

Some Outstanding Issues in the Jet Formation Mechanisms

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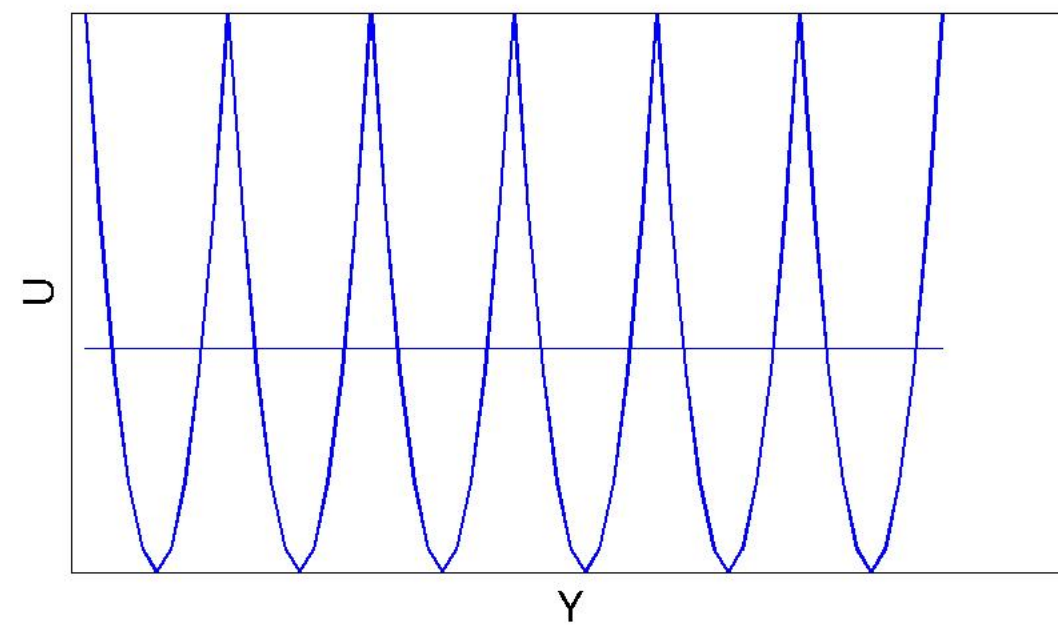
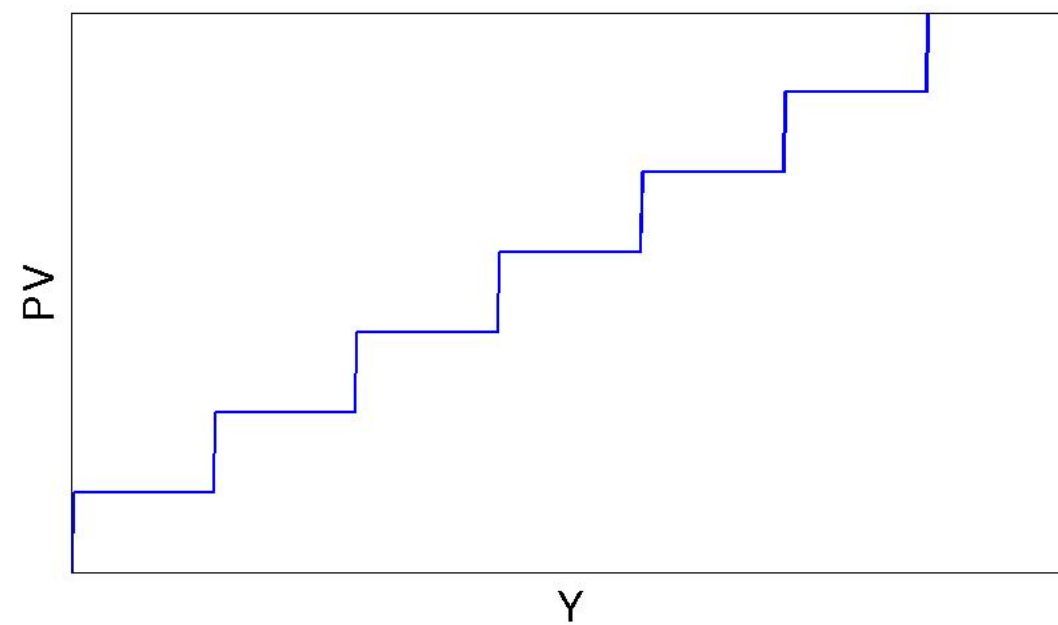


Brief Survey of Jovian Atmosphere

- Gas giant planet
- Convective interior with high temperature
($\sim 10000K$)
- Six alternating jets in each hemisphere

Two perspectives concerning the jet formation in Jupiter's weather layer:

- geostrophic turbulence on β -plane (Rhines, JFM, 1975; Williams, JAS, 1978, 1979; Maltrud, M. and G.K. Vallis. Physics of Fluids A, 1993)
- Potential Vorticity (PV) staircase model
 - turbulence-PV mixing feedback (Baldwin et al, Science, 2007)
 - formation of westly jets and PV mixing (Dritschel and McIntyre, JAS, 2008)



schematical illustration of $u(y)$ for infinite number of PV steps

Different emphasis of these two perspectives

- Geophysical Turbulence—small-scale turbulence and large-scale wave-like motion (upscale energy cascade)
- PV staircase model— understanding the PV structure (and use invertibility to back out flow)

