# Current progress in data assimilation with the Local Ensemble Transform Kalman Filter

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- http://www.weatherchaos.umd.edu
- Mathematical details: B. R. Hunt, E. K., I. Szunyogh, Physica D **230** (2007), 112–126
- Review paper on GFS: I. Szunyogh, E. K., *et al.*, Tellus A **60** (2008), 113–130
- Bias estimation: S. J. Baek, B. R. Hunt, E. Kalnay, E. Ott, I. Szunyogh, Mon. Wea. Rev. **139** (2009), 2349–2364.
- ECOM model: R. N. Hoffman, *et al.*, J. Atmos. Ocean Tech. **25** (2008), 1638–1656.

- The basic approach
- Bias estimation
- Extensions: non-Gaussianity, nonlinear H

### Introduction: the estimation problem

- Observations:  $\mathbf{y}_o = \mathbf{H}(\mathbf{x}_t) + \boldsymbol{\varepsilon}$  with  $\boldsymbol{E}(\boldsymbol{\varepsilon}) = \mathbf{0}$  and  $\boldsymbol{E}(\boldsymbol{\varepsilon}\boldsymbol{\varepsilon}^{\mathrm{T}}) = \mathbf{R}$
- Forecast:  $\mathbf{x}_b = \mathbf{x}_t + \boldsymbol{\eta}$  with  $E(\boldsymbol{\eta}) = \mathbf{0}$  and  $E(\boldsymbol{\eta} \boldsymbol{\eta}^{\mathrm{T}}) = \mathbf{P}_b$
- Minimize the quadratic cost function:  $J(\mathbf{x}) = (\mathbf{y}_o \mathbf{H}(\mathbf{x}))^T \mathbf{R}^{-1} (\mathbf{y}_o \mathbf{H}(\mathbf{x})) + (\mathbf{x} \mathbf{x}_b)^T \mathbf{P}_b^{-1} (\mathbf{x} \mathbf{x}_b)$
- Current NCEP operations: ~ 1.75 million obs assimilated into a ~ 3 billion variable model every 6 hours

- When the errors are Gaussian and the underlying dynamics are linear, the minimizer of *J* is "optimal" (unbiased, minimum variance)
- To evaluate J, we must invert **R** and **P**<sub>b</sub>
- The observation covariance matrix **R** is a nearly diagonal  $p \times p$  matrix with  $p \sim 10^6 10^7$
- $\mathbf{P}_b$  is not diagonal and is  $n \times n$ , where  $n \sim 10^7 10^9$  for typical weather models
- Ensemble Kalman filters attempt a low-rank empirical estimate of **P**<sub>b</sub>

# Our approach: Use dynamics to reduce the dimensionality

- Key finding: Over typical synoptic regions, the forecast uncertainty evolves in a much lower-dimensional space than the phase space
- Example: The Global Forecast System at T62 resolution has ~ 3000 variables in a typical 1000 × 1000-km synoptic region
- An ensemble of 100-200 forecast vectors typically spans a  $\sim 40$ -dimensional subspace over short forecast periods

# The LETKF algorithm: update overlapping local regions

- Consider all observations with a prespecified distance of each grid point
- Given k ensemble forecasts, change coordinates to the (k-1)-dimensional subspace that they span
- Compute the Kalman update in these coordinates
- Naturally parallel algorithm



- The LETKF and 4DVar attempt to minimize the same type of cost function
- LETKF uses a flow-dependent background covariance (determined empirically from an ensemble of forecasts)
- 4DVar requires integration of the nonlinear model and its linearization
- LETKF is model independent

- NCEP Global Forecast System (GFS)
- ECOM (Estuarine & Coastal Ocean Model)—with Ross Hoffman
- NCEP Regional Spectral Model—Dasa Merkova
- CO and O<sub>3</sub> assimilation—Dave Kuhl
- GFDL model (Mars Microwave Sounder)—John Wilson, Matt Hoffman
- CAM/CASA' model (OCO)—Junjie Liu, Inez Fung
- CPTEC and JMA working towards operational implementation

# Other advantages of the LETKF

- Uses flow-dependent covariances, including the analysis uncertainty
- Assimilates all data at once—full 4-d scheme
- Nonlinear observation operators are readily accommodated
- The local region and ensemble size are the only free parameters
- No adjoint necessary!
- Model parameters can be estimated as part of the state vector

