldsymbol{x}_L | y)$ The non Gaussian gives good point predictions.

Reverse conditioning!  $[\boldsymbol{x}_L | \boldsymbol{x}_G]$ Posterior for  $\boldsymbol{x}_G$  is just  $p(\boldsymbol{x}_G | Y)$ 

Posterior draw for  $x_L$  is ( up to the mean)

 $E[\boldsymbol{x}_L|\boldsymbol{x}_G] + [A\boldsymbol{x}_L|y]$ 

The matrix A is chosen so that complete ensemble agrees with posterior covariance for EnKF.

# Summary

Results

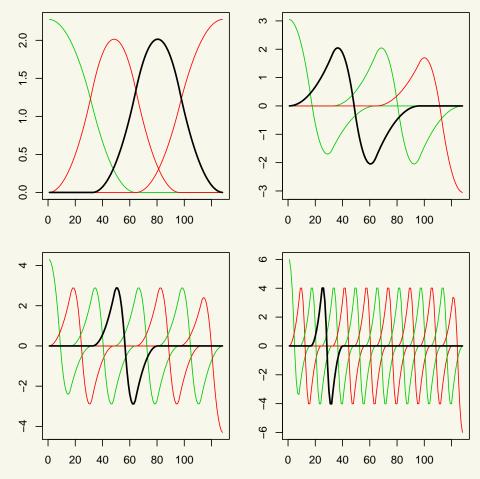
- We have some evidence that the practical version of the EnKF actually handles non-Gaussian distributions better than an exact Kalman filter.
- The Local-Local filter clearly out performs EnKF in a simple 3-d system especially in places where g is very nonlinear.
- A version of the L-L filter also performs better than EnKF with about 5% improvement (without any extensive tuning) for the 40 variable model.

#### Issues

- The local linear fitting seems important but it is hard to beat the EnKF
- The spread of the ensemble may have nothing to do with Bayes; All that is important is to generate a good regression relationship.
- RMSE as a criterion has little to do with non-Gaussian distributions.
- Are components far away from observations more amenable to a Gaussian update?

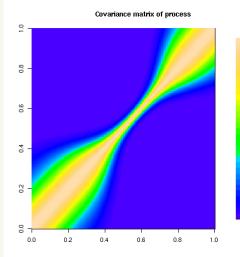
### Better models for covariance functions

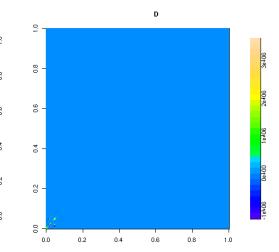




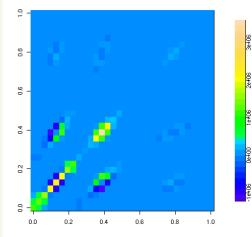
# Sparseness with wavelets

Decomposition of a 1-d covariance matrix





32X32 sub-matrix of D



32X32 sub-matrix of H where D= H\*H