Our goal for this problem

- understand and predict the PV structure

Two experiments

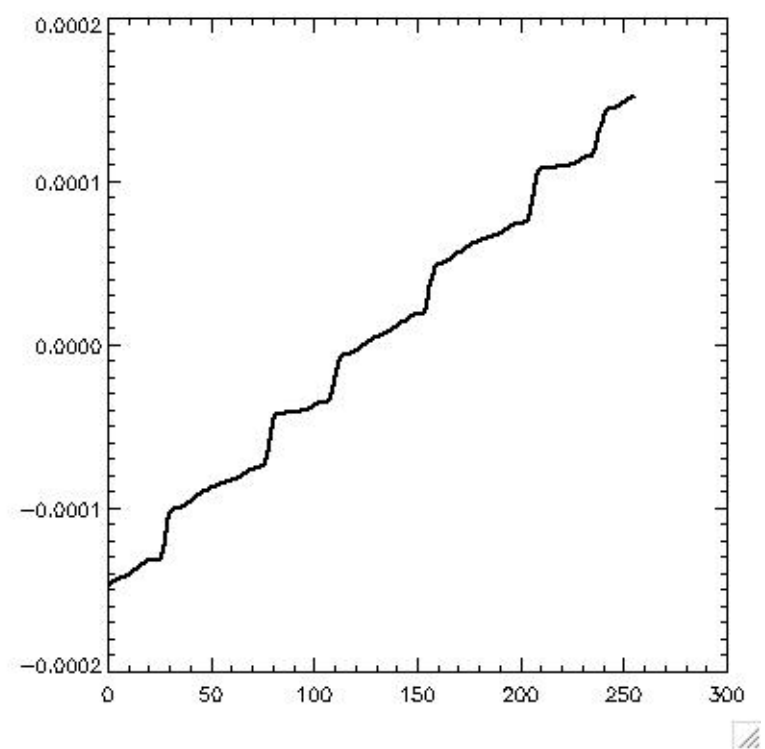
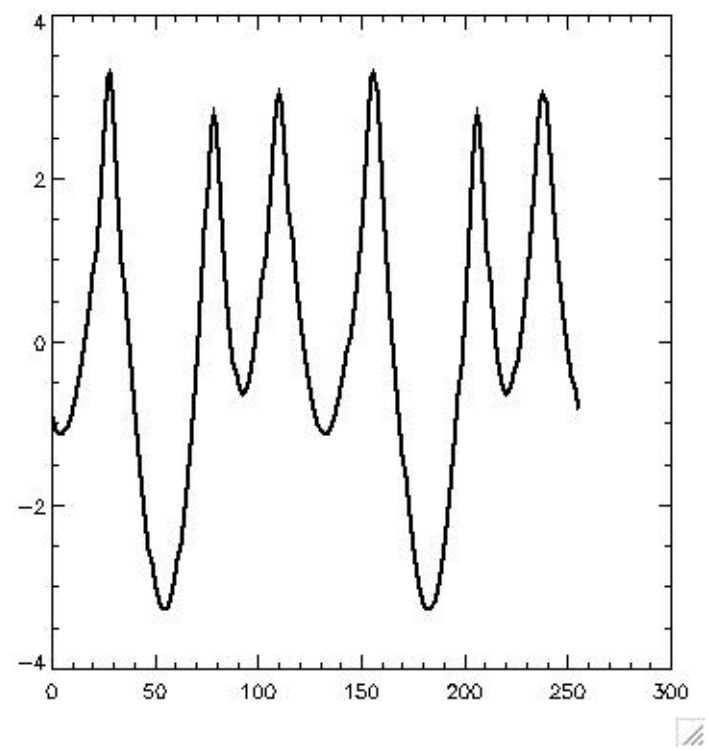
- beta-plane turbulence
- stationary Rossby wave breaking

Geostrophic Turbulence Experiment

Barotropic Vorticity Equation on β -plane with isotropic small scale random forcing(F) and linear damping ($\bar{\gamma}$) in the zonal direction