- Goal: Find the linear combination of ensemble solutions that best fits the observations
- Start with background ensemble  $\{\mathbf{x}_b^i\}_{i=1}^k$  with mean  $\bar{\mathbf{x}}_b$  and perturbation matrix

$$\mathbf{X}_b = \left( \mathbf{x}_b^1 - \bar{\mathbf{x}}_b \ \mathbf{x}_b^2 - \bar{\mathbf{x}}_b \ \cdots \ \mathbf{x}_b^k - \bar{\mathbf{x}}_b \right)$$

- Compute the analysis mean  $\bar{\mathbf{x}}_a$  and analysis ensemble perturbations  $\mathbf{X}_a = \mathbf{X}_b \mathbf{W}_a$
- The analysis is  $\mathbf{X}_a + \bar{\mathbf{x}}_a$

# Change of basis to the background perturbations

- The ensemble covariance matrix  $\mathbf{P}_b = (k-1)^{-1} \mathbf{X}_b \mathbf{X}_b^{\mathrm{T}}$ has rank k-1
- Since  $X_b$  is 1–1 on its column space *S*, we minimize *J* on *S*, where  $P_b^{-1}$  is defined
- Treat  $\mathbf{X}_b$  as a linear transformation from an abstract *k*-dimensional space  $\tilde{S}$  to S
- If  $\mathbf{w} \in \tilde{S}$  is Gaussian with mean **0** and covariance  $(k-1)^{-1}\mathbf{I}$ , then  $\mathbf{x} = \bar{\mathbf{x}}_b + \mathbf{X}_b \mathbf{w}$  is Gaussian with mean  $\bar{\mathbf{x}}_b$  and covariance  $\mathbf{P}_b$

# The objective function

• The cost function

$$J(\mathbf{x}) = (\mathbf{x} - \mathbf{x}_b)^{\mathrm{T}} \mathbf{P}_b^{-1} (\mathbf{x} - \mathbf{x}_b) + (\mathbf{y}_o - \mathbf{H}(\mathbf{x}))^{\mathrm{T}} \mathbf{R}^{-1} (\mathbf{y}_o - \mathbf{H}(\mathbf{x}))$$

#### becomes

$$\widetilde{J}(\mathbf{w}) = (k-1)^{-1} \mathbf{w}^{\mathrm{T}} \mathbf{w} + [\mathbf{y}_{0} - \mathbf{H}(\bar{\mathbf{x}}_{b} + \mathbf{X}_{b} \mathbf{w})]^{\mathrm{T}} \mathbf{R}^{-1} [\mathbf{y}_{0} - \mathbf{H}(\bar{\mathbf{x}}_{b} + \mathbf{X}_{b} \mathbf{w})]$$
  
and we linearize  $\mathbf{H}(\bar{\mathbf{x}}_{b} + \mathbf{X}_{b} \mathbf{w}) \approx (\bar{\mathbf{y}}_{b} - \mathbf{Y}_{b} \mathbf{w})$  where

$$\mathbf{\bar{y}}_b = \mathbf{H}(\mathbf{x}_b^i)$$
 and  $\mathbf{Y}_b = (\mathbf{H}(\mathbf{x}_b^1) - \mathbf{\bar{y}}_b \cdots \mathbf{H}(\mathbf{x}_b^k) - \mathbf{\bar{y}}_b)$ 

- Only the components of  $\mathbf{H}(\mathbf{x}_b^i)$  belonging to the current local region are selected to form  $\bar{\mathbf{y}}_b$  and  $\mathbf{Y}_b$
- If  $\mathbf{w}_a$  minimizes  $\tilde{J}$ , then  $\bar{\mathbf{x}}_a = \bar{\mathbf{x}}_b + \mathbf{X}_b \mathbf{w}_a$  minimizes the original J
- Modifications are required for satellite radiance observations

# The analysis ensemble

- The minimizer  $\mathbf{w}_a = \mathbf{Q}\mathbf{Y}_b^{\mathrm{T}}\mathbf{R}^{-1}(\mathbf{y}_o \bar{\mathbf{y}}_b)$ , where  $\mathbf{Q} = [(k-1)\mathbf{I} + \mathbf{Y}_b^{\mathrm{T}}\mathbf{R}^{-1}\mathbf{Y}_b]^{-1}$
- The analysis mean is  $\bar{\mathbf{x}}_a = \bar{\mathbf{x}}_b + \mathbf{X}_b \mathbf{w}_a$  with covariance matrix  $\mathbf{P}_a = \mathbf{X}_b \mathbf{Q} \mathbf{X}_b^{\mathrm{T}}$
- The analysis perturbations are  $\mathbf{X}_a = \mathbf{X}_b \mathbf{W}_a$  where  $\mathbf{W}_a = [(k-1)\mathbf{Q}]^{1/2}$  and we take the symmetric square root
- W<sub>a</sub> then depends continuously on Q and assures that the columns of X<sub>a</sub> sum to 0 (for the correct sample mean)