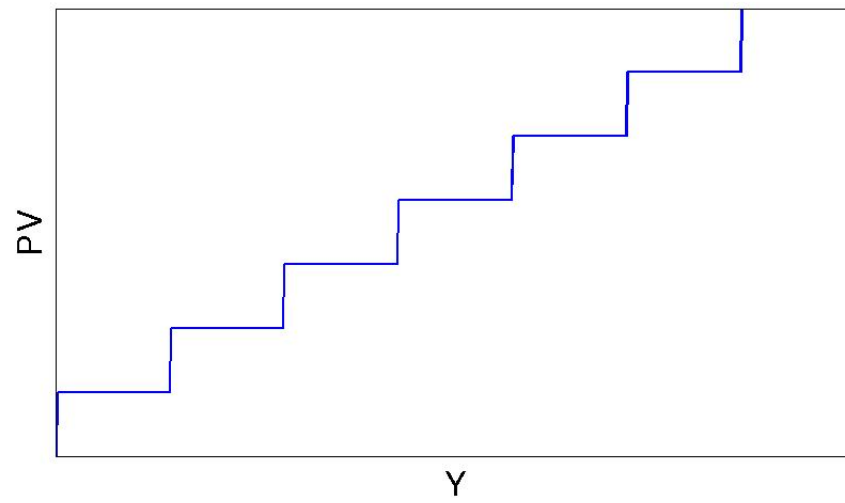
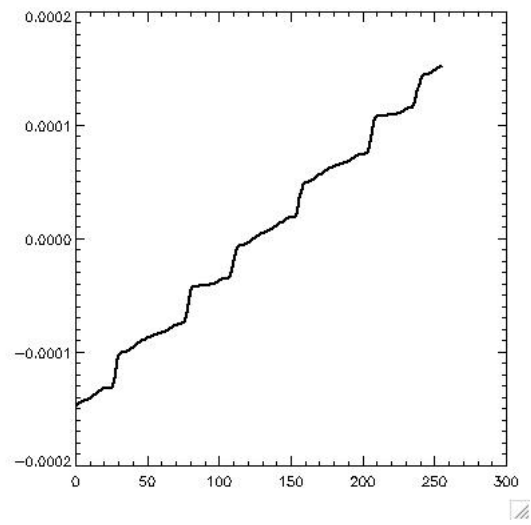
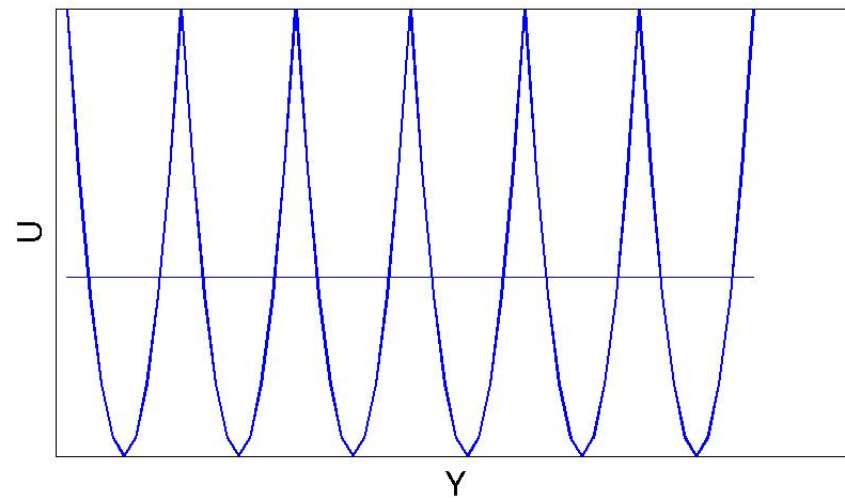
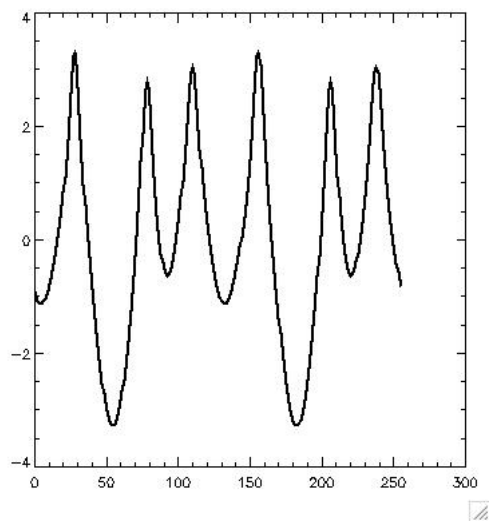
$$\frac{\partial \nabla^2 \psi}{\partial t} + J(\psi, \nabla^2 \psi) + \beta \frac{\partial \psi}{\partial x} = -\nu \nabla^6 \psi + \bar{\gamma} U_y + F$$

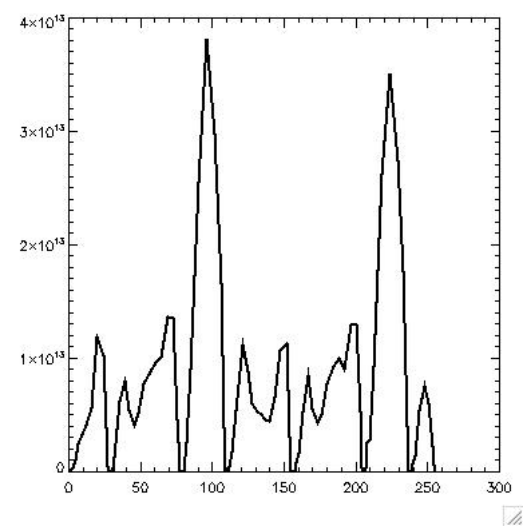
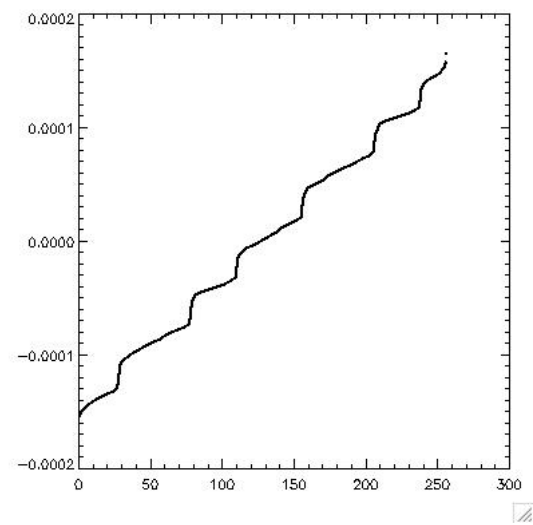
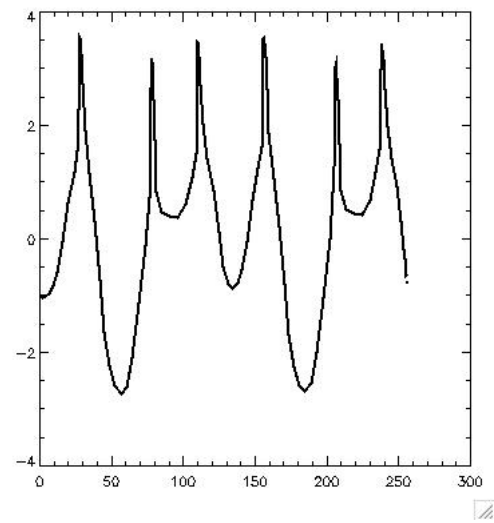
- Spatial resolution: 256×256 (T228)
- Periodic in zonal and meridional direction
- effect of anisotropic damping is very modest compared to isotropic result (Zhu and Nakamura, AMS poster, 2007; Danilov and Gurarie, Physical Review E, 2002)



time mean zonal mean(TMZM) U and PV vs. latitude

comparison to PV staircase model





TMZM U, PV, effective diffusivity(κ_{eff}) vs. equivalent latitude

Q1: Why is PV not homogenized between the steps in the geostrophic turbulence? Is it due to relaxation ($\bar{\gamma}$)?

Q2: Is the jet spacing considered to be Rhines scale?

Q3: Is the mixing done by the large scale or small scale eddies?

Rossby Wave Breaking Experiment

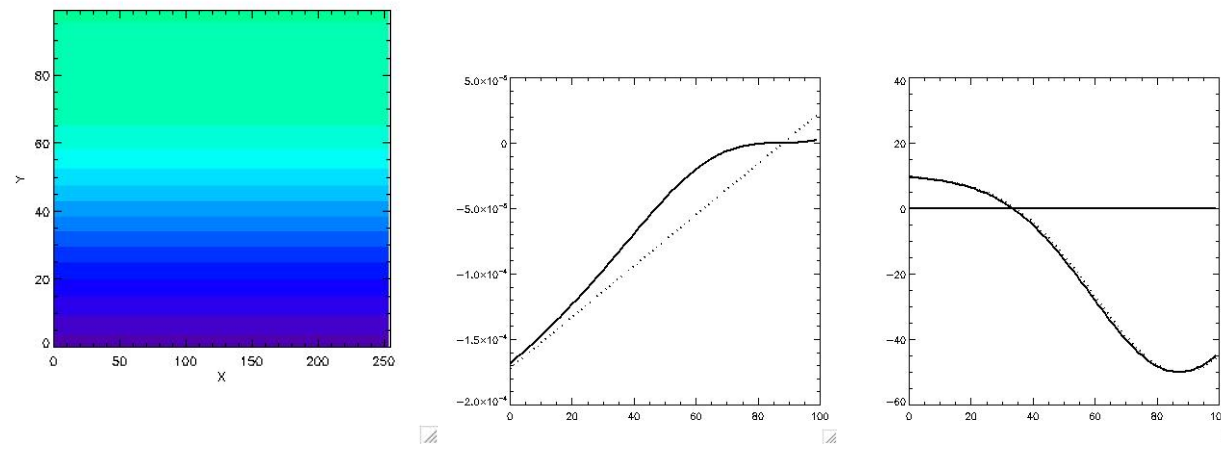
Forced-dissipative QG β plane model

$$\frac{\partial \nabla^2 \psi}{\partial t} + J(\psi, \nabla^2 \psi) + \beta \frac{\partial \psi}{\partial x} = -\nu \nabla^6 \psi - T(x, y_0, \delta)$$