# Computational efficiency

- The forecasts are the most expensive part!
- Given an ensemble of size *k* and *s* observations in a local region
- Most expensive step ( > 90% of cpu cycles): computing  $\mathbf{Y}_{b}^{\mathrm{T}}\mathbf{R}^{-1}\mathbf{Y}_{b}$ , which is  $O(k^{2}s)$
- Second most expensive step: computing the symmetric square root, which is  $O(k^3)$
- Observation lookup is  $O(\log L)$  where L is the size of the total observation set
- Wall-clock time: 40 sec on 16 quad-core processors for 500,000 obs and T62 GFS with a 40-member ensemble

- Reference system ("truth"): NCEP operational analyses (T254L64) using the SSI (Parris & Derber, 1992) with all available data (including satellite radiances), truncated to T62L28
- Benchmark system: NCEP operational system, omitting radiance data, with forecasts run at T62L28
- Comparison: (48-h forecast truth)<sup>2</sup> with forecasts started from the LETKF analysis and from the Benchmark system

### 48-hour forecast error: temperature in extratropics





Geographical comparison of 48-h forecast error



- Many sources of model bias: finite resolution, parametrizations of subgrid processes, errors in boundary conditions, etc.
- Baek *et al.* (2006 and 2009) proposed methods for doing local ensemble Kalman filtering in presence of model bias

- Atmospheric state:  $\mathbf{u}(\mathbf{r}, t_n) = \mathbf{u}_n$
- Atmospheric dynamics:  $\mathbf{u}_n = \mathscr{F}(\mathbf{u}_{n-1})$
- Forecast model:  $\mathbf{x}_n = \mathbf{f}(\mathbf{x}_{n-1})$  approximates  $\mathbf{u}_n$
- Assumption: There exists a projection  $\mathscr{P}$  from the infinite-dimensional space of the atmospheric dynamics to the finite-dimensional model space
- Data assimilation computes an estimate  $\mathbf{x}_n^a$  of  $\mathscr{P}(\mathbf{u}_n)$  using the forecast  $\mathbf{x}_n^b$  and observations  $\mathbf{y}_n^o$

# Kalman filter data assimilation scheme

- Forecast:  $\mathbf{x}_n^b = \mathbf{f}(\mathbf{x}_{n-1}^a)$
- Analysis:  $\mathbf{x}_n^a = \mathbf{x}_n^b + \mathbf{K}_n \left[ \mathbf{y}_n^o \mathbf{H}(\mathbf{x}_n^b) \right]$
- Observation operator:  $\mathbf{y}_n^o = \mathbf{H}(\mathbf{x}_n^b) + \boldsymbol{\varepsilon}_n$
- "Noise:"  $\varepsilon_n \sim N(\mathbf{0}, \mathbf{R}_n)$ , assuming  $\mathbf{x}_n^b \approx \mathscr{P}(\mathbf{u}_n)$

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- "Noise:"  $\varepsilon_n \sim N(\mathbf{0}, \mathbf{R}_n)$ , assuming  $\mathbf{x}_n^b \approx \mathscr{P}(\mathbf{u}_n)$
- Model error:  $\mathscr{P}(\mathbf{u}_n) = \mathbf{f}(\mathscr{P}(\mathbf{u}_{n-1})) + \mathbf{b}_n$
- Model error evolution:  $\mathbf{b}_n = \mathbf{G}(\mathbf{b}_{n-1})$

# Two potential approaches to handle model bias

- Bias model I: Correct the background using the estimated bias and compute the analysis using the bias-corrected background and latest observations
- Assumption: The model is subject to the same error whether it is started from the projection 𝒫(u<sub>n-1</sub>) or from the best estimate thereof, x<sup>a</sup><sub>n-1</sub>

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- Bias model I: Correct the background using the estimated bias and compute the analysis using the bias-corrected background and latest observations
- Assumption: The model is subject to the same error whether it is started from the projection 𝒫(**u**<sub>n-1</sub>) or from the best estimate thereof, **x**<sup>a</sup><sub>n-1</sub>
- Bias model II: Assume the forecast dynamics evolve on a model attractor that is shifted from the "true" one
- Assumption: For every finite-dimensional projection *P*(**u**<sub>n</sub>) of the "true" atmospheric state, there is a corresponding point *P* on the model attractor