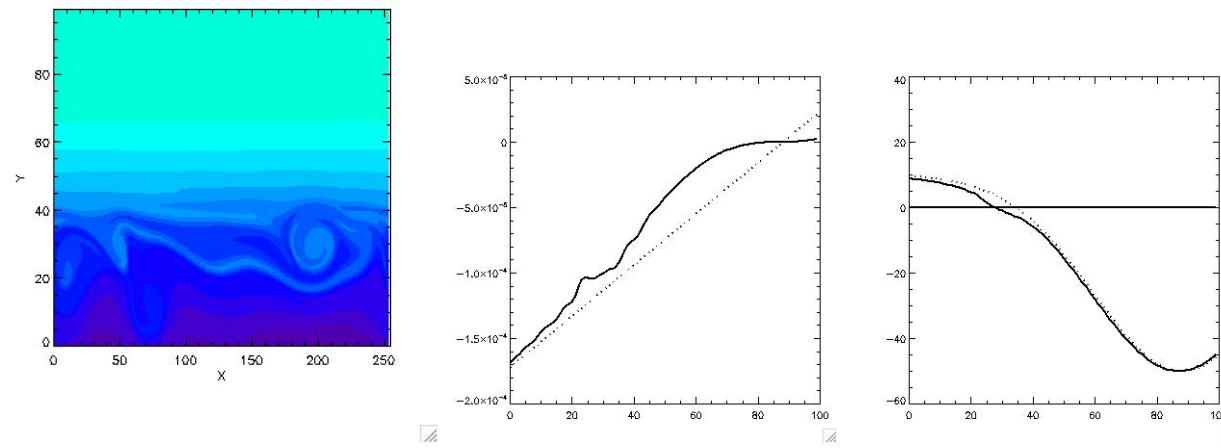
where $T(x, y_0, \delta) = \sin(\hat{k}x)G(y_0, \delta)$ mimics the mountain sinusoidal in x direction and bell-shaped(centered at y_0) in y direction. The wind profile is westerly over the mountain.

PV is expected to be irreversibly rearranged and mixed at the critical latitude $U = 0$.

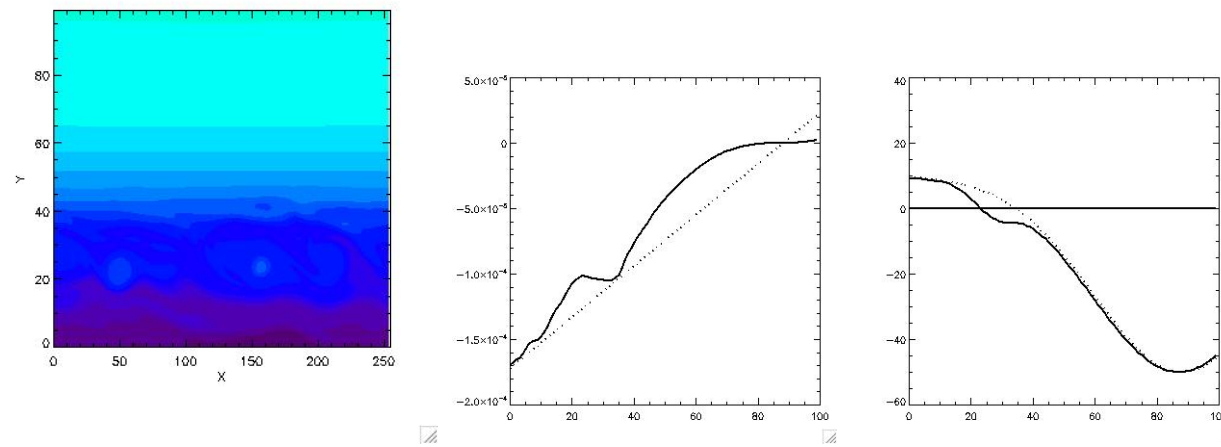
PV contours / zonal mean PV / zonal mean U



$t=0$

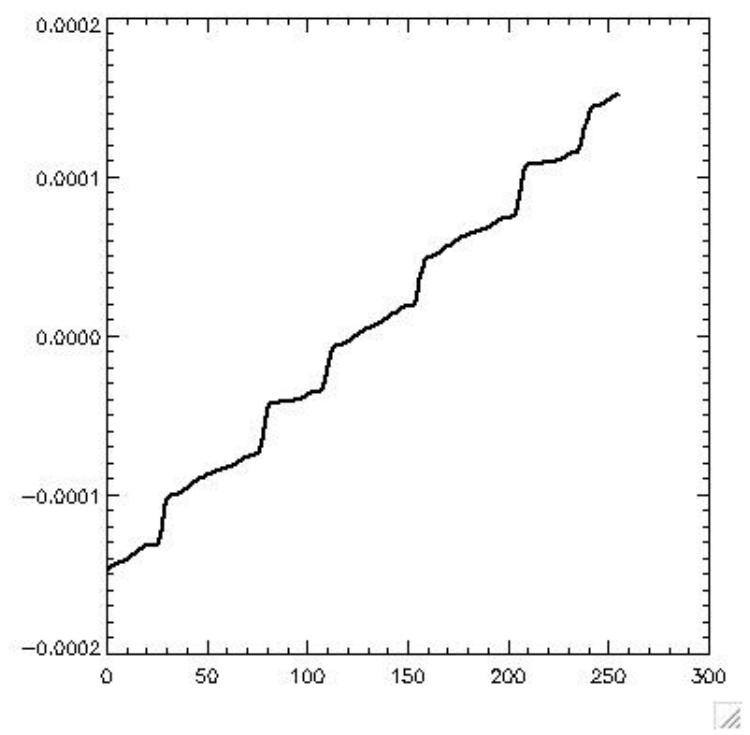
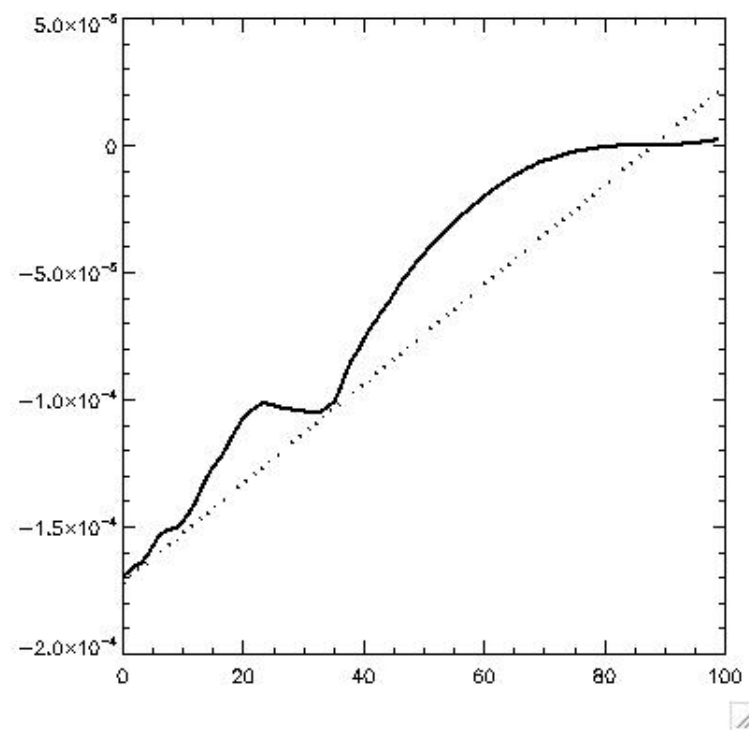


$t=26000$



$t=78000$

comparison to beta-plane results



1-D Diffusion-Reaction Model

Idea: Relaxation($\bar{\gamma}$) + Effective Diffusivity(κ_{eff}) \Rightarrow PV slope ?

Governing Equation:

$$0 \approx \frac{\partial q}{\partial t} = \frac{\partial}{\partial y} \left(K_{eff}(y) \frac{\partial q}{\partial y} \right) - \gamma(q - \beta y)$$

with boundary Condition:

$$\begin{aligned} \frac{dq}{dy}(0) &= \frac{dq}{dy}(L) \\ \frac{d^2q}{dy^2}(0) &= \frac{d^2q}{dy^2}(L) \end{aligned}$$

(1) Blue dot-dashed—constant $\max(K_{eff}), \gamma_1$

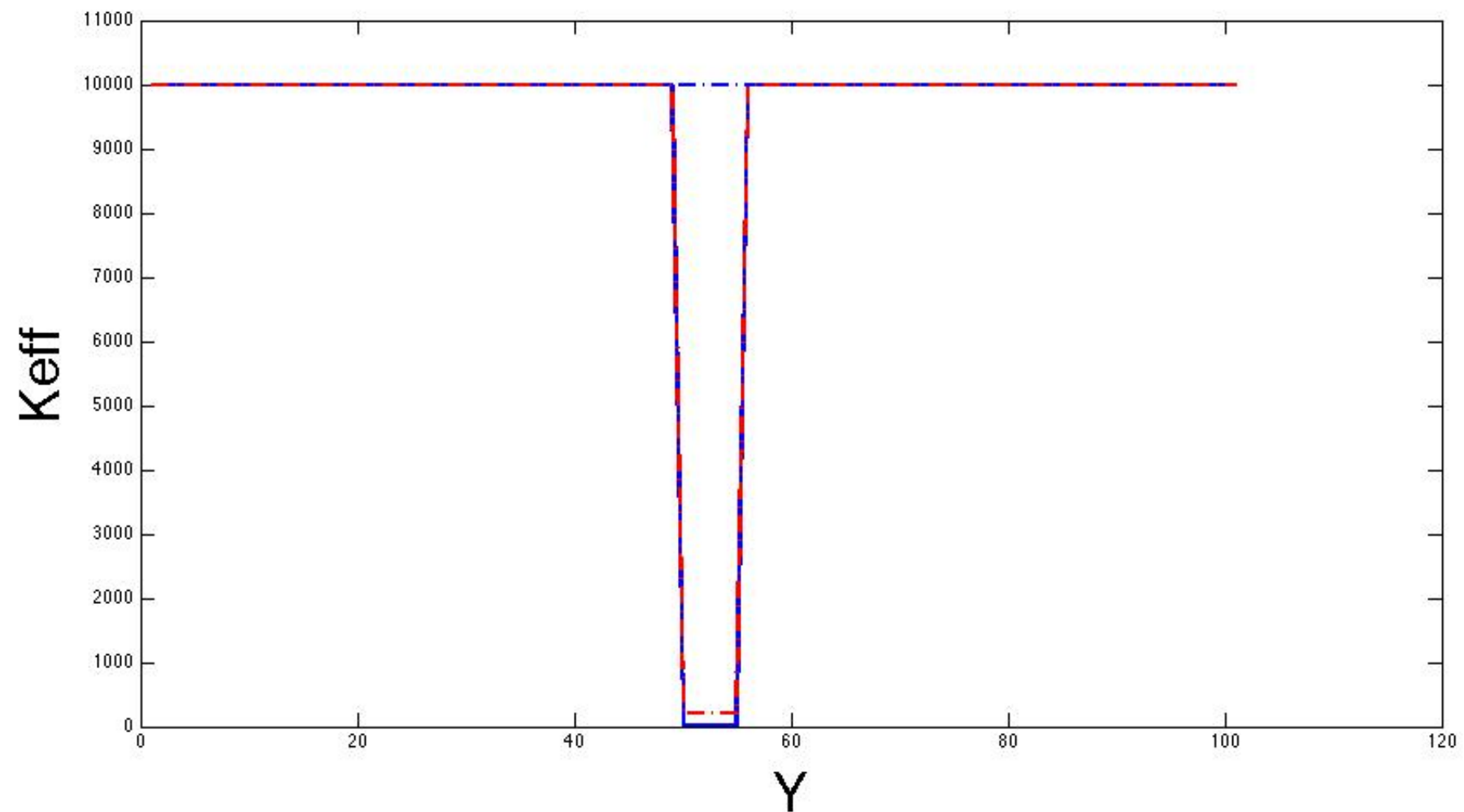
(2) Blue SOLID line— $\min(K_{eff1}), \gamma_1$