# Schematic illustration



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# Revised Kalman filter implementation

# • Bias model I:

$$\begin{aligned} \mathbf{x}_n^b &= \mathbf{f}(\mathbf{x}_{n-1}^a) + \mathbf{b}_n \\ \mathbf{x}_n^a &= \mathbf{x}_n^b + \mathbf{K}_n \left[ \mathbf{y}_n^o - \mathbf{H} \left( \mathbf{x}_n^b \right) \right] \end{aligned}$$

• Bias model II:

$$\begin{aligned} \mathbf{x}_n^b &= \mathbf{f}(\mathbf{x}_{n-1}^a) \\ \mathbf{x}_n^a &= \mathbf{x}_n^b + \mathbf{K}_n \left[ \mathbf{y}_n^o - \mathbf{H} \left( \mathbf{x}_n^b + \mathbf{b}_n \right) \right] \end{aligned}$$

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• Orography correction:

 $\Delta z =$ model orography – true orography

at the observing location

- Hydrostatic correction:  $p_s = p_s^{\text{model}} \times \exp(g\Delta z/R\bar{T})$
- Bias model II:  $\mathbf{H}(\mathbf{x}_n^b) = \widehat{\mathbf{H}}(\mathbf{x}_n^b) \mathbf{c}_n$
- Bias model II plus hydrostatic correction:

$$\mathbf{H}(\mathbf{x}_n^b) = \widehat{\mathbf{H}}(\mathbf{x}_n^b) \cdot \exp(g \, \Delta z / R \bar{T}) - \mathbf{c}_r$$

- "Truth:" Generate surface pressure data using a free run of the T62L28 GFS (144 × 73 × 28 grid) for Jan.–Feb. 2004 at model grid points approximating a real observing network
- Model: Assimilate synthetic obs (surface pressure, temperature, wind) using a 60-member emsemble with the T30L7 Simplified Parametrization Primitive Equation (SPEEDY) model
- Compute RMS averages of 6-hour surface pressure forecast errors using the different bias correction schemes

# Results (verified against background)



- Bias estimation in a radiative transfer model in assimilation of satellite radiance data (with José Aravéquia, CPTEC)
- Apply Baek's scheme to surface pressure measurements in the GFS
- Quantify effects on forecast and analysis accuracy using high-resolution NCEP analyses

- Nonlinearities in model evolution cause the ensemble eventually to underestimate the true uncertainty
- Compensate by variance inflation
- In ensemble coordinates, inflation gives (for  $\rho \ge 0$ )

 $\frac{(k-1)\mathbf{w}^{\mathrm{T}}\mathbf{w}}{1+\rho}$ 

• Harlim and Hunt (2007) suggested the modification

 $\frac{(k-1)\mathbf{w}^{\mathrm{T}}\mathbf{w}}{1+\alpha\sqrt{\mathbf{w}^{\mathrm{T}}\mathbf{w}}}$ 

# Digital filtering with the LETKF (Kostelich & Szunyogh)

- Sometimes the analyzed fields excite spurious dynamics in the model forecasts due to balance issues between atmospheric variables
- Add an extra constraint to penalize (say) divergence or frequency of changes in the analysis window

$$\begin{aligned} \widetilde{J}(\mathbf{w}) &= (k-1)^{-1} \mathbf{w}^{\mathrm{T}} \mathbf{w} + \\ & [\mathbf{y}_0 - \mathbf{H}(\bar{\mathbf{x}}_b + \mathbf{X}_b \mathbf{w})]^{\mathrm{T}} \mathbf{R}^{-1} [\mathbf{y}_0 - \mathbf{H}(\bar{\mathbf{x}}_b + \mathbf{X}_b \mathbf{w})] \\ & + \mathbf{w}^{\mathrm{T}} (\mathbf{X}_b - \mathbf{F})^{\mathrm{T}} \mathbf{Q} (\mathbf{X}_b - \mathbf{F}) \mathbf{w} \end{aligned}$$

where  $\mathbf{F}$  is a suitable linear combination of background forecast perturbations.

# Conclusions

- The LETKF is a fast and accurate data assimilation scheme for meteorological applications
- Lends itself to potential applications to nonlinear models and digital filtering (though issues like multiple minima will arise)
- Local low dimensionality of the dynamics is key
- Requires a reasonably dense observing network for continuity
- The LETKF and ensemble diagnostics for predictability may be useful for climate simulations on decadal scales

- National Science Foundation
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- Preprints: www.weatherchaos.umd.edu