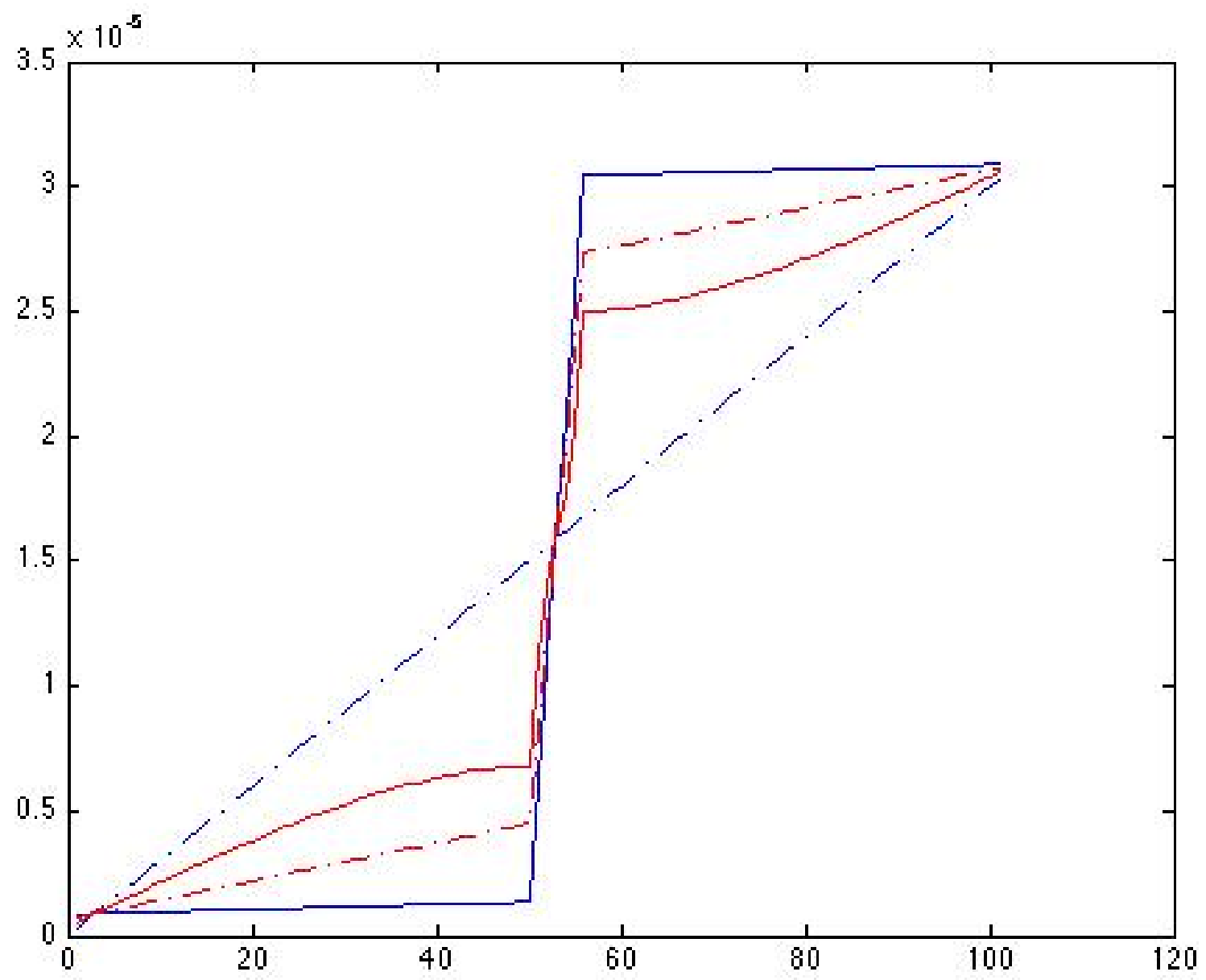
(3) Red dot-dashed line— $\min(K_{eff2}), \gamma_1$

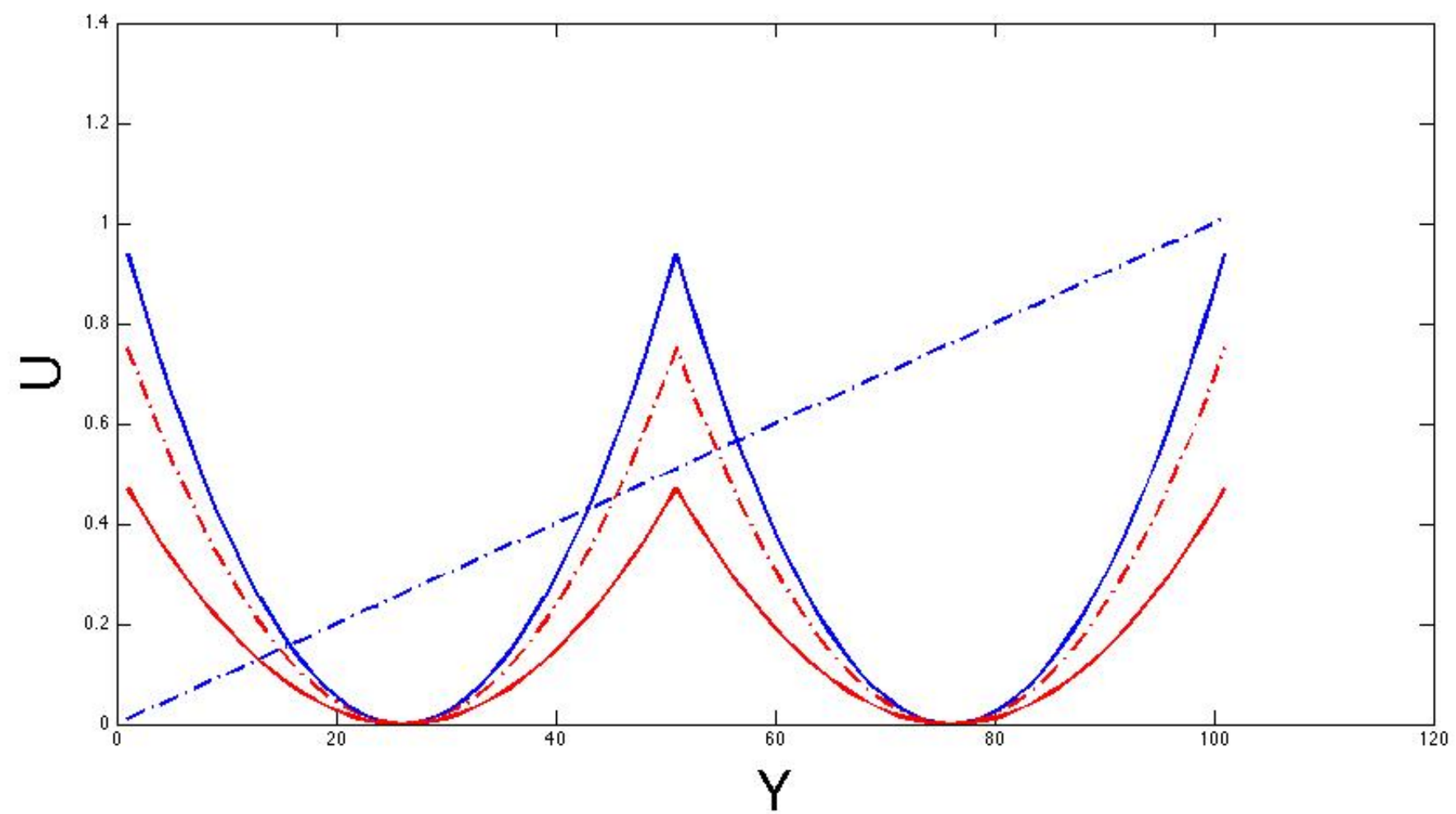
($\min(K_{eff2}) > \min(K_{eff1})$)

(4) Red SOLID line— $\min(K_{eff1}), \gamma_2$

($\gamma_2 > \gamma_1$)







Summary

- the PV gradient in the surf zone is determined by the combination of the relaxation γ and $\min(K_{eff})$, NOT K_{eff} in the surf zone. This finding is consistent with Nakamura's most recent result. (Nakamura, JAS, to appear, 2008)
- getting the "leakiness" of the barriers($\min(K_{eff})$) correctly is important to predict PV profile (and hence the flow) correctly.
- Wave-mean flow interaction and barotropic instability might have an important impact in homogenizing the PV within some particular regions.